

Instrument Development

STELLES



SIFS



Our Problem

Steve's Problem

Your Problem

- ← Concept
- ← CDR
- ← PDR
- ← FDR
- ← Construction
- ← Lots of fun
(weeping and teeth gnashing)
- ← P.s.Tests
- ← Engineering Com.
- ← Science Com.
- ← Early Science
- ← Shared Risks
- ← Science
- ...
- Junkyard

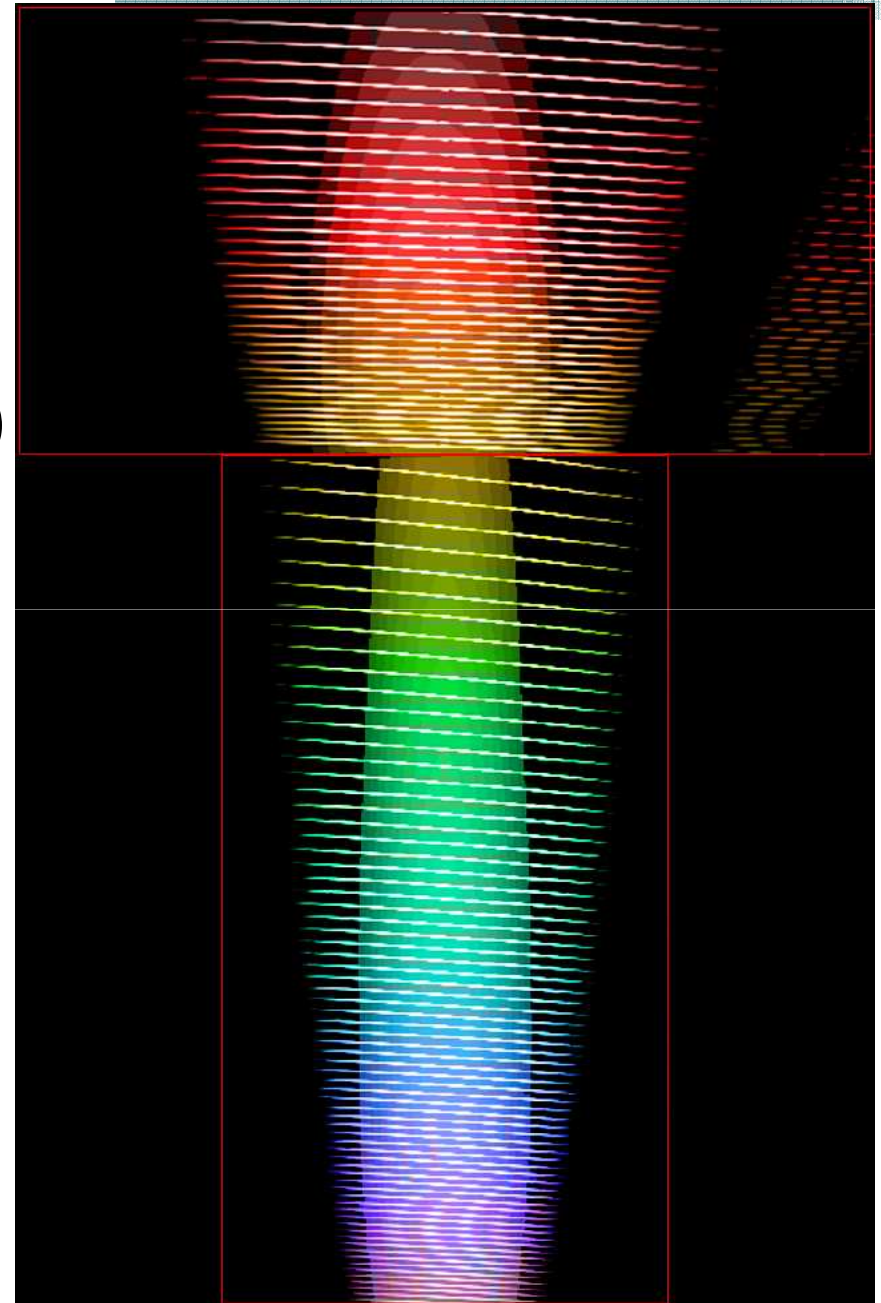


Observing with STELES

Bruno Castilho

STELES

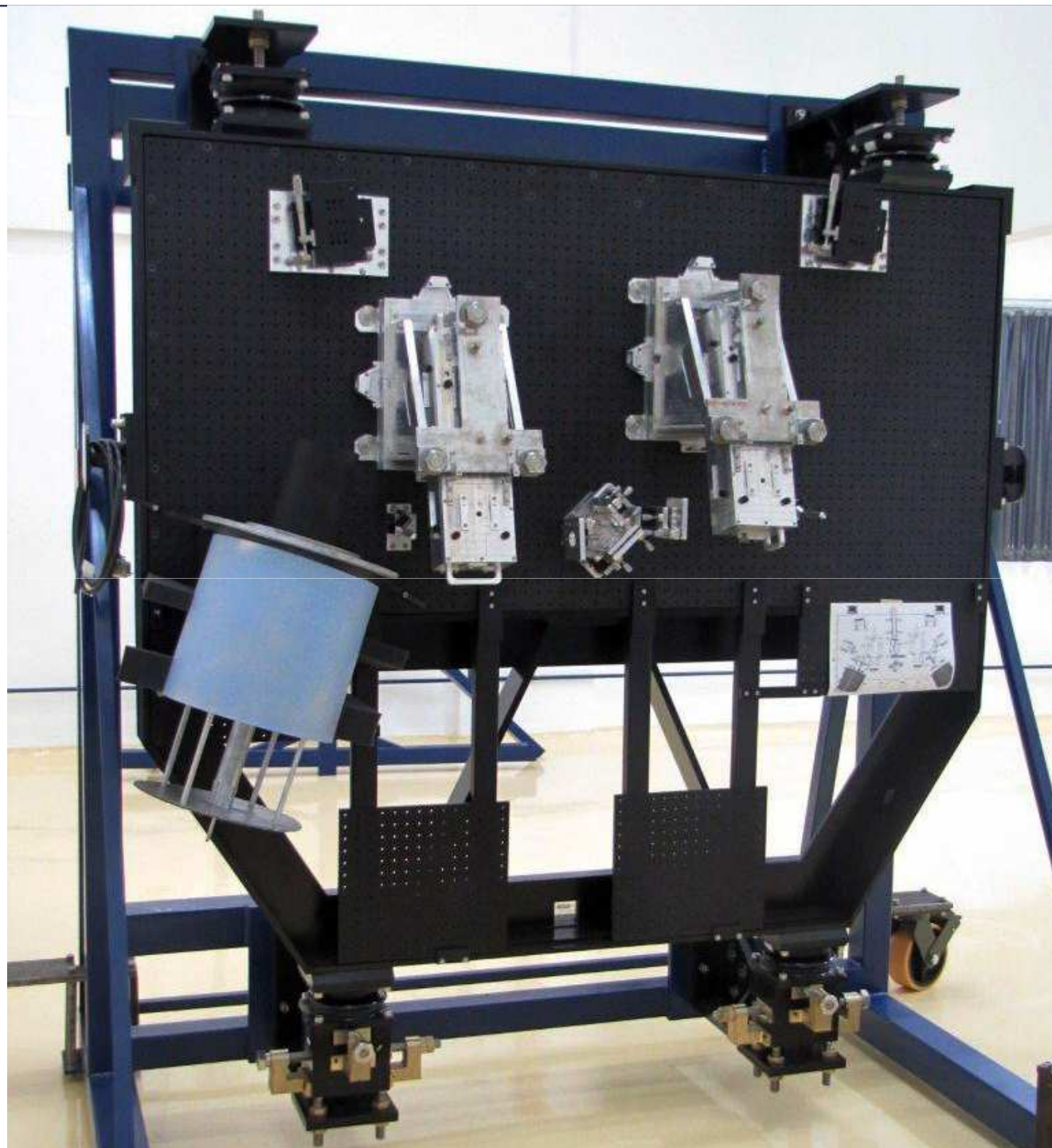
- $\lambda\lambda$ - 300-890nm (in one shot)
- R4 echelle gratings
50k (75K narrow slits)
- 2 independent slits
0.8" (0.5-5")
8.0" (0.5-20")
- CCD - 2 x 4096x2048
 - graded coating
- Reduction pipeline





