

The Industrial Revolution of Galaxies Cluster Finding in the Dark Energy Survey

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Science with the LSST: A Brazilian/US Joint Workshop

APRIL 1-4, 2012















Outline



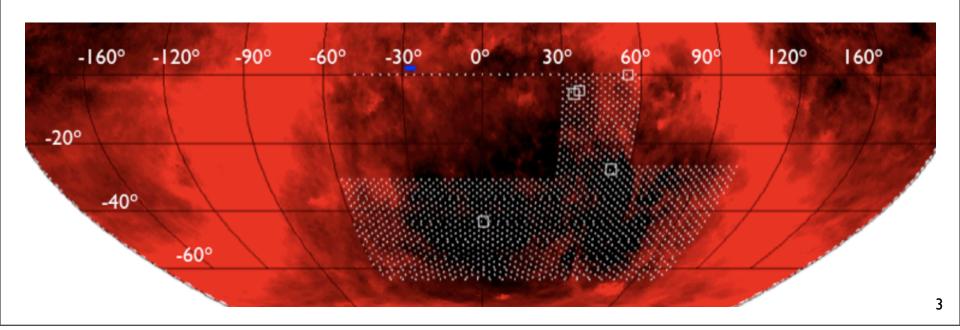
- Introduction
- Cluster Finders Industrialization
- Data handling
- Parallelization
- Results
- Scientific collaboration network
- Fine tuning and Cluster finding challenge
- Lessons learned



Introduction



- Dark Energy Survey 5000 square degrees
- 400 million galaxies spread on the cosmic web
- Four probes to cosmology: SN, BAO, WL and Clusters
- Several methods to find clusters of galaxies
- Understand their selection functions → better cosmology.





Linea Cluster Finders Industrialization



- Run Cluster Finders (CFs) on mock catalogs
- Match clusters to halos in simulations
 - Calculate completeness and purity
- Have several CFs in a controlled environment
 - Keep track of data provenance, code version, and configurations
- Faster processing response time on data releases



Lines Cluster Finders Industrialization

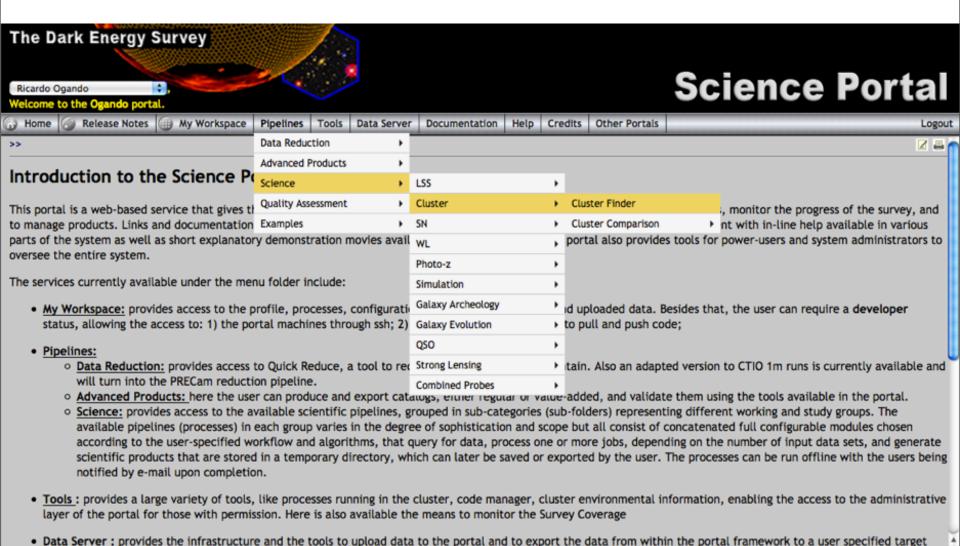


- Matched Filter (L. Olsen), WAZP (C. Benoist), VTRS (W. Barkhouse), GMBCG (J. Hao), Lambda (E. Rykoff), C4+ (L. Baruah), and zVT (M. Soares-Santos)
- Framework to compare results allowing:
 - **Fine tuning**. Finding the best set of CF input parameters, e.g. M* and threshold.
 - -Cluster Finding challenge. Pros and cons of each CF method. Assess their selection functions.
- Compliance: CFs produce common output: RA, Dec, z, mass proxy rank



