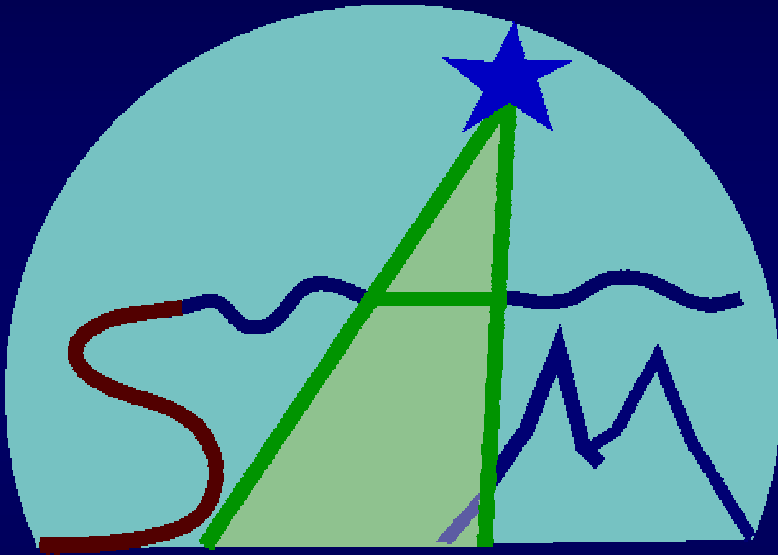


SOAR

Adaptive Module



Seeing-limited

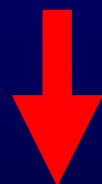
(>90% of ground-based astronomy!)



Ground
Layer
Adaptive
Optics

=

Better
seeing
in wider
field



Diffraction limit
(full AO or MCAO)

- **GLAO works in the visible**
- **Complete sky coverage**

**Matches
SOAR
strategy!**

GLAO systems

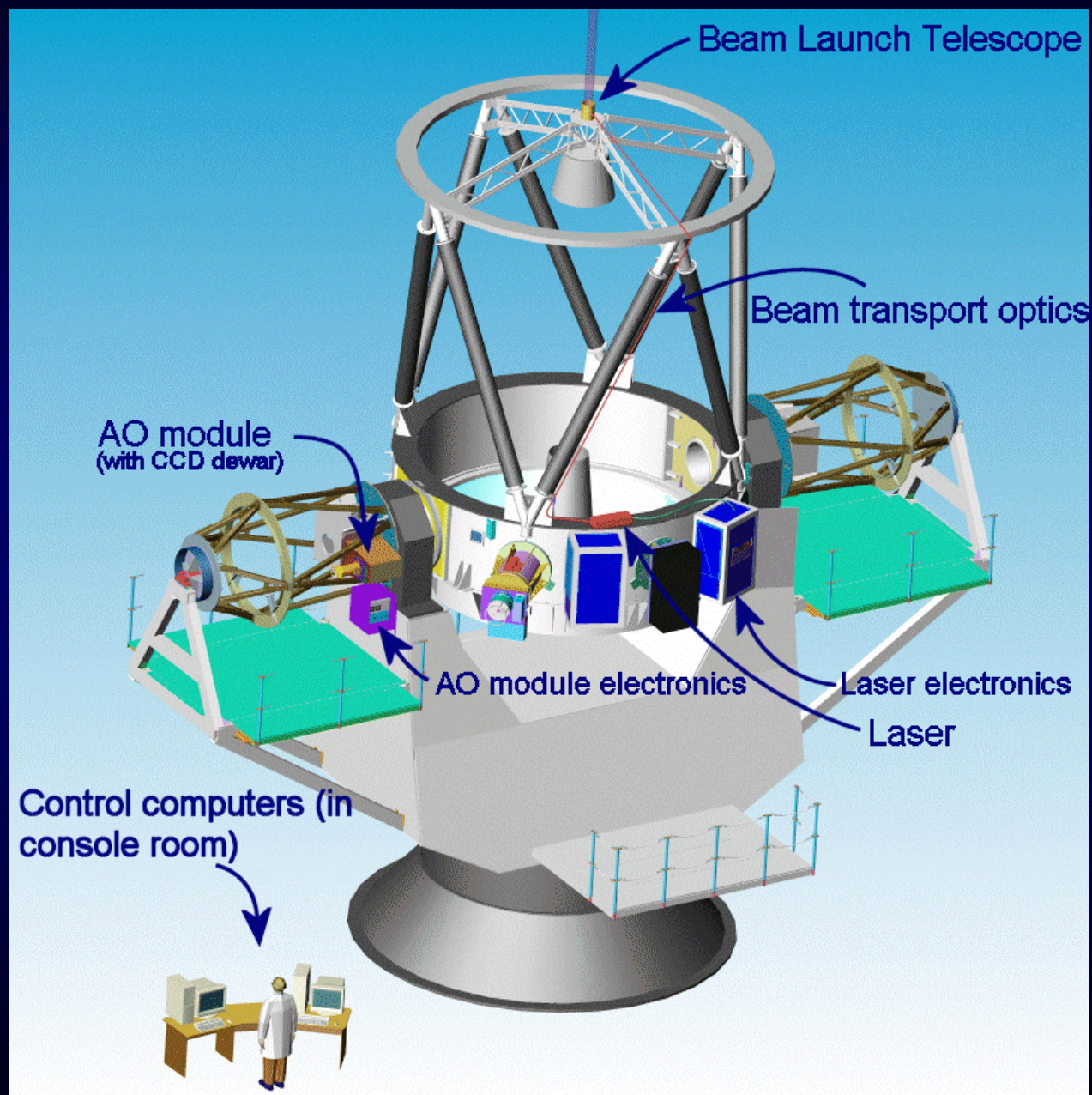
Different ways to sense and correct the ground layer...