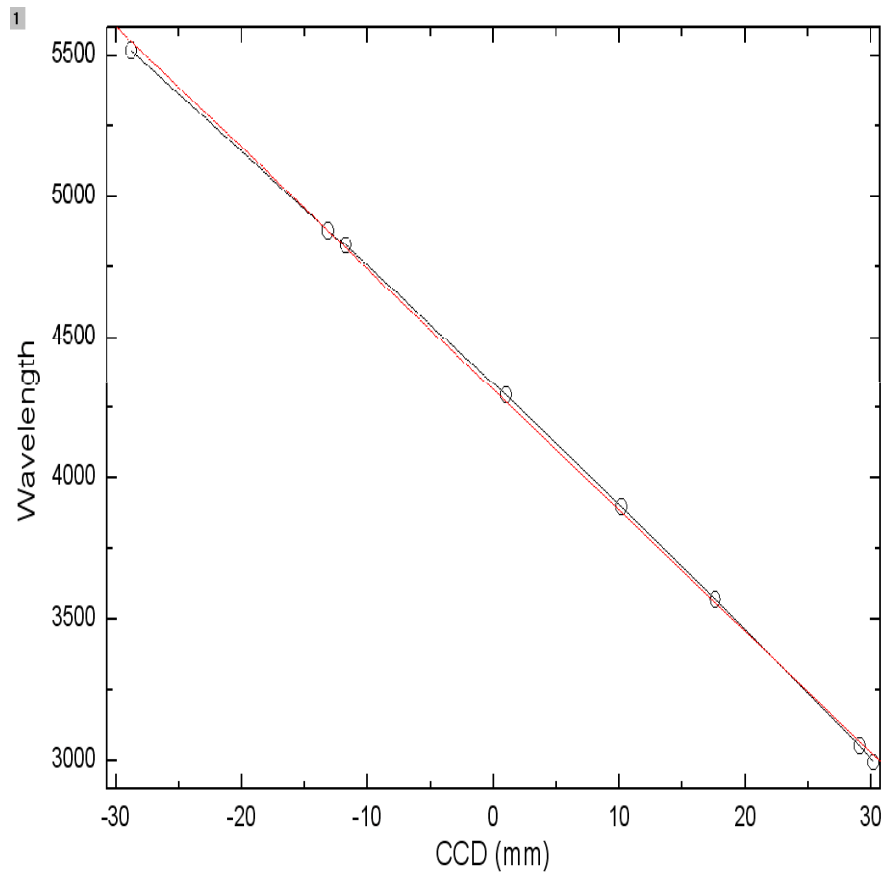
Status

- Optics 100% - arrived, to be tested
- CCDs 100% - at CTIO, in characterization
- Dewars design finished, to be fabricated at CTIO
- Mechanics fabrication 50%
 - imported parts 80% arrived, 20% ordered
 - in fabrication at LNA and contractors
- Electronics 75% arrived
- Data reduction - see poster 26

