

Deconvolution of fiber spectra in IFU spectrographs

SIFS data reduction package

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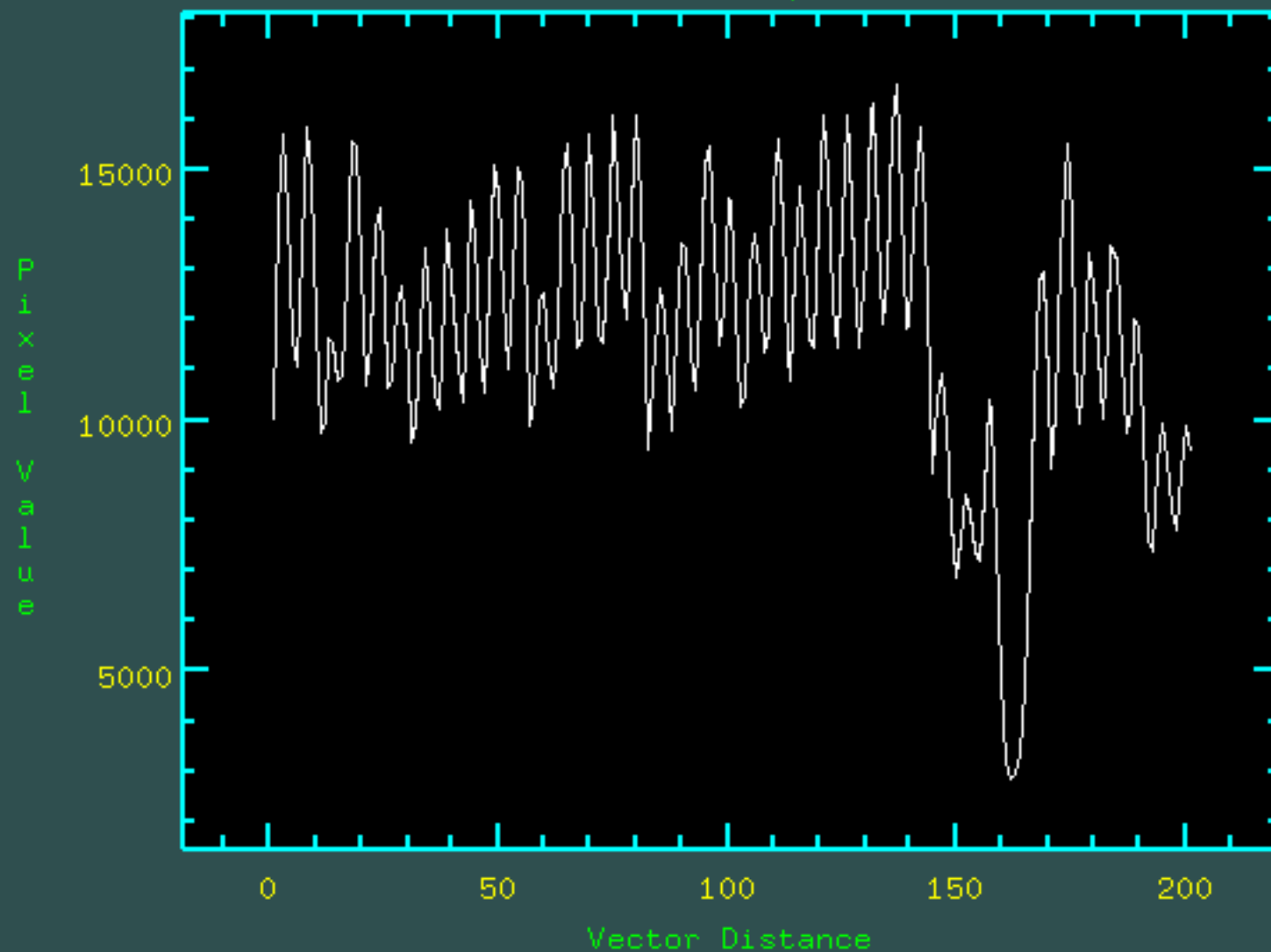
The problem

Fibers are very packed and the spectrum of one superposes to the other in the spatial direction.