- Gemini: 5 sodium lasers, Deformable Mirrors
- ESO/MUSE: 4 sodium lasers, adaptive secondary
- Magellan: several wide NGS
- GLAS: Rayleigh LGS at ~20km, green
- SAM: Rayleigh LGS at ~10km, UV

Rayleigh LGS senses preferentially low, strong layers

→ wide-field correction

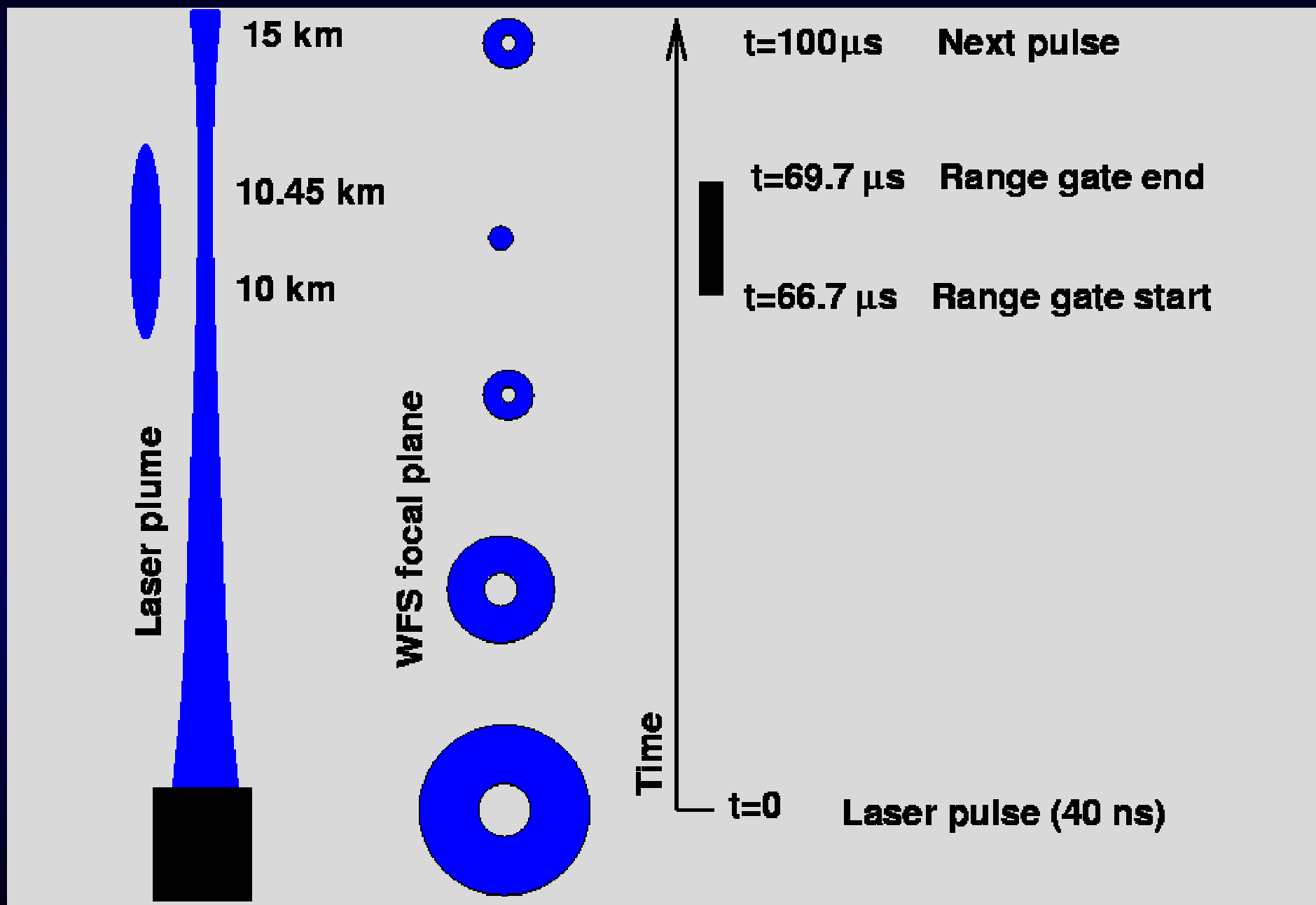
SAM at a glance



**1:1 re-imaging
3'x3' field**

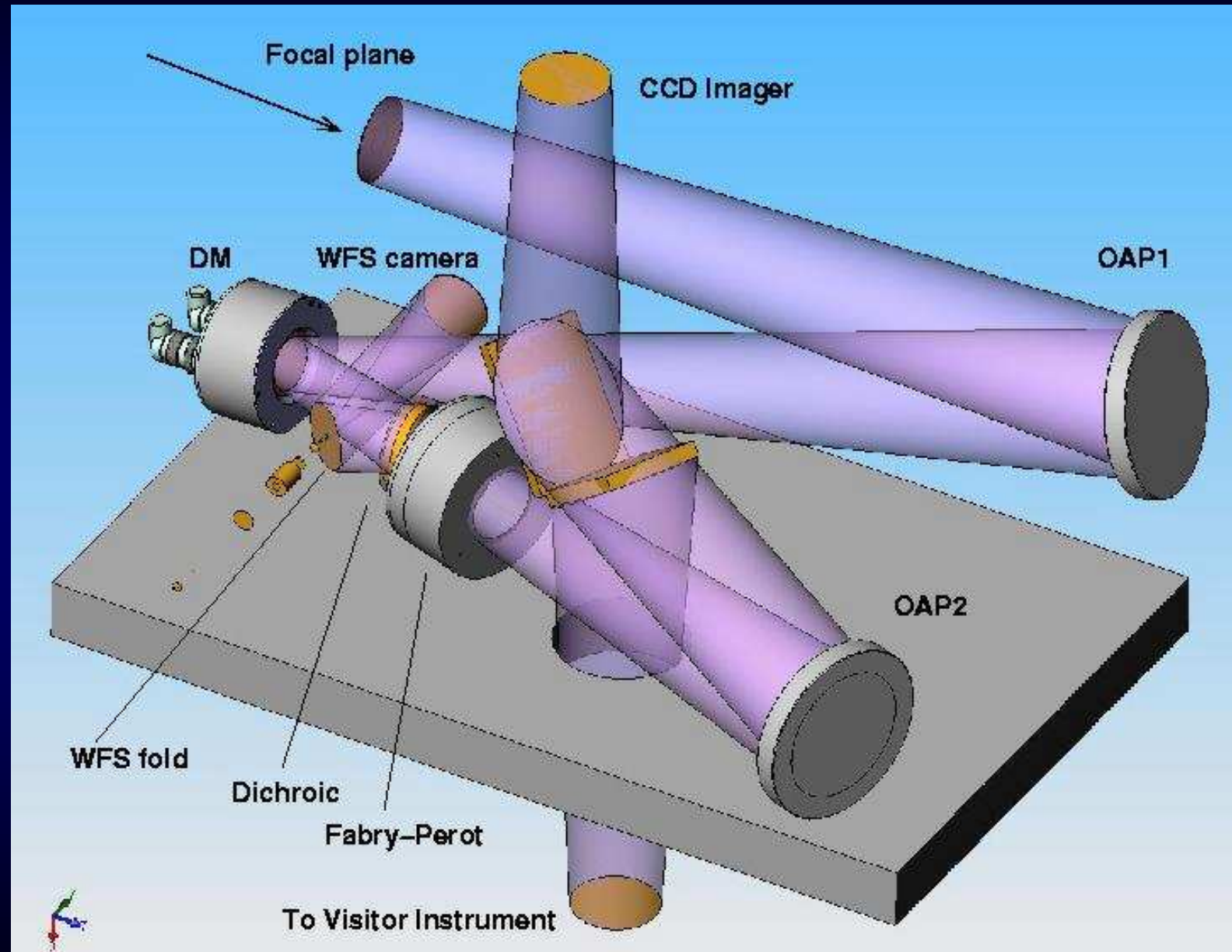
**feeds a “Visitor
Instrument” &
internal imager**