The Cluster Finder pipeline



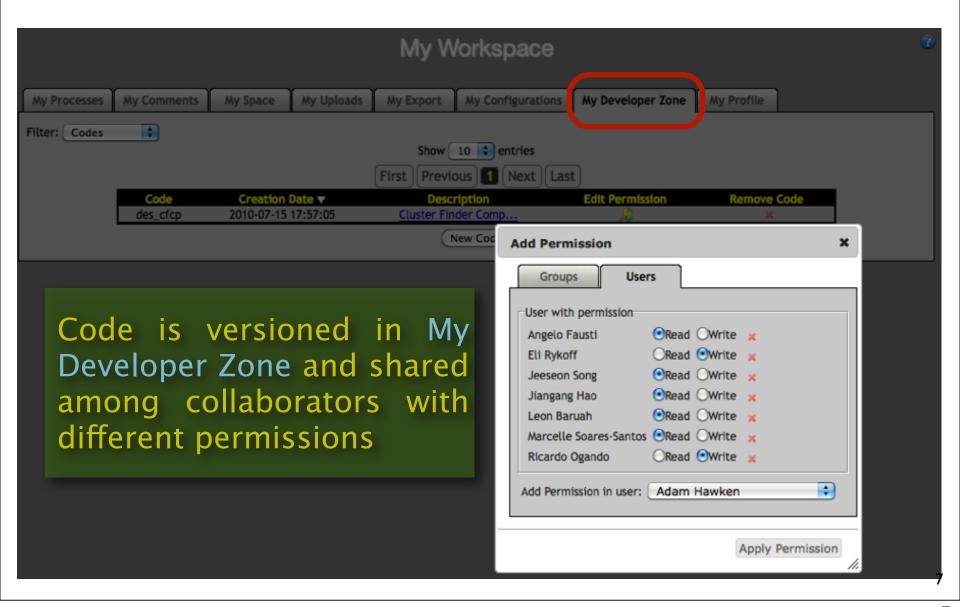


http://ogando.linea.gov.br/# 012 - 14:17:59 Copyright



Lines Code versioning and share



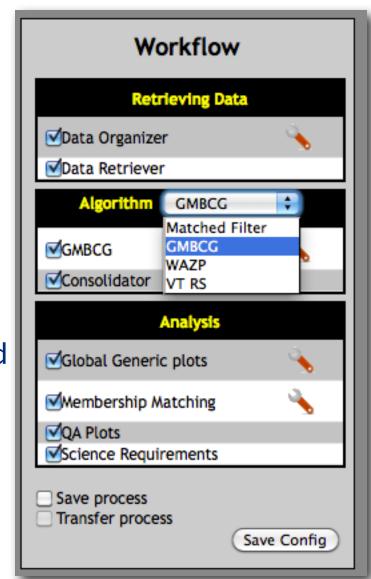


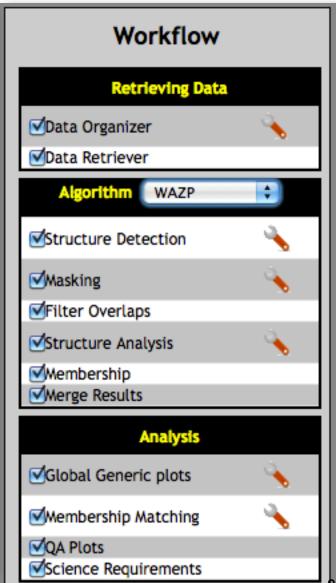


Workflow



- CF codes wrapped in python modules
- Modules concatenated in XML workflows

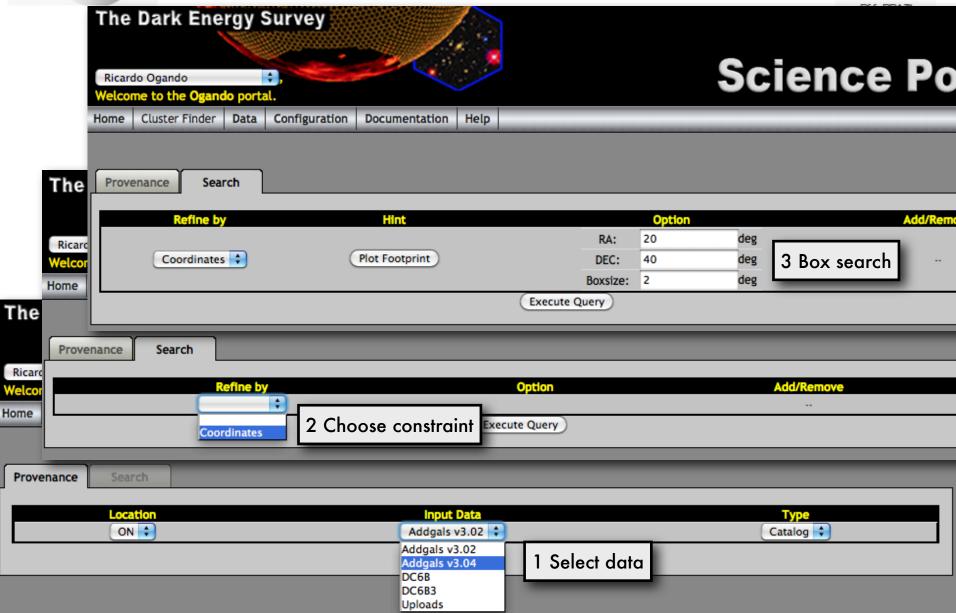






Data search







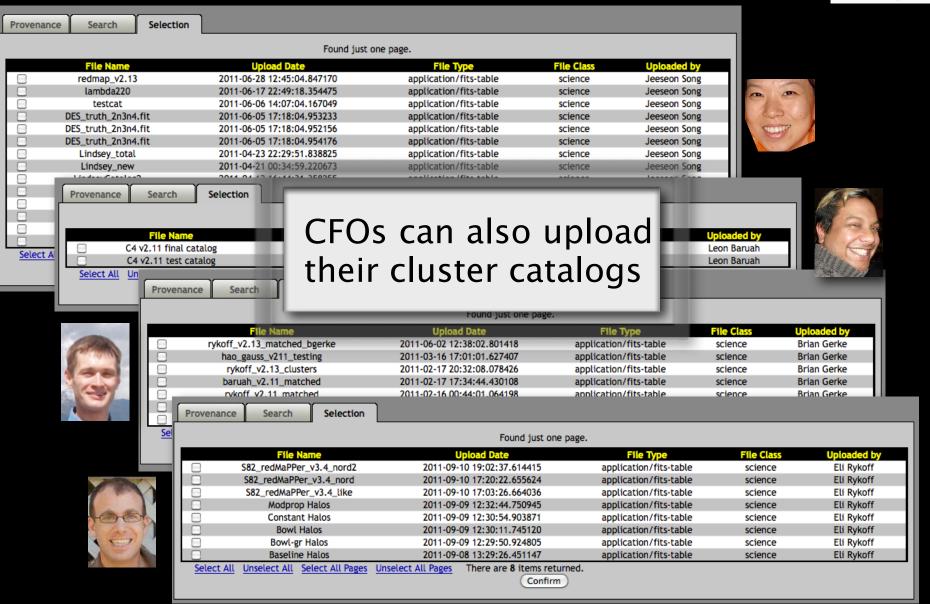
Data handling



- ADDGALS: mock catalogs (200sqdeg)
- Data Challenges: mock observations (100sqdeg)
- Blind Cosmology Challenge: unknown cosmology mock catalog (5000sqdeg)
- Other data: CFHT, S82, and BOSS
- Ingested to PostgreSQL database.
- Q3C indexing for large astronomical catalogs







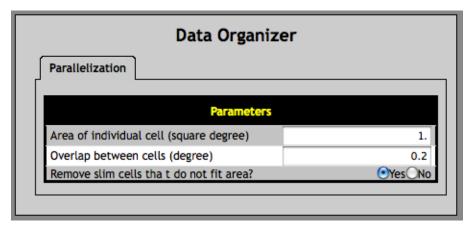


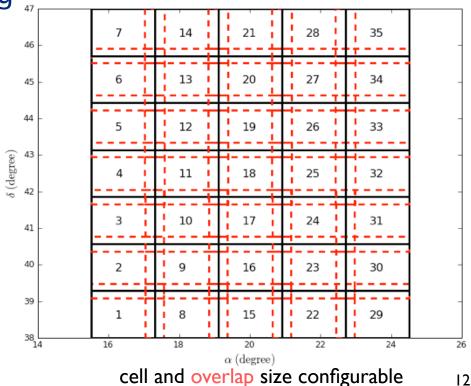
Parallelization



- Important for large area analysis
- (almost) Embarrassingly parallel
- Resolve overlaps by bisection before stitching

Some CF needs post-processing





60 sqdeg/35 cells

12

View Processes

Results Comments



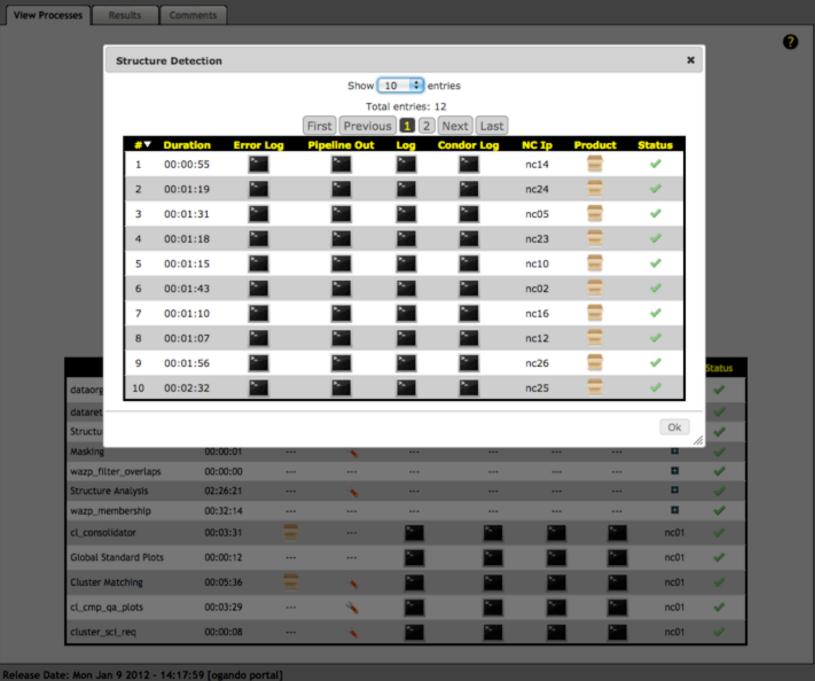
Pipeline: Cluster Finder					
User	Ricardo Ogando				
Process	10007936				
Start	2012-02-16 02:17:10				
End	2012-02-16 08:50:15				
Duration	06:33:05				
Ingest Time	00:00:02				
Expiration	2012-02-21 08:50:15				
Jobs	78				
Automatically Saved	Yes				
Algorithm	WAZP				
Configuration	View				
Code Version	View				

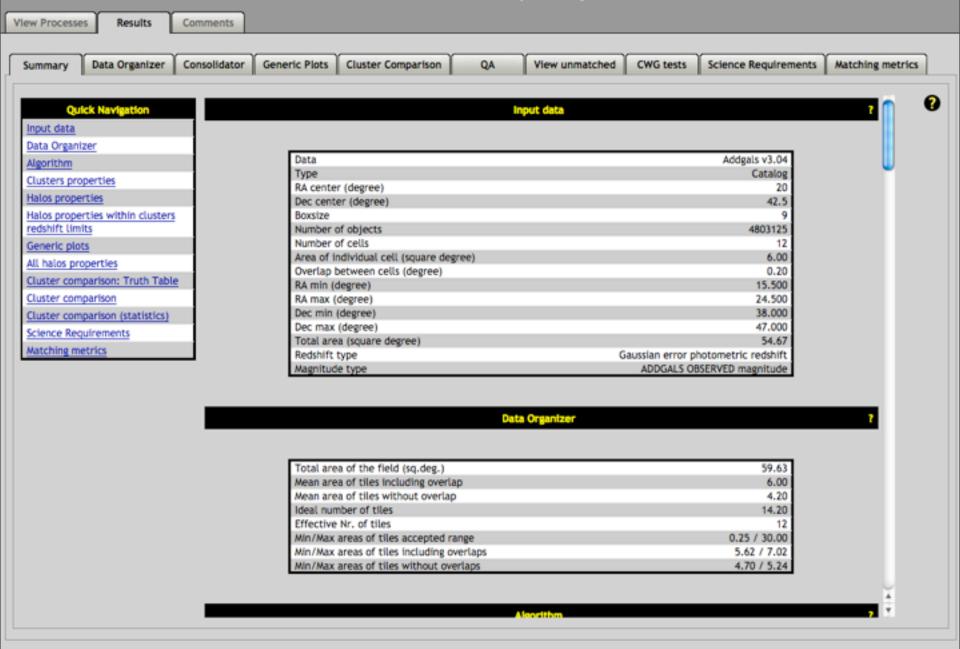
Input Data						
Input data	Addgals v3.04					
Type	Catalog					
RA (deg)	20					
DEC (deg)	42.5					
Selected Region (deg)	9					
Objects	4803125					
Area (sq deg)	54.67					
Density (objects/sq arc min)	24.4					

Module	Duration	Products	Config	Error Log	Pipeline Out	Log	Condor Log	NC	Status
dataorganizer	00:02:05		*	*	Sec. 1	~	~	nc01	4
dataretriever	00:01:35	***		***				D	4
Structure Detection	00:01:24		*	•••					4
Masking	00:00:01		4	***		•••	•••	D	4
wazp_filter_overlaps	00:00:00	***	***	•••		***		•	4
Structure Analysis	02:26:21		-						4
wazp_membership	00:32:14	***	•••	***		•••		+	~
cl_consolidator	00:03:31	=		*	200		2	nc01	4
Global Standard Plots	00:00:12			*	~	2		nc01	~
Cluster Matching	00:05:36	=	4	*	-	200	~	nc01	4
cl_cmp_qa_plots	00:03:29		*	*	S-1	*	~	nc01	~
cluster_sci_req	00:00:08		4	-	*	*	-	nc01	4

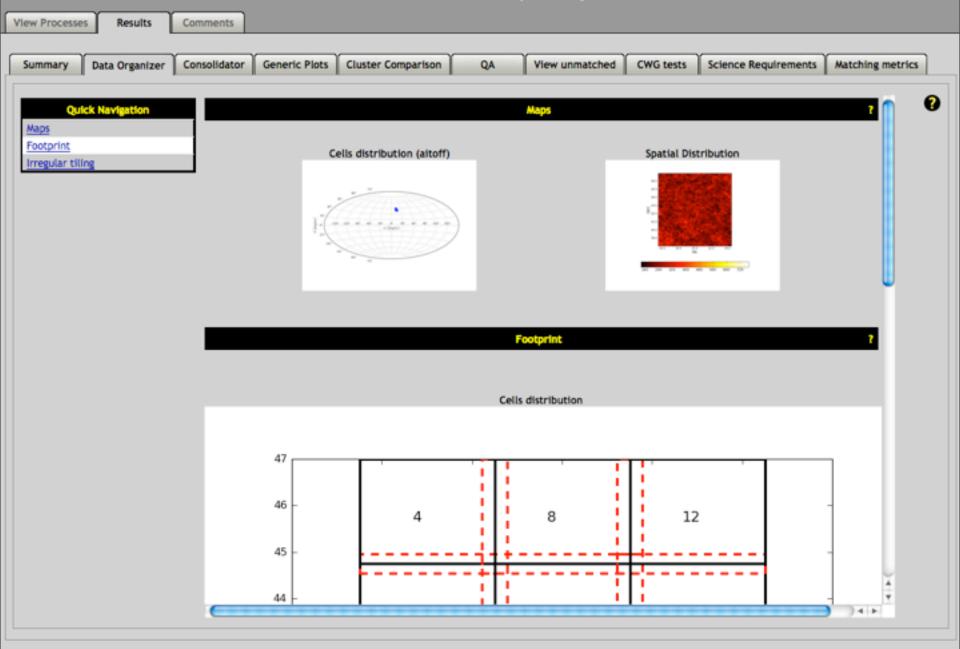
Release Date: Mon Jan 9 2012 - 14:17:59 [ogando portal]