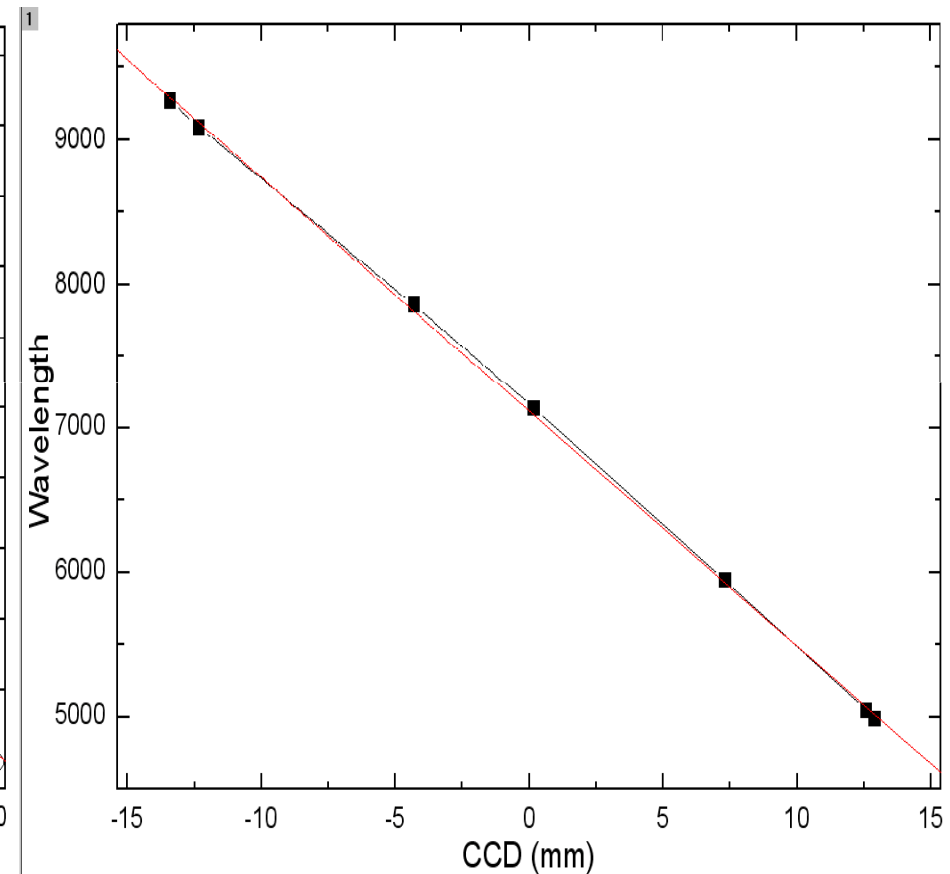


Spectral format / Detectors

Gradient Coating Format



Blue chip



Red chip

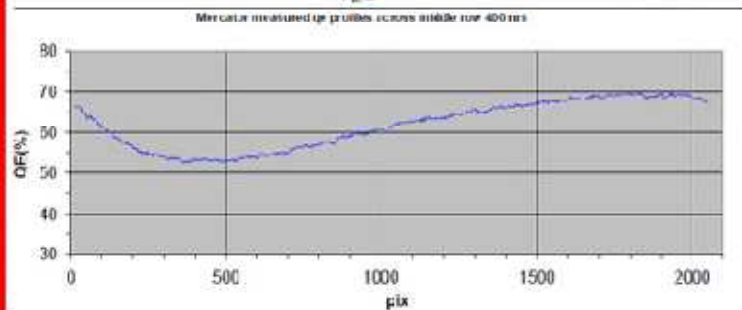