Rayleigh LGS timing

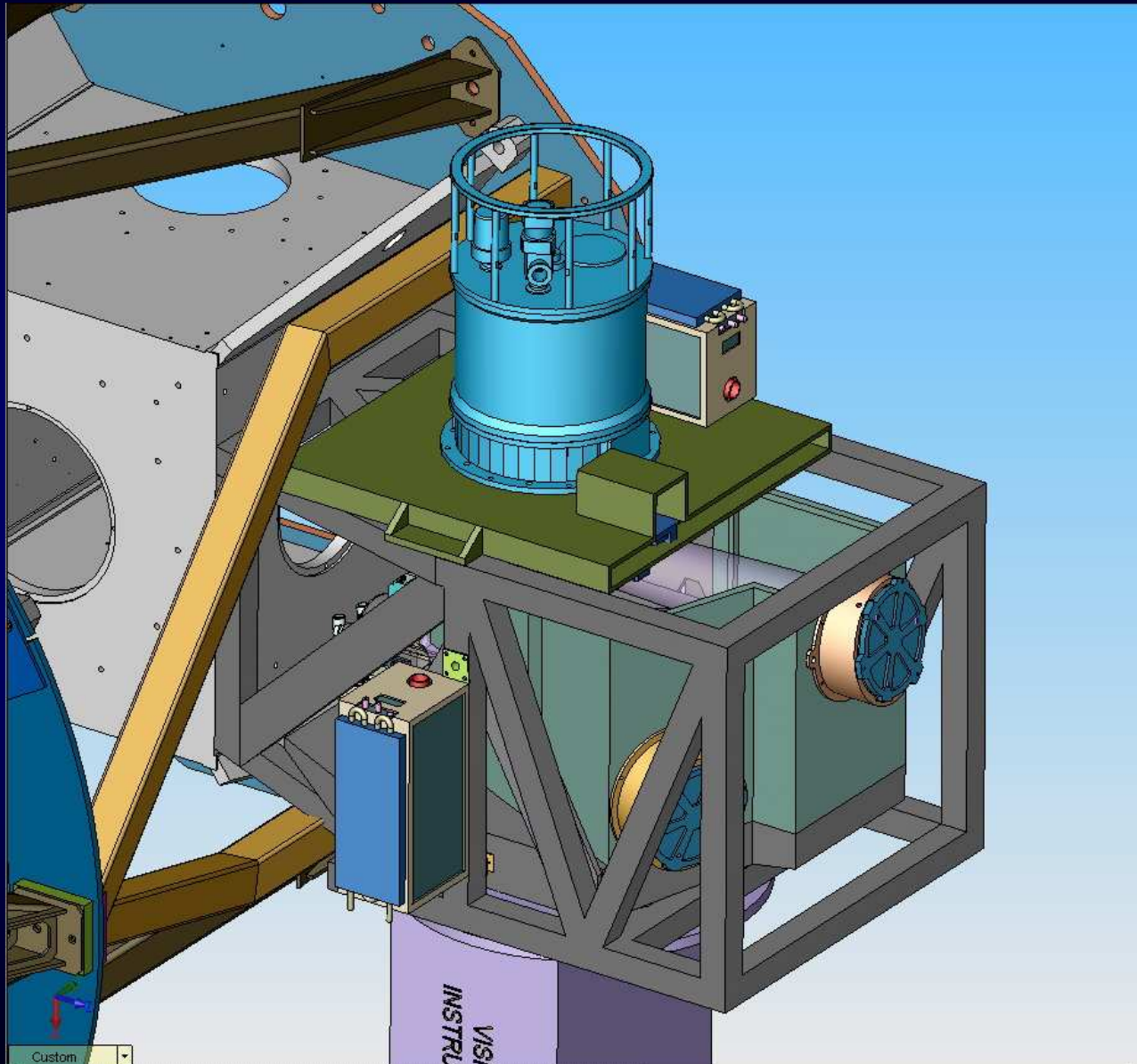


Optics

Throughput
0.85-0.9
at
wavelengths
0.4 to 0.9 μm

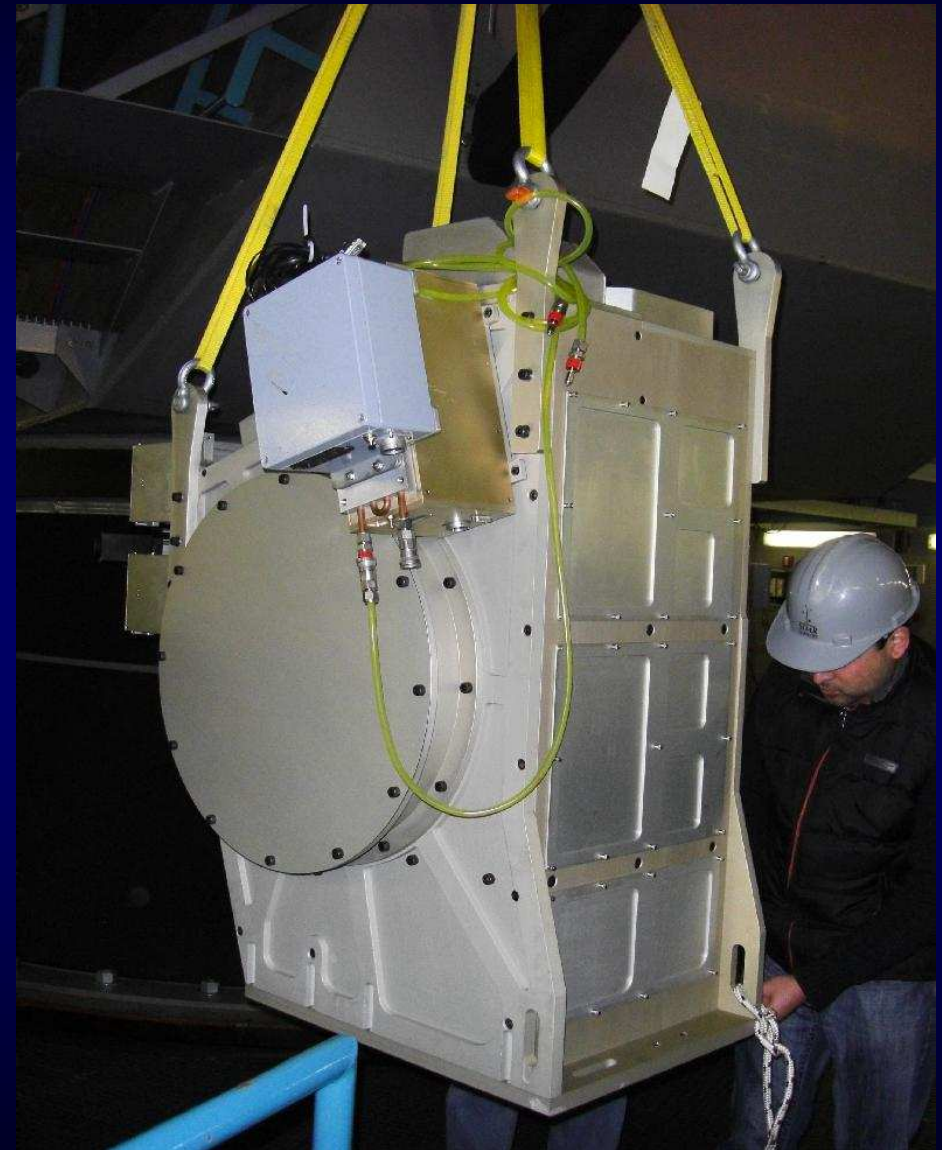


Mechanics



- Bench
- Frame
- Displaced axis
- Modified truss

Reality



Laser



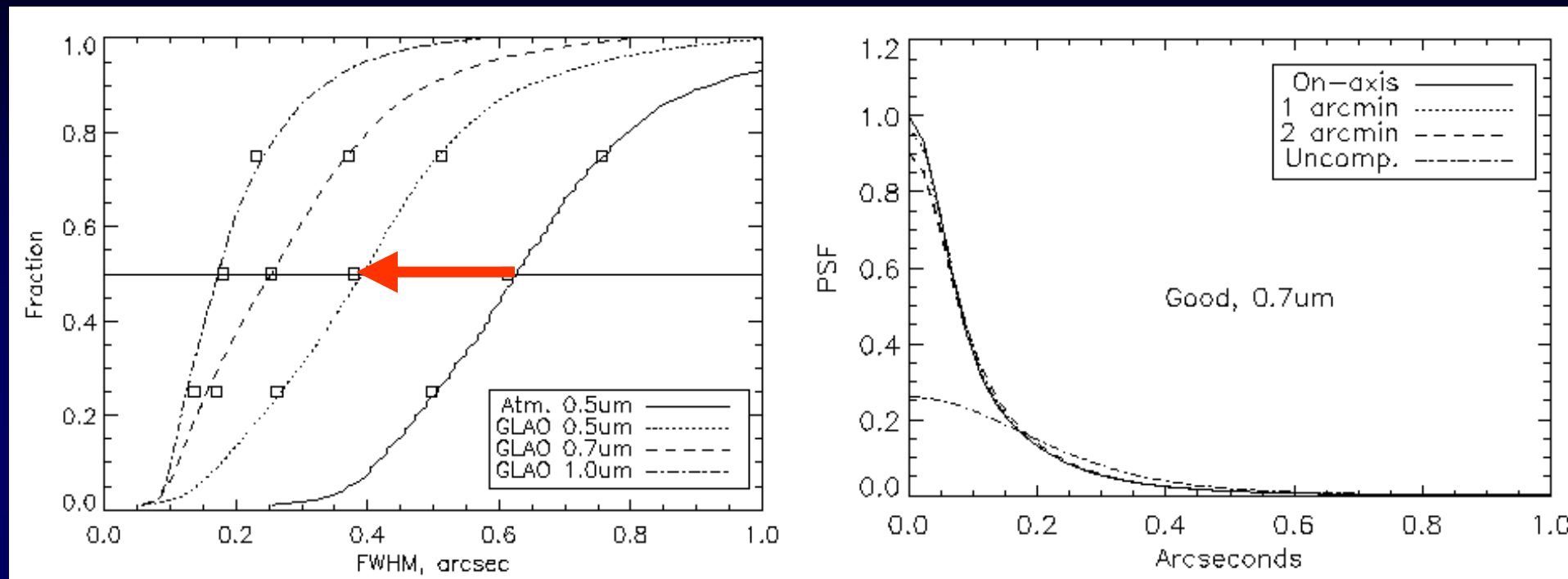
- **355nm, 10W tripled Nd:YAG Laser**
Q301-HD is used in the microprocessor industry 24/7.
 - **Standard industrial LASER**
 - **Similar (532nm) LASER used for MMT AO system**

Instruments with SAM



- **SAM-I – a Simple Optical imager matched to AO corrected images**
 - **3 x 3 arcmin FOV @ 0.07 arcsec/pixel**
 - **E2V 4k x 4k CCD**
 - **Filter wheel for 7 filters**
- **Visitor Instruments – SAM reproduces the SOAR focal plane at 1:1 scale, and the provides a copy of the ISB bolt circle allowing any (relatively small and light weight) instrument to be mounted**
 - **HR cam – commercial EM CCD camera for fast time series & speckle**
 - **SIFS**
 - **BTFI**
 - **Spartan**
 - **.....**
- **Accessible space has been reserved in the collimated beam to allow the installation of e.g. a Fabry-Perot Etalon**

SAM Performance Predictions



- New data for Cerro Pachon (MASS+DIMM)
- Models representative of good, typical, bad conditions
- LGS @ 10km, two tip-tilt NGS at 2.5'