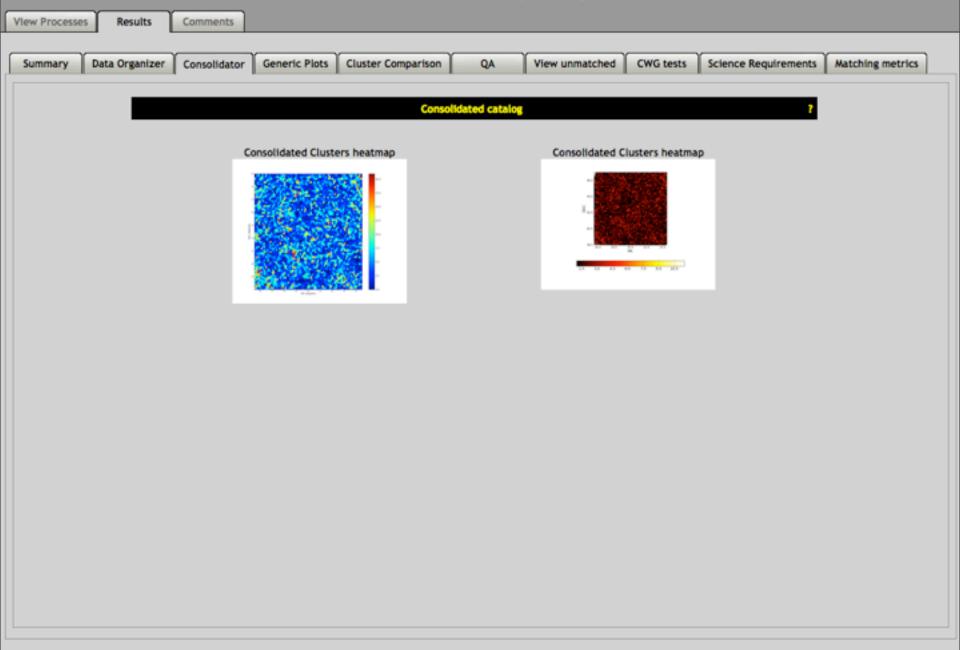


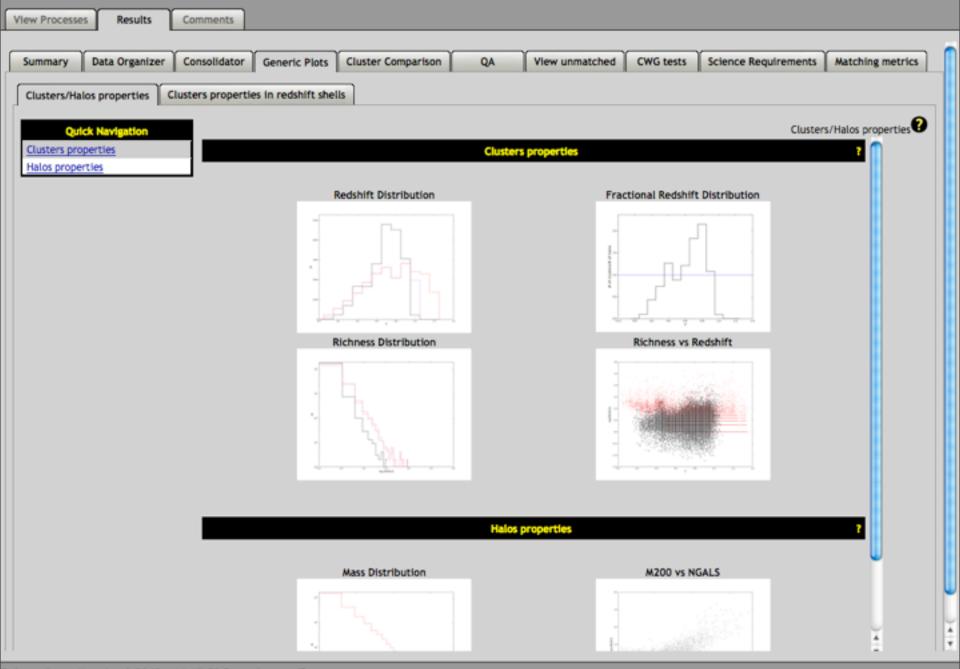


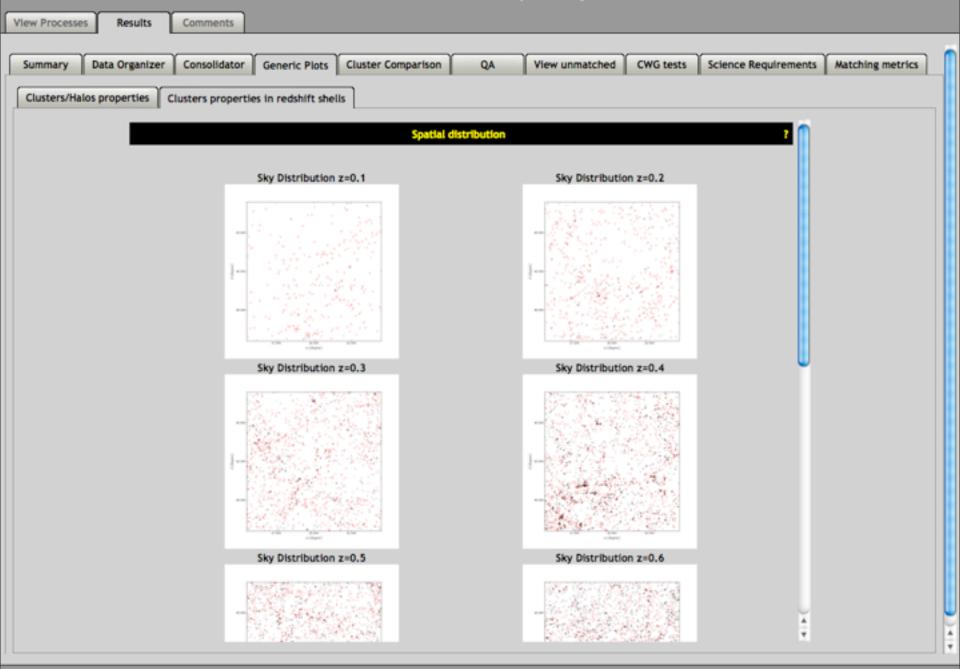
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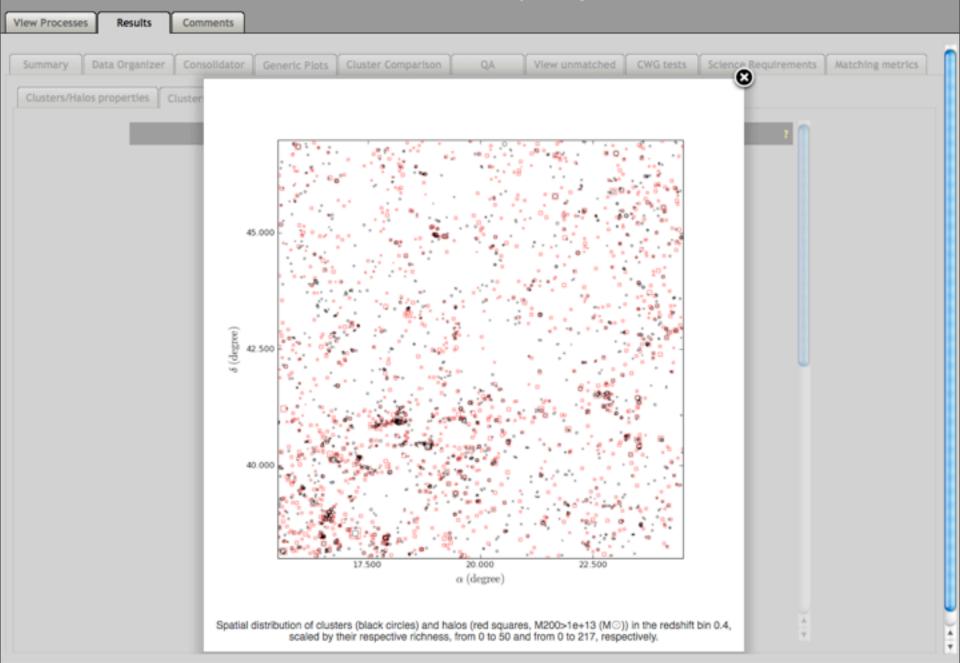
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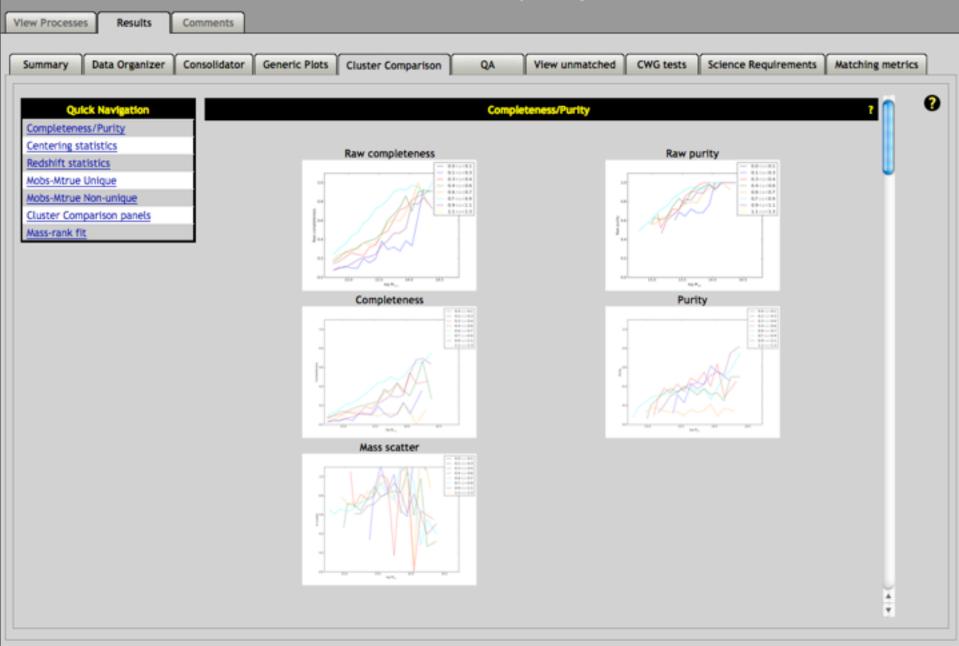