X xgterm <2>

NOAO/IRAF V2.12EXPORT kanaan@chamaeleon Mon 19:35:42 17-Nov-2003
Fla0001: Vector 1028.0,1463.0 to 1028.0,1663.0 naverage: 1
Flat Cup



The Solution

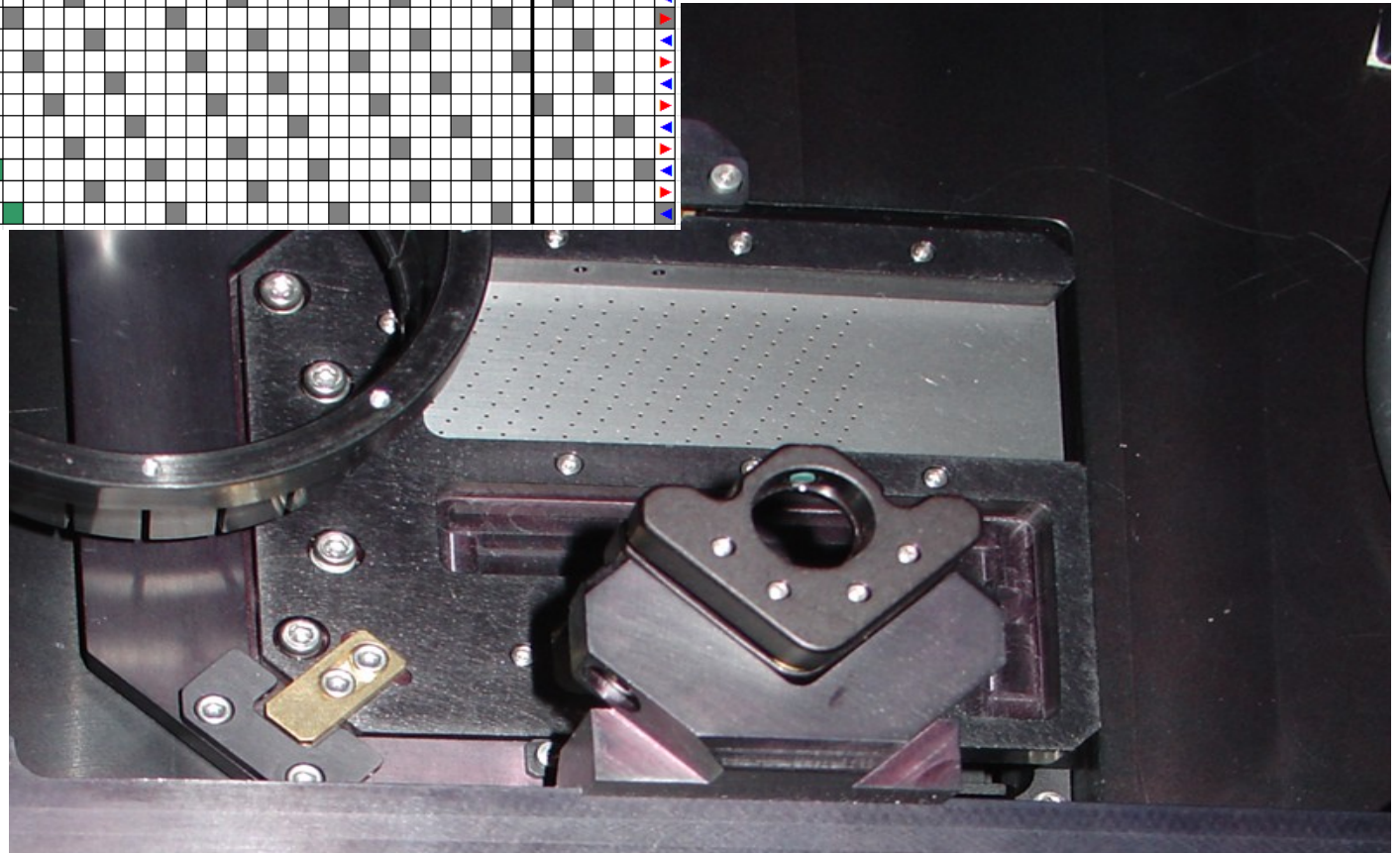