Detailed predictions

Wavelength 0.7 μm , SOAR \rightarrow SAM

Conditions	FWHM, arcsec	EE(0.3'')
Good, $z=0^\circ$	0.45 \rightarrow 0.17	0.20 \rightarrow 0.39
Good, $z=50^\circ$	0.60 \rightarrow 0.28	0.12 \rightarrow 0.25
Typical, $z=0^\circ$	0.55 \rightarrow 0.25	0.14 \rightarrow 0.28
Typical, $z=50^\circ$	0.74 \rightarrow 0.43	0.09 \rightarrow 0.16
Bad, $z=0^\circ$	0.68 \rightarrow 0.37	0.10 \rightarrow 0.19
Bad, $z=50^\circ$	0.91 \rightarrow 0.60	0.06 \rightarrow 0.10

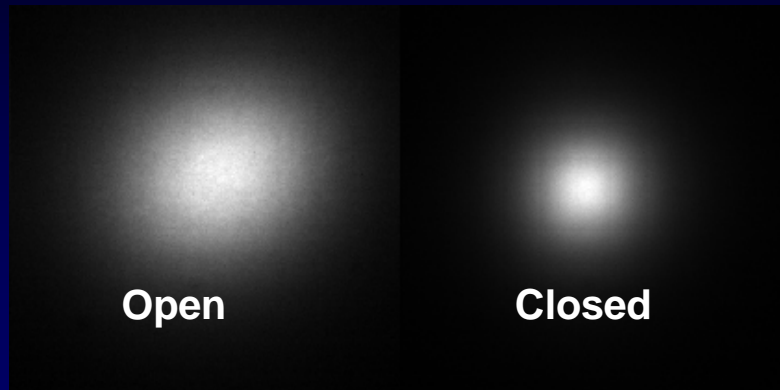
FWHM gain $\sim 2x$; EE(0.3'') gain $\sim 2x$

SAM First Light



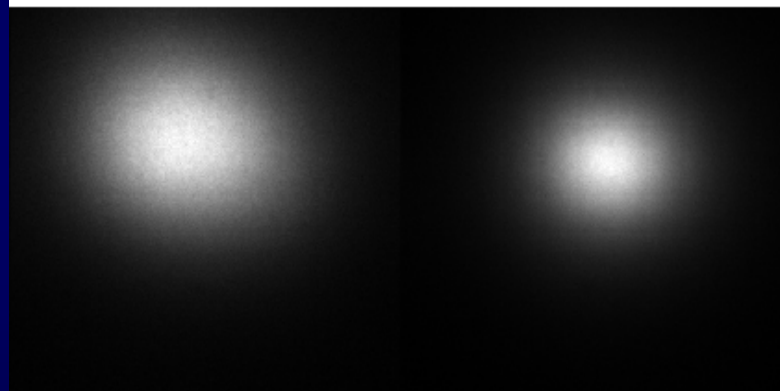
- SAM was installed on SOAR on 3 Aug 2009
 - Goal was to test closure of AO loop in NGS mode
- The weather did not cooperate
 - Aug 6 – 7: two cloudy nights
 - Aug 31 – Sep 3: 2nd run, some data with mostly bad seeing
- Nonetheless served to whet everyone's appetites

NoAO → YesAO



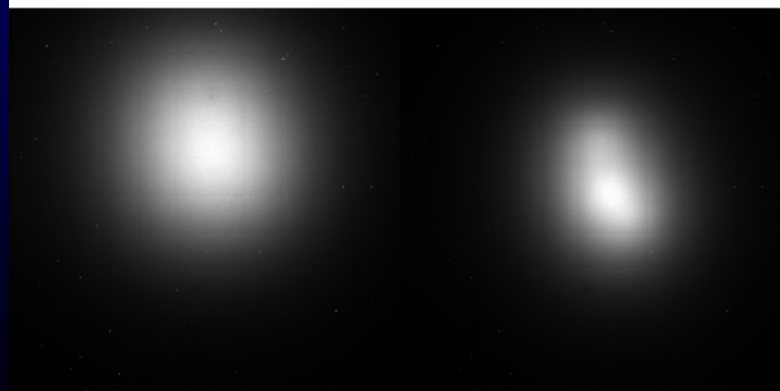
HIP 86846, 550nm

(1.29", 1.09") → (0.73", 0.75")



FIN 307, 550nm

(1.28", 1.08") → (0.90", 0.83")