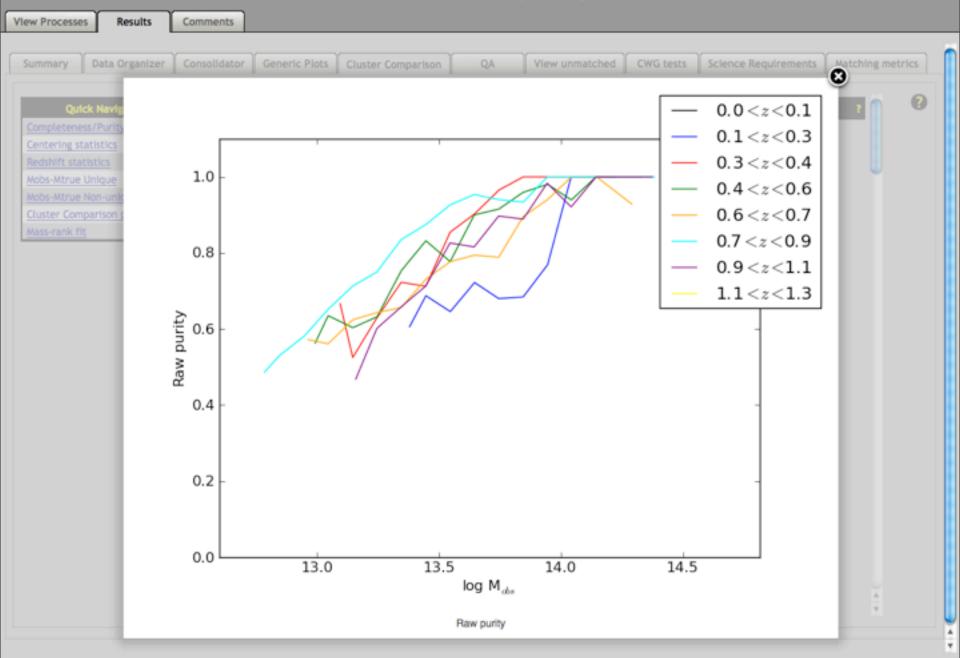


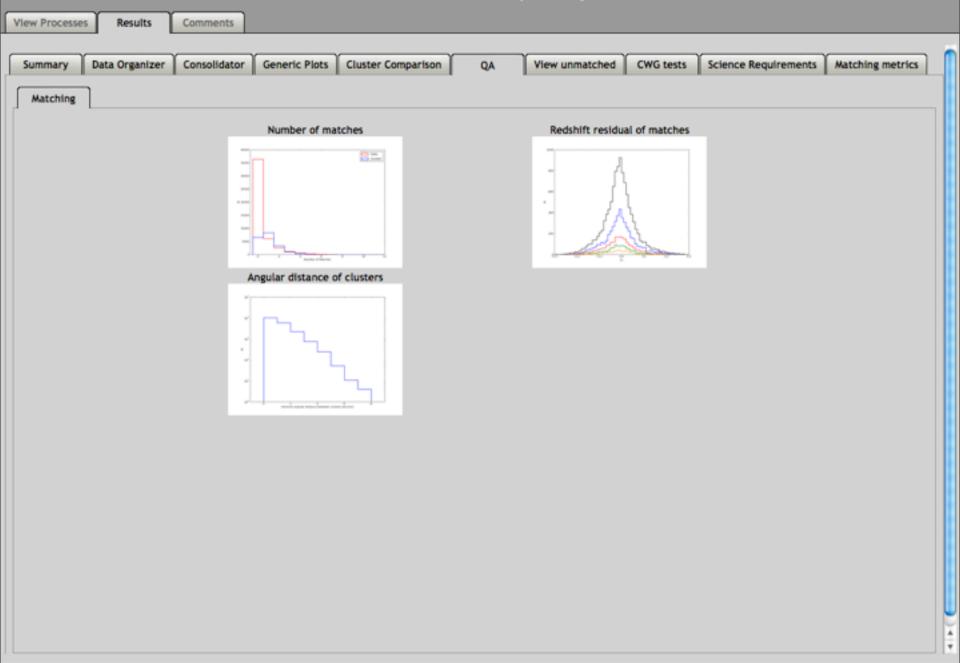


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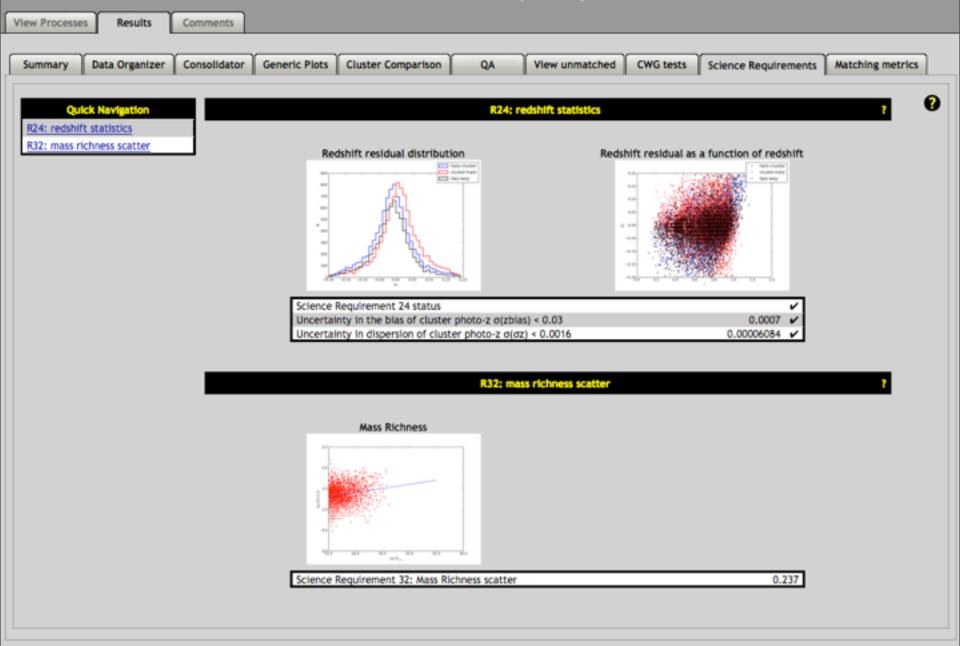








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Comments

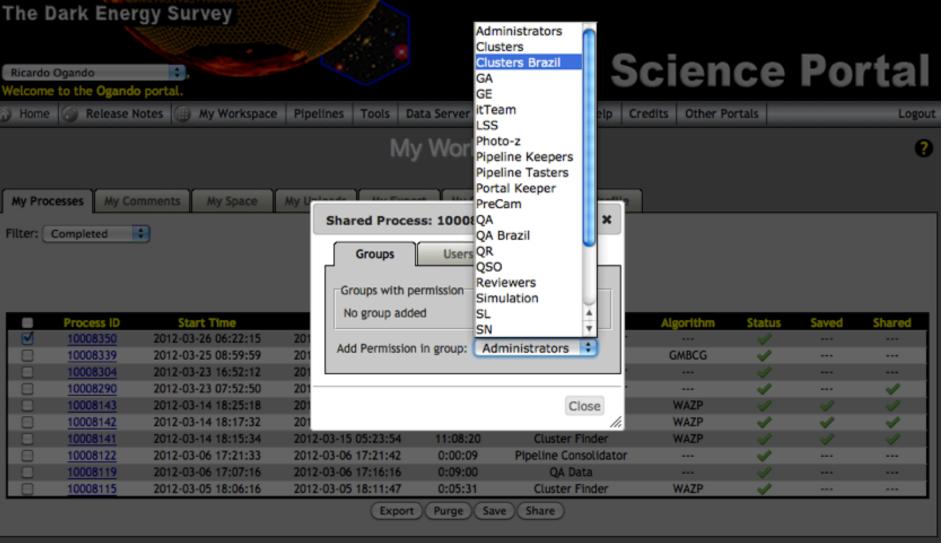


	Pipeline Consolidator	
View Processes Results Comments		
		•
User: Ricardo Ogando	ve Comment	
	e parameter. ANNZ and all photozs are calculated in SLAC and included in the catalog. I will update the lengthy titles, just wanted to share the	good stuff first. 2012-03-23 10:26:06
User: Luiz da Costa is the ANNZ run done by us ? or is this a different catalog?		2012-03-23 09:45:09
User: Luiz da Costa GMBCG has no explicit magnitude limit cut ? no threshold	? we need to make these parameters explicit I still think that the titles of the tables should be shorter and the table caption should be placed	d in the help 2012-03-23 09:41:18
	n GMBCG is using, has worst performance for this version of Addgais than other photozs, GMBCG, nevertheless, has better redshift performance. If uses BCGs to center the cluster, and the mock catalog, by construction, is centered also on BCGs, this is quite remarkable.	Another surprise is 2012-03-23 08:04:03
User: Ricardo Ogando This includes reorganization of Summary and a new tab ab can vary considerably.	cout centering and redshift offset. In order to compare the runs besides the regular histogram, I also present the normalized histogram, since the	e number of detections 2012-03-23 08:02:47
User: Ricardo Ogando Cluster Finder pipeline Consolidator update.		2012-03-23 00.02.47
		2012-03-23 07:55:31