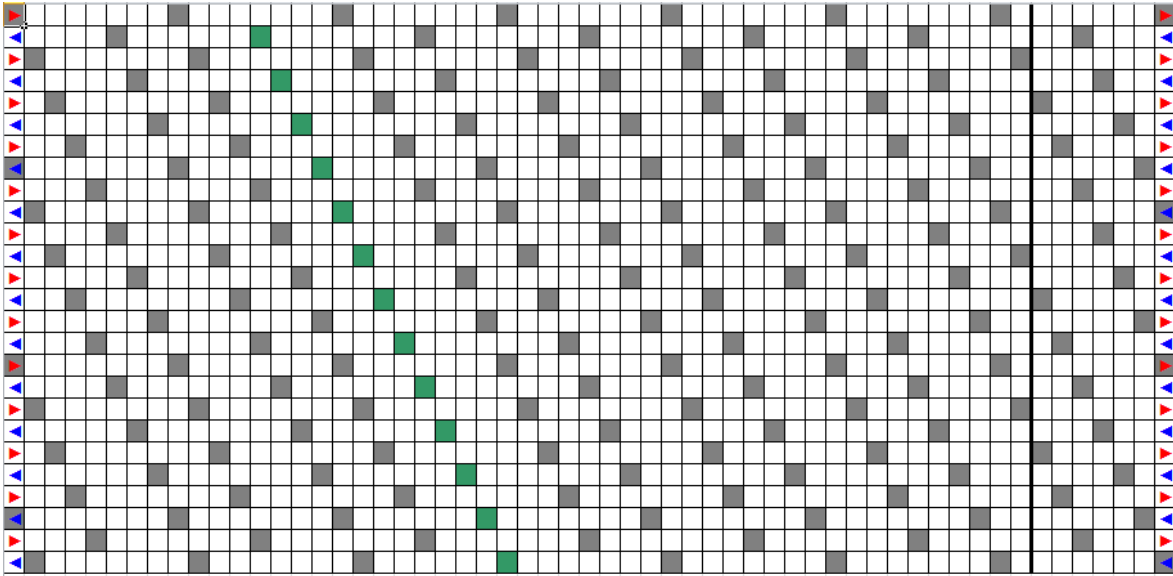
If (and only if), we know the position and width of the spectrum, of each fiber for all wavelengths on the CCD, then it is possible to determine the amount of light incident upon each fiber.