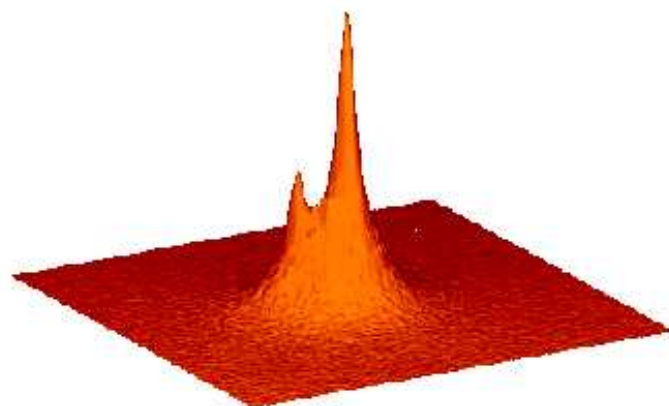
I 385, R-band

1.12" → 0.75"

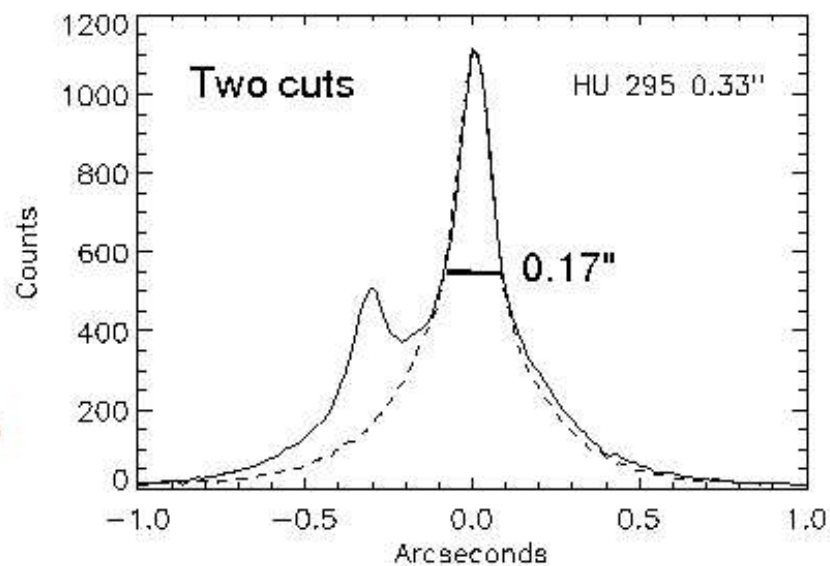
**Stable, centered
PSF**

**Focus and
astigmatism
are corrected!**

Best Images



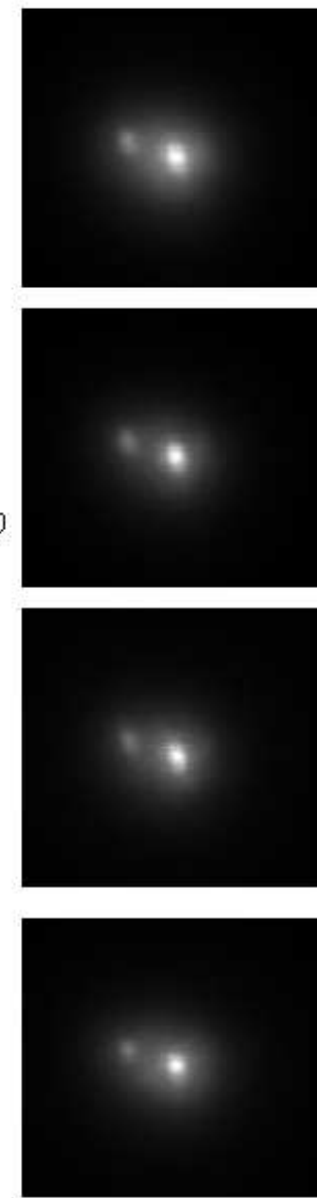
HU 295 0.33"
I-band, 200x0.02s = 2s
SAM, September 3, 2009



Four 2-s images



I 22 0.41"
I-band, 200x0.1s = 20s
SAM, August 31 2009



Next Steps



- SAM is now back in the lab
 - Integration & test of Tip-Tilt Sensors & ADC
 - Fabrication of SAM-I Dewar nearing completion
 - E2V 4kx4k CCD in hand
- LASER under test in Lab
- Waiting for delayed delivery of optics for LASER launch telescope
 - All mechanical parts designed and being fabricated
- Anticipate going back to the telescope in July/August
 - Full shake down of SAM in NGS mode
 - First Laser Launch
- Commissioning during 2010B/2011A