Jobs can be **shared** with groups or individuals, which can comment the results

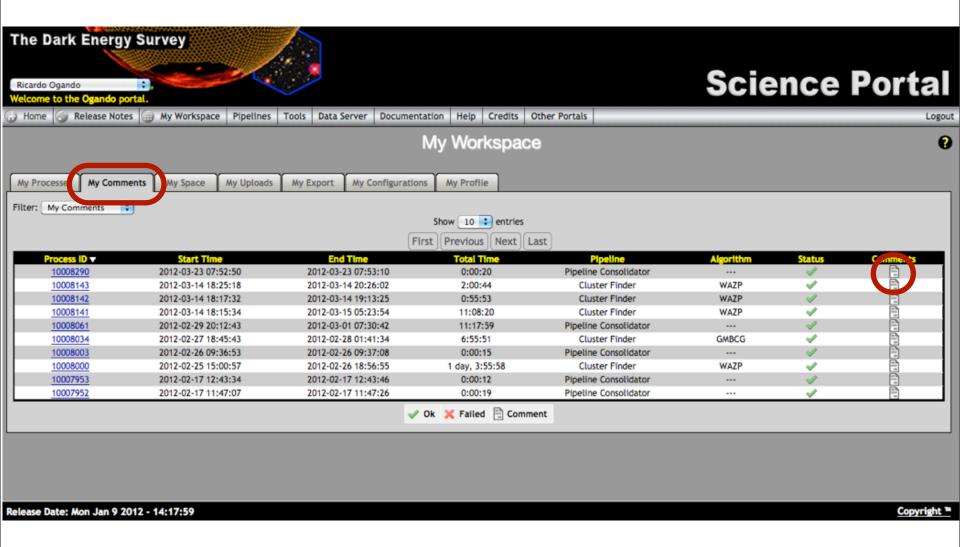






Comments





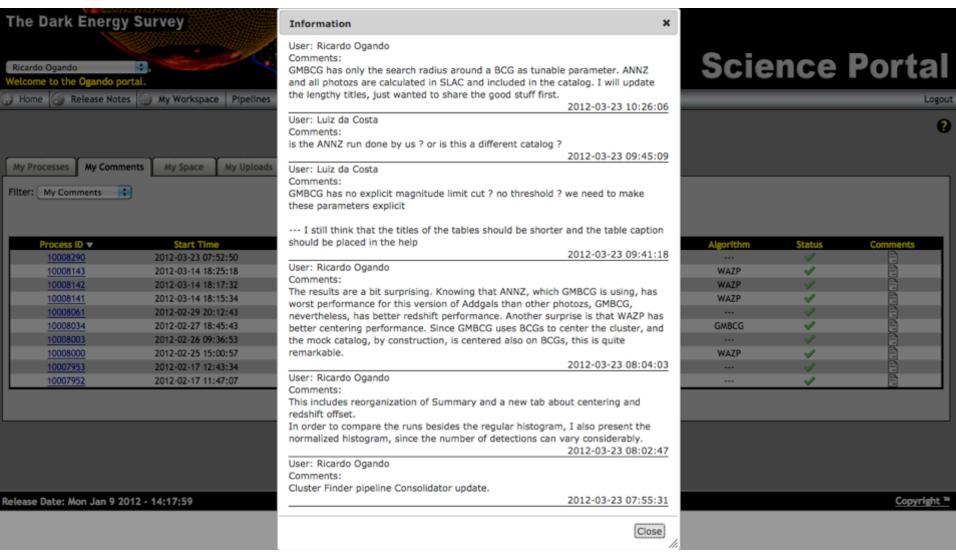
Browse comments through My Workspace

29



Comments







Lines Pipeline Consolidator



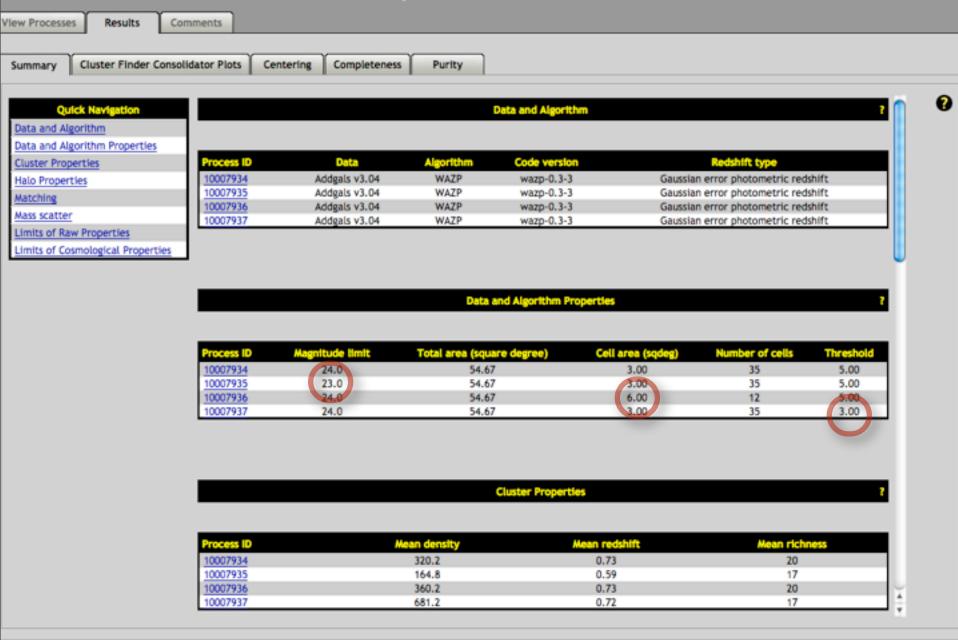
- Access saved processes
- Compare side by side
 - Published metrics
 - -Common output from halo-cluster matching code
- Cases:
 - Fine Tuning: WAZP
 - Cluster Challenge: WAZP VS GMBCG