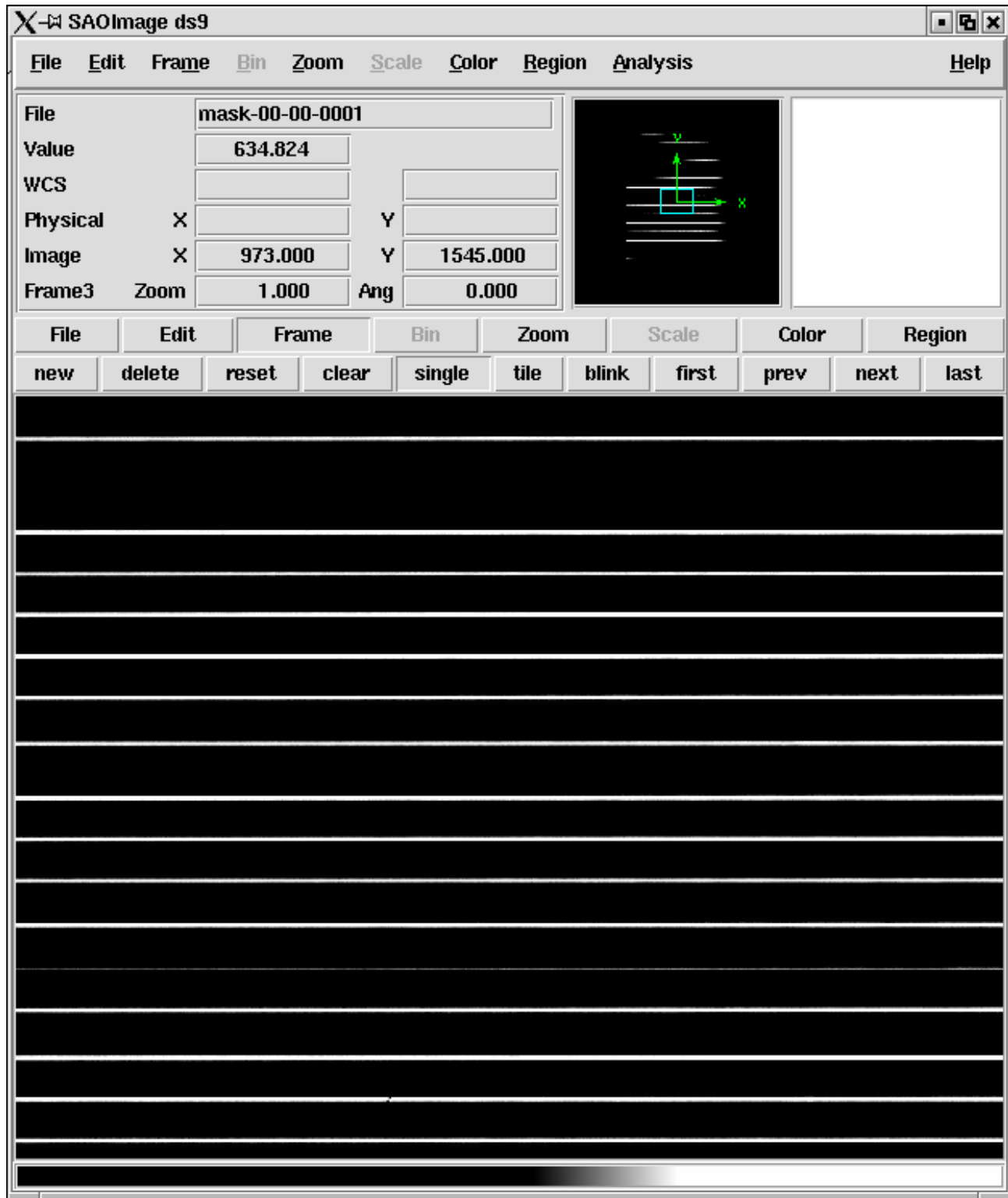
We can do that by fitting the amplitude of Gaussian profiles to column cuts (perpendicular to dispersion) across the detector. The number of Gaussians in the fit equals the number of fibers.