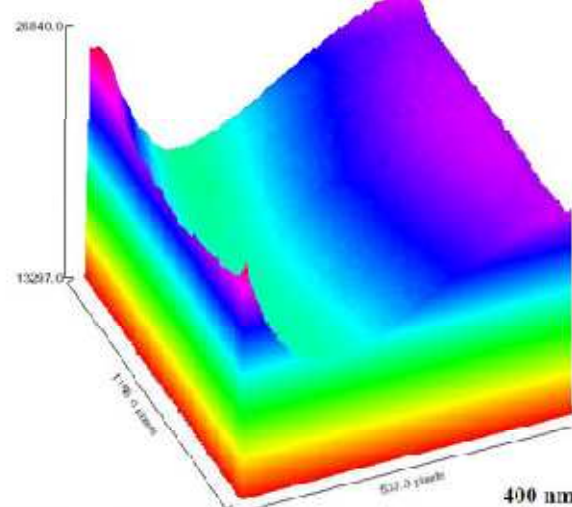
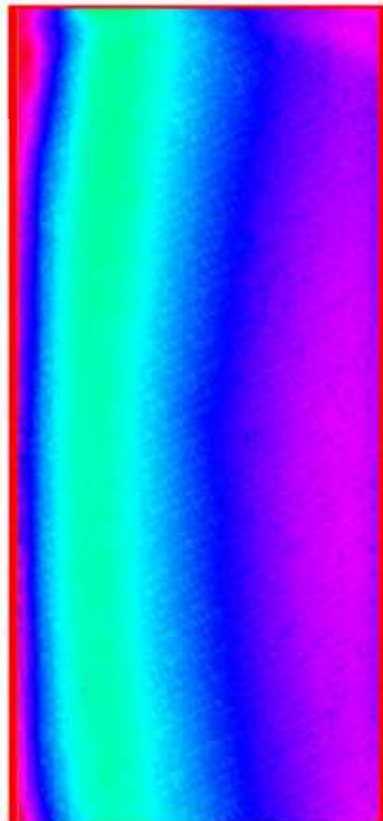
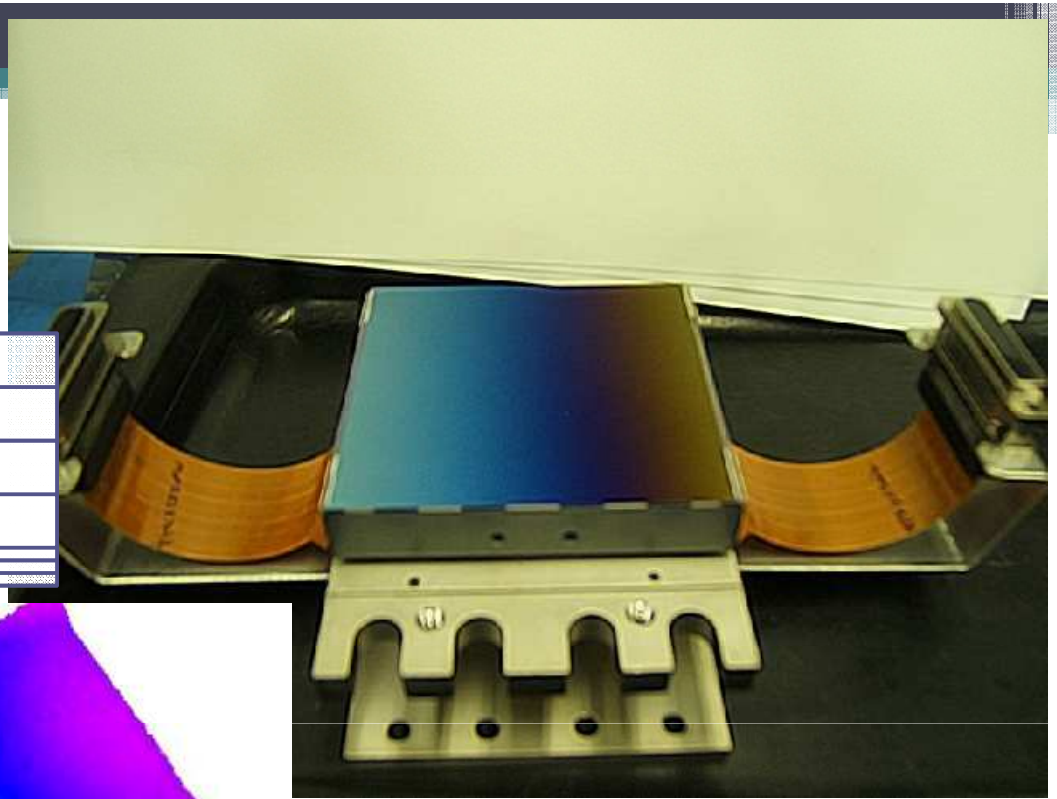
Spectral format / Detectors Results for coating