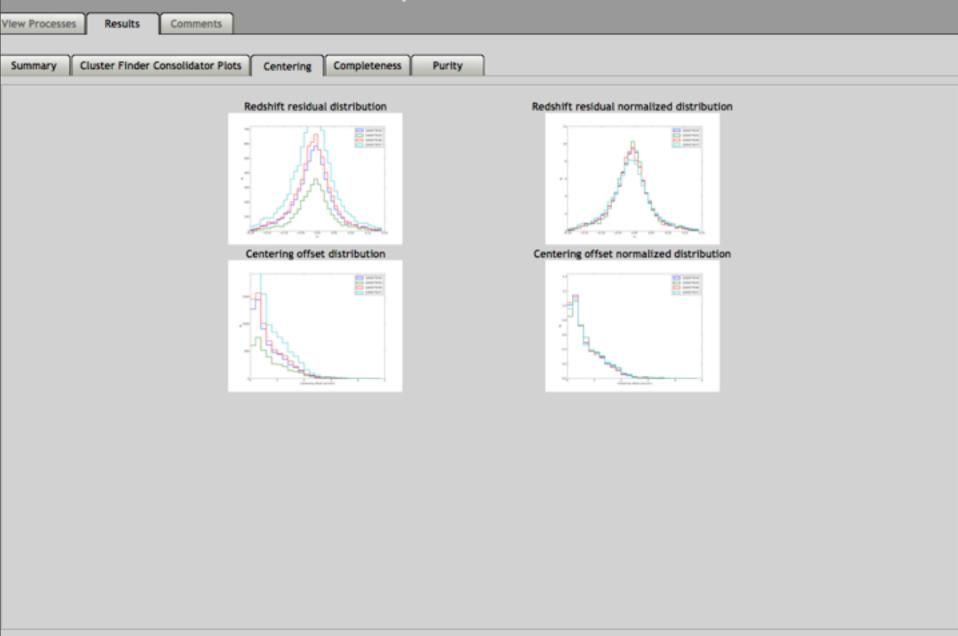


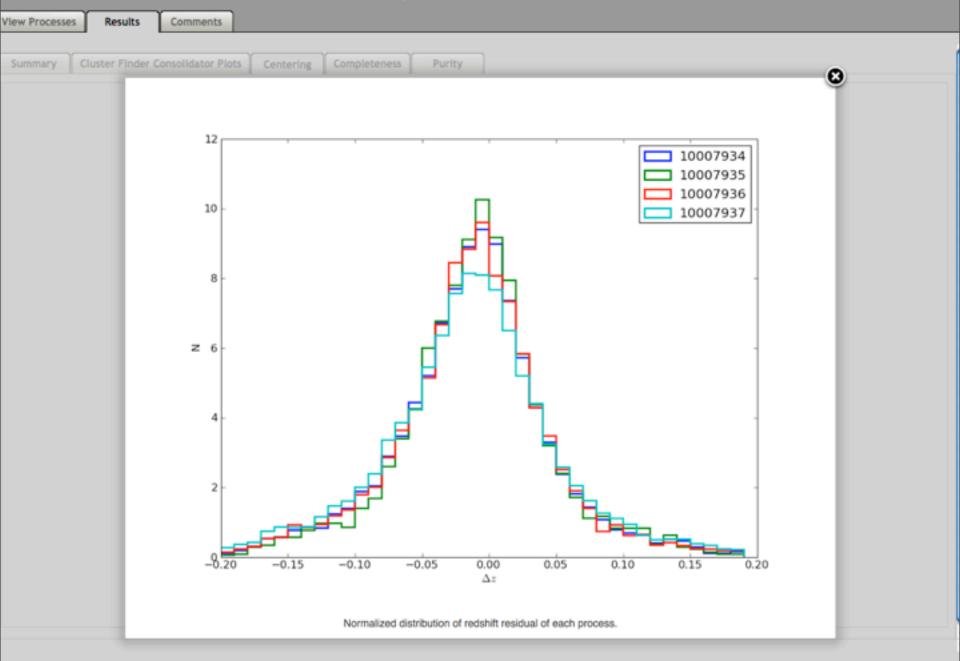
Uner WAZP Fine Tuning

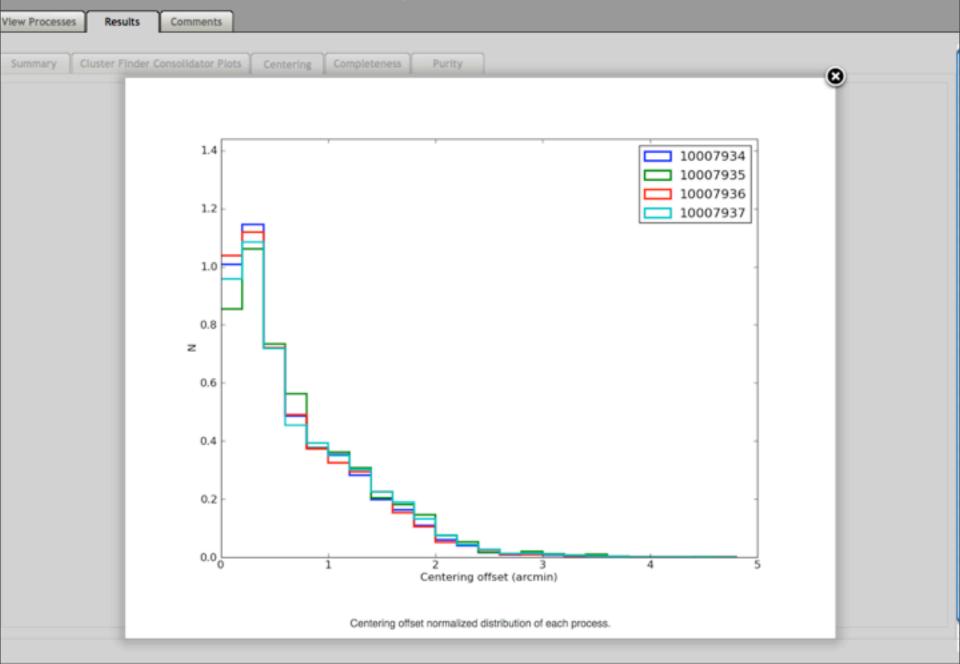


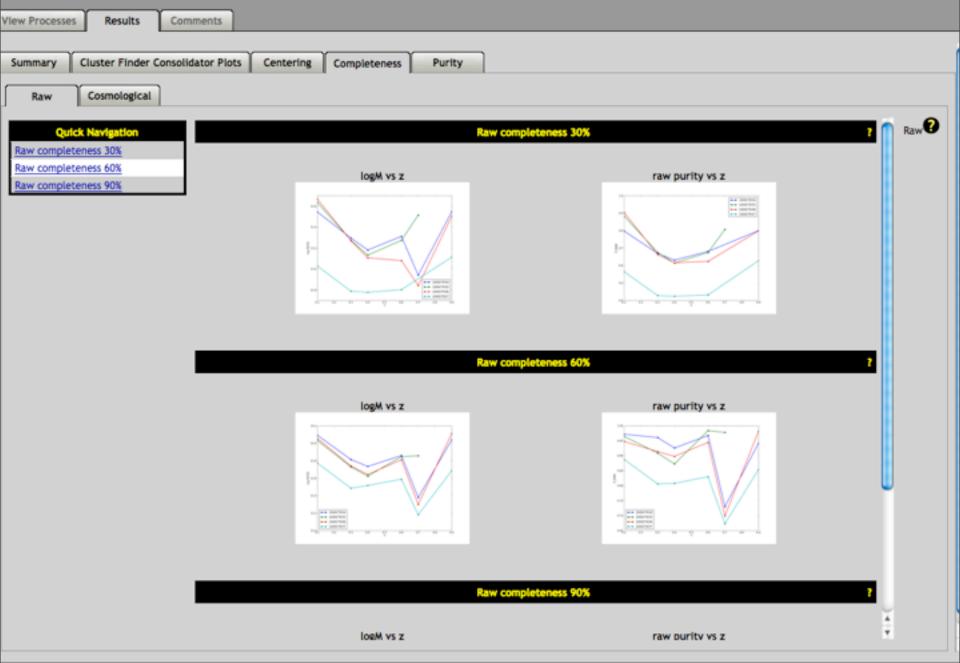
- WAZP (Christophe Benoist/OCA)
 - Identification of galaxy over-densities in Ra, Dec, zphot space.
 - —The underlying algorithm uses 2D and 1D density field reconstruction based on wavelet transforms.
- Tunable parameters:
 - —Threshold
 - Magnitude limit
 - Parallelization (size of cells in sky partition)