As the amplitude is the only free parameter the fit is linear (fast and with a unique solution).

SIFS fore-optics mask

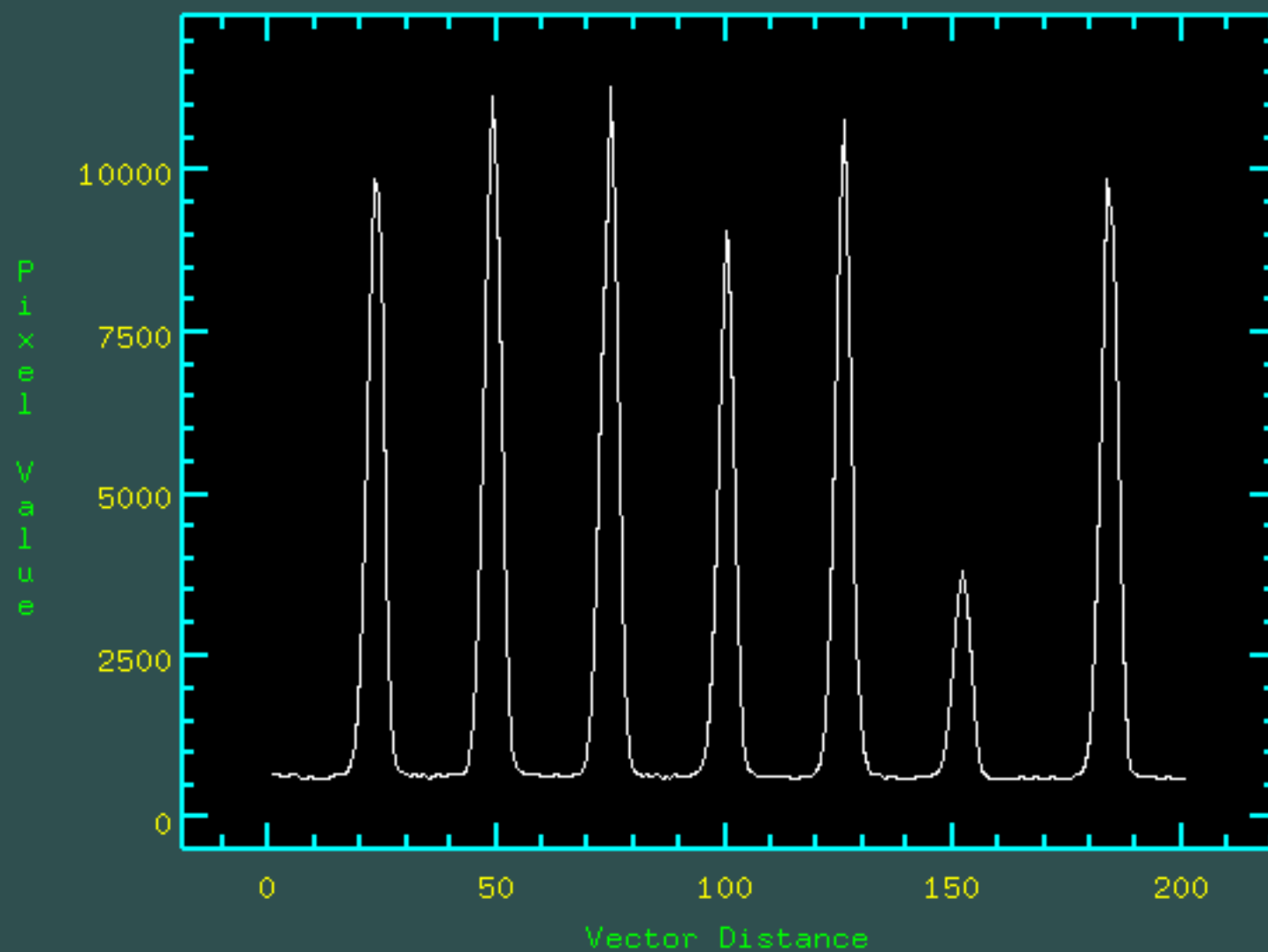
* masked flat as an extra calibration step



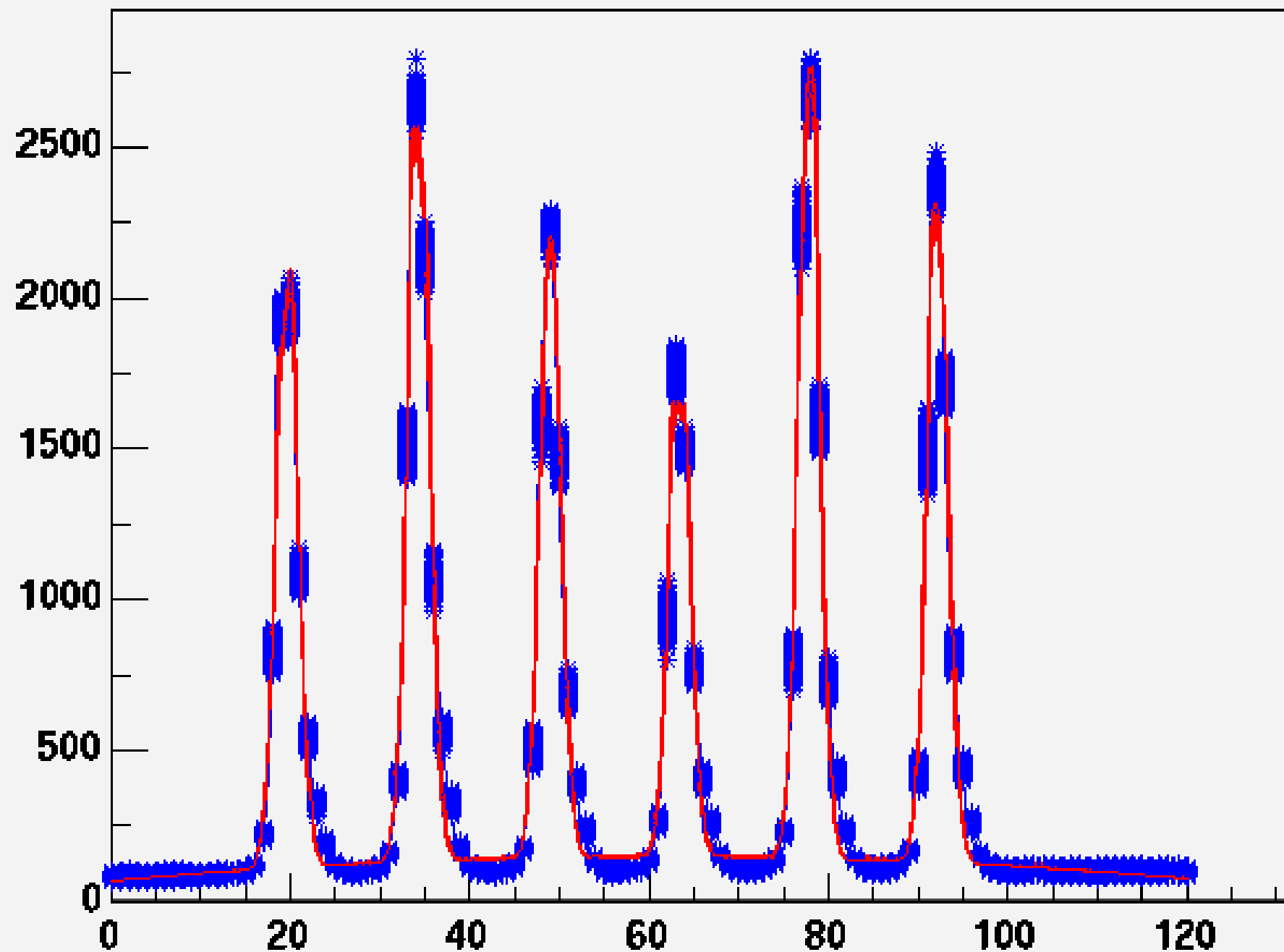


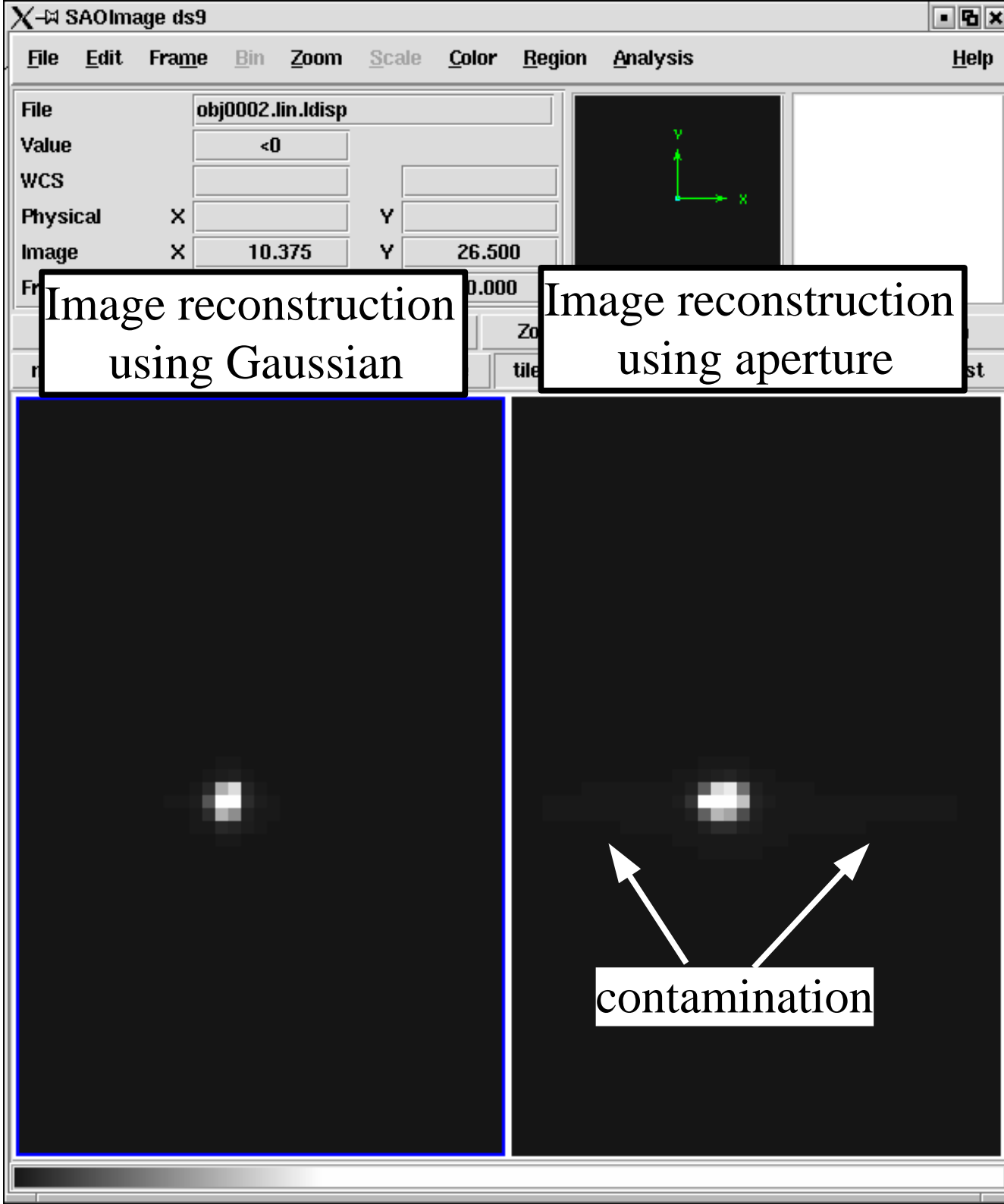
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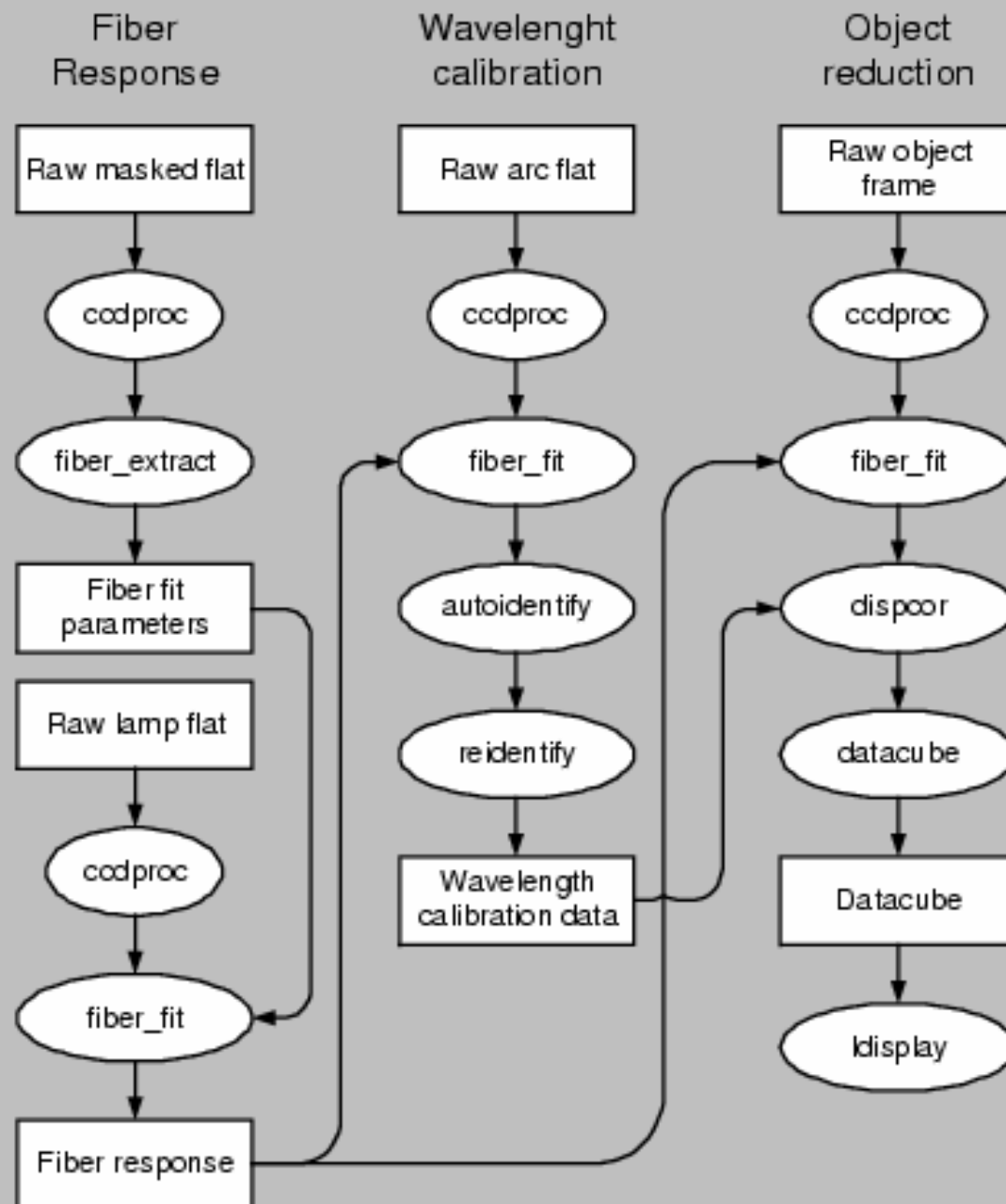
NOAO/IRAF V2.12EXPORT kanaan@chamaeleon Mon 19:40:40 17-Nov-2003
mask-00-00-0001: Vector 1028.0,1463.0 to 1028.0,1663.0 naverage: 1
Mascara Pos 1



Blue - data / Red - Fit







A complete IRAF package is currently in use for EUCALYPTUS data reduction.

<http://www.astro.ufsc.br/~kanaan/ifu>

The software will need some adaptations to run smoothly on SOAR, but it is **modular** enough that most fundamental differences can simply be reconfigured by changing configuration files.