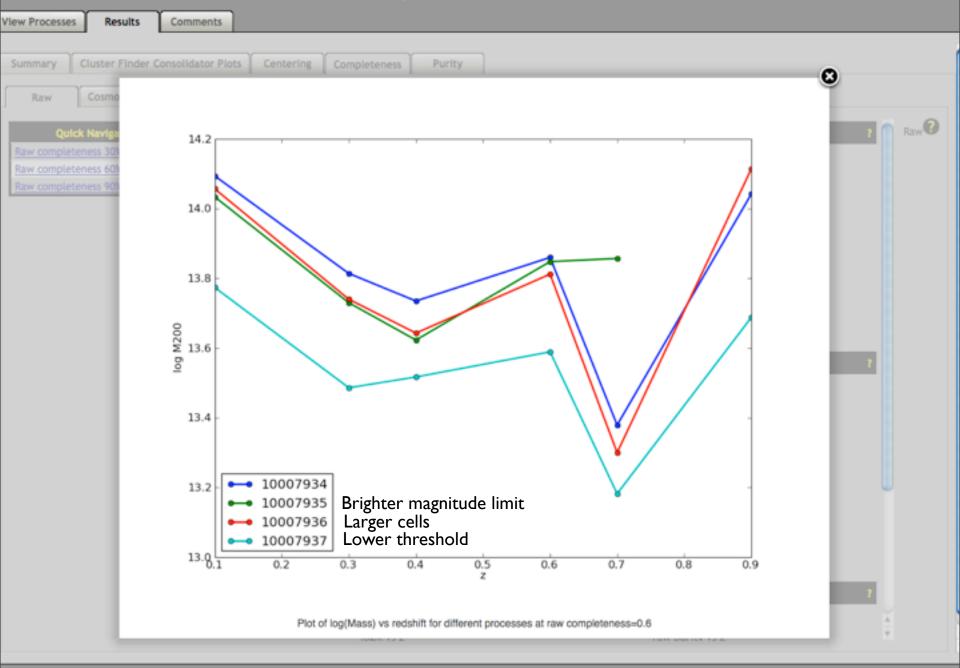










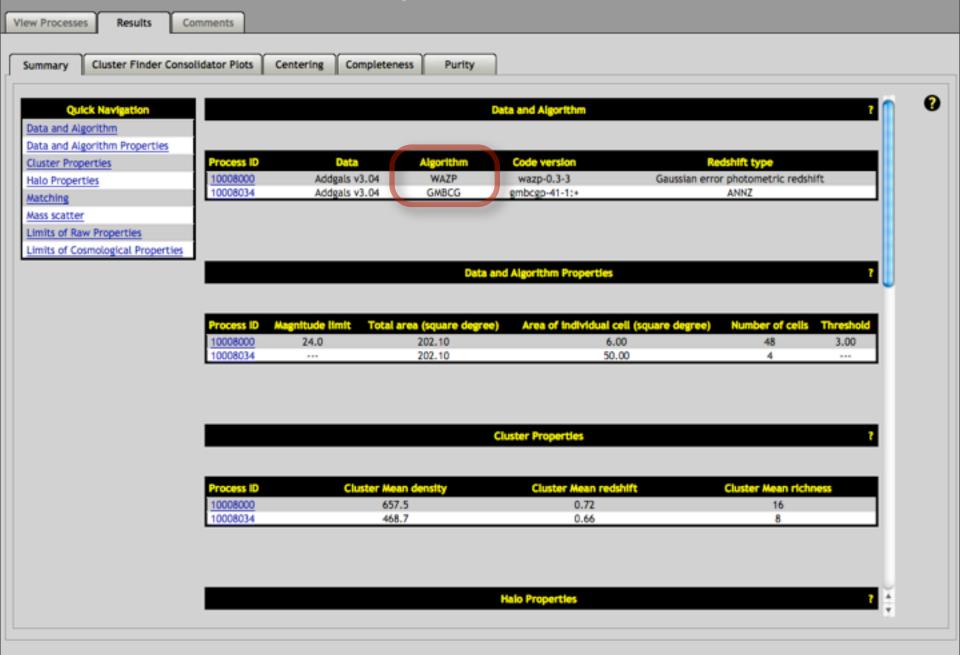




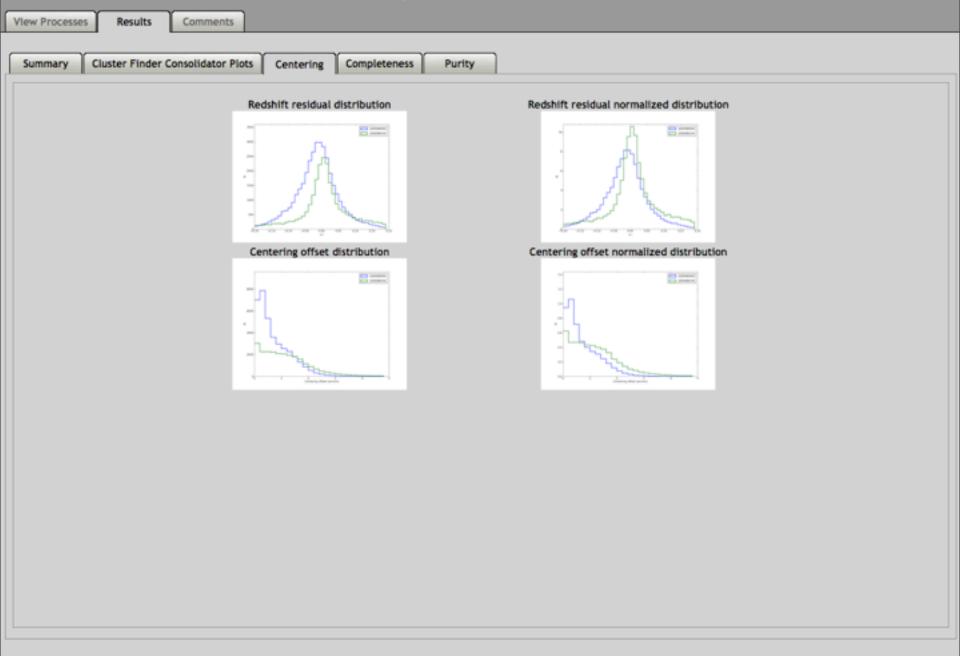
Cluster Challenge

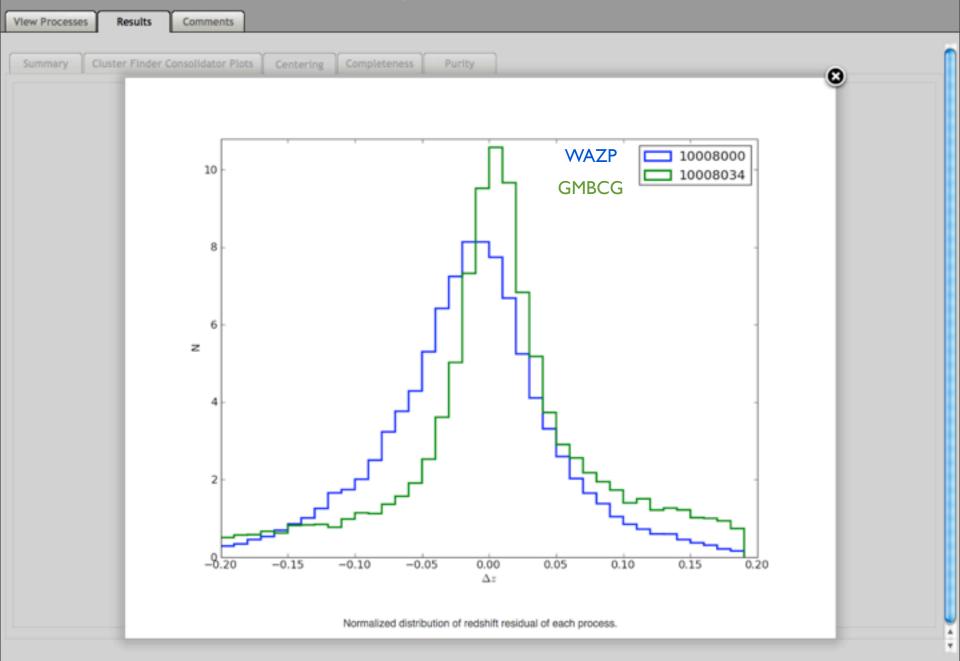


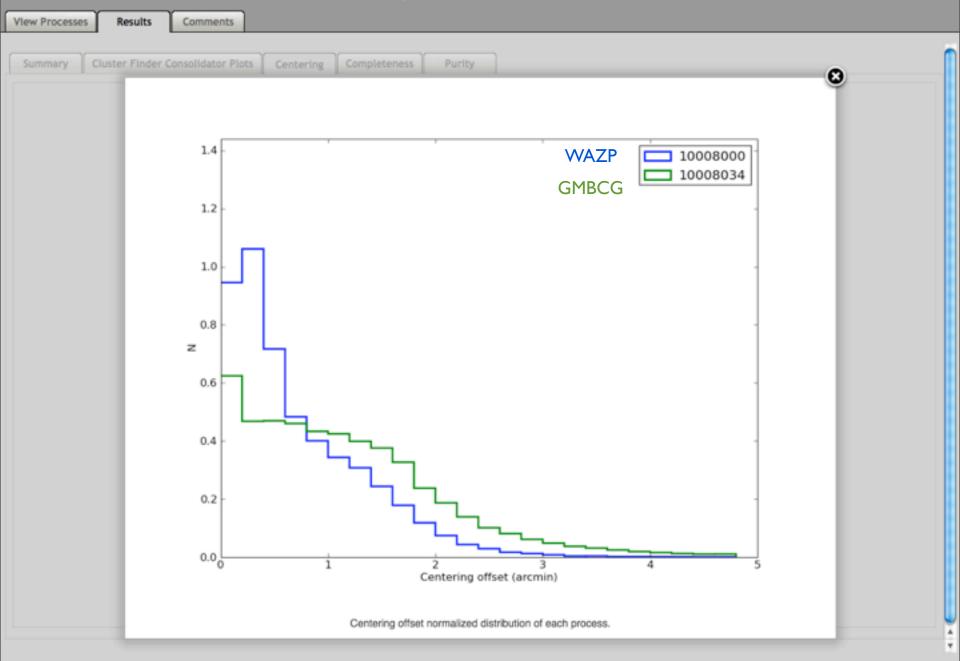
- WAZP VS GMBCG
- GMBCG searches for over-densities of red sequence galaxies + BCG. It uses photometric redshifts to assign the cluster's redshift.

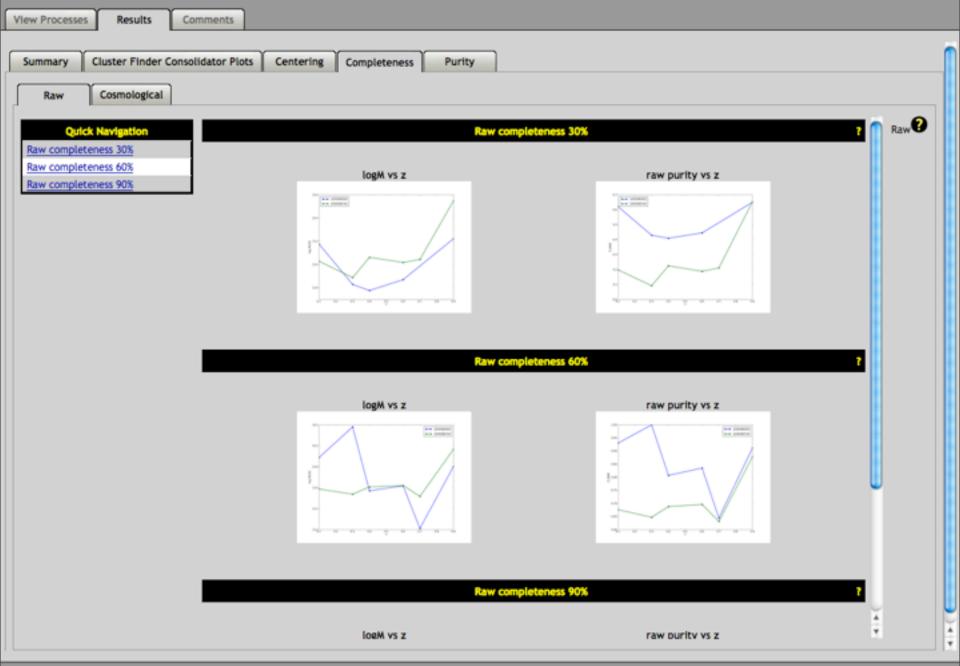


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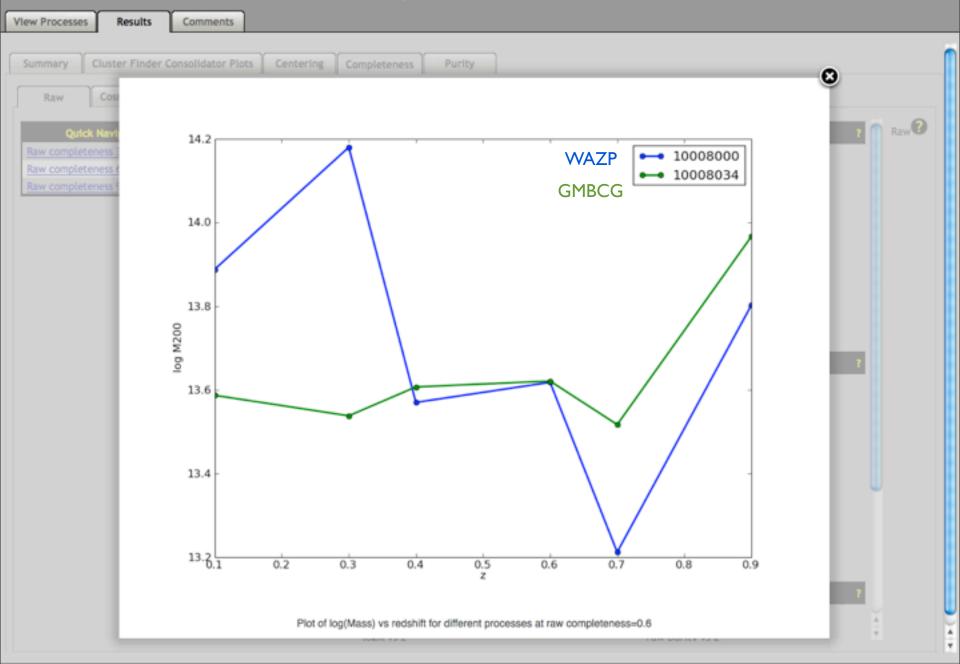
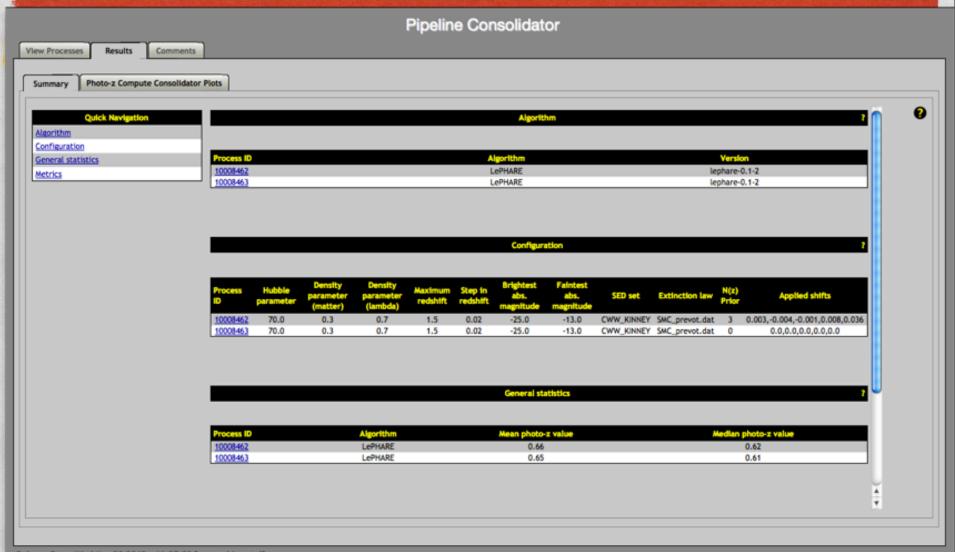
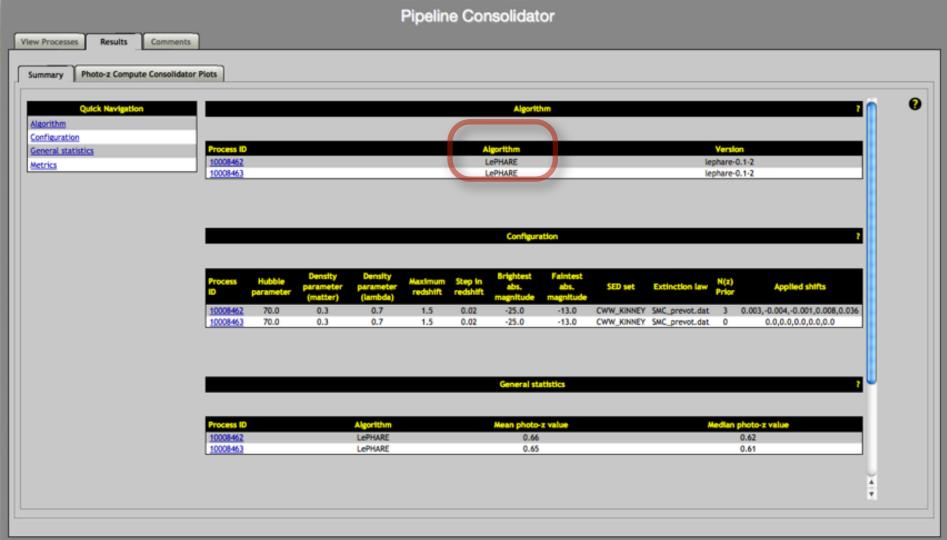


PHOTO-Z



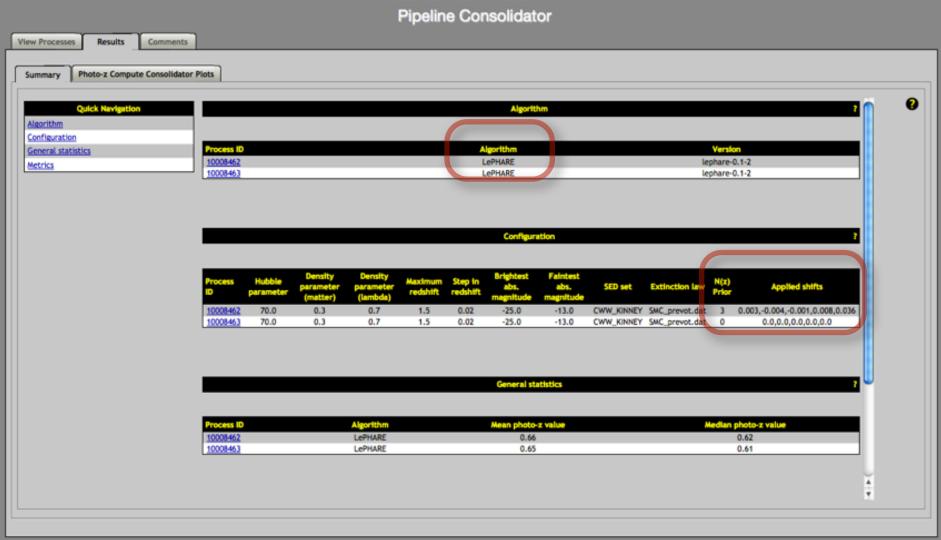
Release Date: Wed Mar 28 2012 - 16:27:02 [ostrovski portal]

PHOTO-Z



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PHOTO-Z



Release Date: Wed Mar 28 2012 - 16:27:02 [ostrovski portal]



Lessons learned



- CFs are in C, C++, Fortran, IDL, Perl, Python.
- Need special libraries.
- All wrapped by python modules
- Time to complete code wrapping depends on
 - have source code running
 - length of workflow,
 - length of configuration,
 - data needs (e.g. from FITS to ASCII),
 - language, libraries installation, and
 - understanding code.
- Some of those are first-timers' difficulties.



Conclusions



- Lots of people, codes, languages, cultures and schedules
 - Several Cluster Finders wrapped.
- Legacy. Production of cluster catalogs at the push of a button
- Pipeline Consolidator allowing easy comparison
- Future
 - close the loop with cosmology module
 - Run on real data

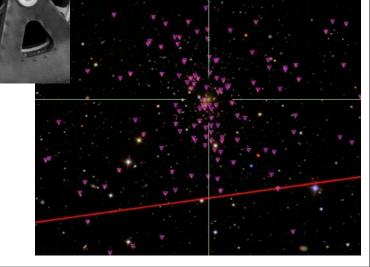


Thanks



Thanks







Extra Slides



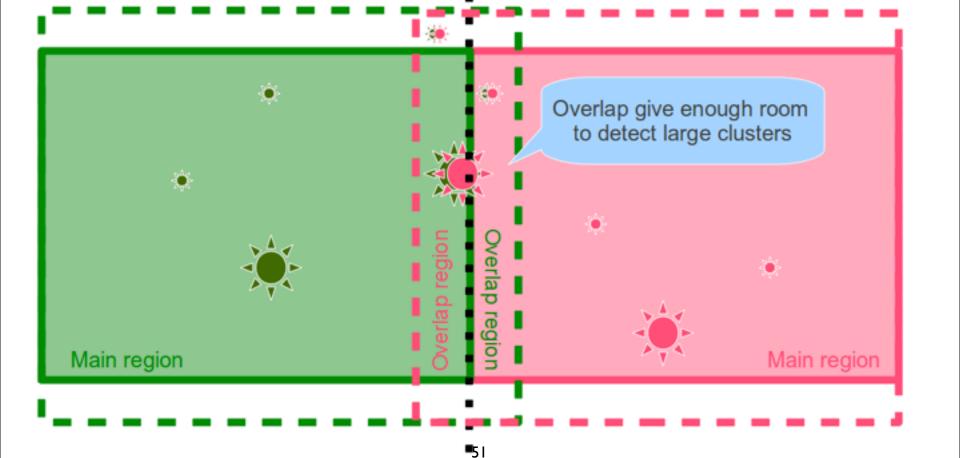
Catalog stitching

- Detection on green cell
 Detection on red cell

Only green detections kept on this side

Small centering offset on detections for display purposes

Only red detections kept on this side





the documentation.

Legacy



Brian Gerke (SLAC) was leading this project.

He left, but the code goes on. The Dark Energy Survey Science Porta The Dark Energy Survey ome Cluster Matching Data Configuration Documentation Help This pipeline matches detected clusters to true halos in a mock catalog, and vice-versa, using either shared membership or spatial proximity to make the associations. It compares the observed "masses" of the membership or spatial proximity to make the associations. It compares the observed finasses of the membership or spatial proximity to make the associations. It compares the observed finasses of the membership or spatial proximity to make the associations. It compares the observed finasses of the membership or spatial proximity to make the associations. It compares the observed finasses of the membership or spatial proximity to make the associations. It compares the observed finasses of the membership or spatial proximity to make the associations. It compares the observed finasses of the membership or spatial proximity to make the associations. It compares the observed finasses of the membership or spatial proximity to make the associations. It compares the observed finasses of the membership or spatial proximity to make the associations. It compares the observed finasses of the membership or spatial proximity to make the associations. Eli Rykoff ne to the Testing rank-ordering produced by the cluster-finding algorithm. (In essence, at a given redshift gets the mass we Cluster Matching Configuration Documentation Help would expect for the top-ranked halo in that bin, and so on.) The fraction of matches, and the accuracy of Data The Dark Energy Survey ketease notes Science Porta Description Matching Known Bugs Home Cluster Matching Data Configuration Documentation Help Outputs Matches clusters detected in a mock catalog and vice-versa. Output of the cluster matching pipeline FAQ ect it (one Code by Brian Gerke (SLAC) and wrapping by in Gerke. References The code outputs a set of plots, as well as a FITS file containing the matching information for the halos and clusters and a summary of quality-assessment statistics. portal domain. An easy Output Plots In order to use this pipeline, first you should Credits hange this is to use the Upload tool that can be accessed on Data Server > Import > Products. It is also The plots in the output log show the following: re-use the results of a previous run. See the Documentation for more details. The Cluster-Halo Comparison tab atter plots of M_obs versus M_true for the case of The initial input catalog must have the following format. It shot The Dark Energy Survey non-unique matching, small points show matches primary (zeroth) HDU (since this HDU is reserved for image data Science Porta HDU should be a catalog of clusters ranked by some mass proxy catalog must have the following tags: 'RANK', 'RA', 'DEC', 'Z'. The Louise Matching Data Configuration Documentation Help that belong to the clusters. It MUST contain tags 'ID' (the ID num scatter plots of the centering offsets for the the cluster to which it belongs). It can also optionally contain a d in redshift space. probability that this galaxy belongs to the cluster with the sped omitted if cylindrical matching is being used. shows the mass-rank relation in the truth table (blue · For now, the relation between angular separation and comoving distance assumed for computing the cylinder in cylindrical matching assumes LCDM with Omega M=0.3. This should be changed to in bins of redshift. It is advisable to check this plot incorporate the appropriate value of Omega_M for the mock being used. However, any impacts should The pipeline produces a product log with several plots and one be small for reasonable Omega M values.

Linea Parallelization in the portal