Other consumers:

CCD42-90 2kx4k5 pixels to Mercator telescope

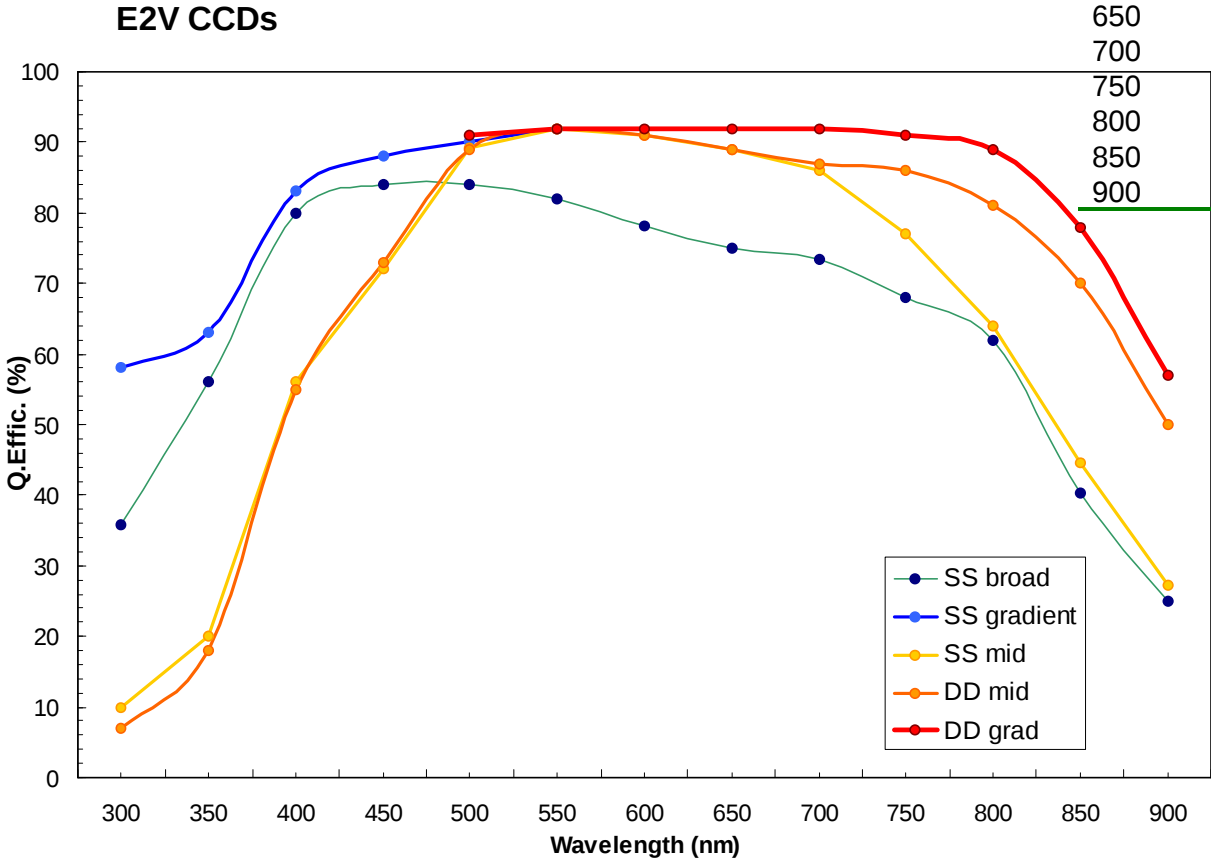
CCD231-84 4k4k pixels to SAO (Harvard)

CCD231-84 set of 25 devices for ESO MUSE project



Spectral format / Detectors

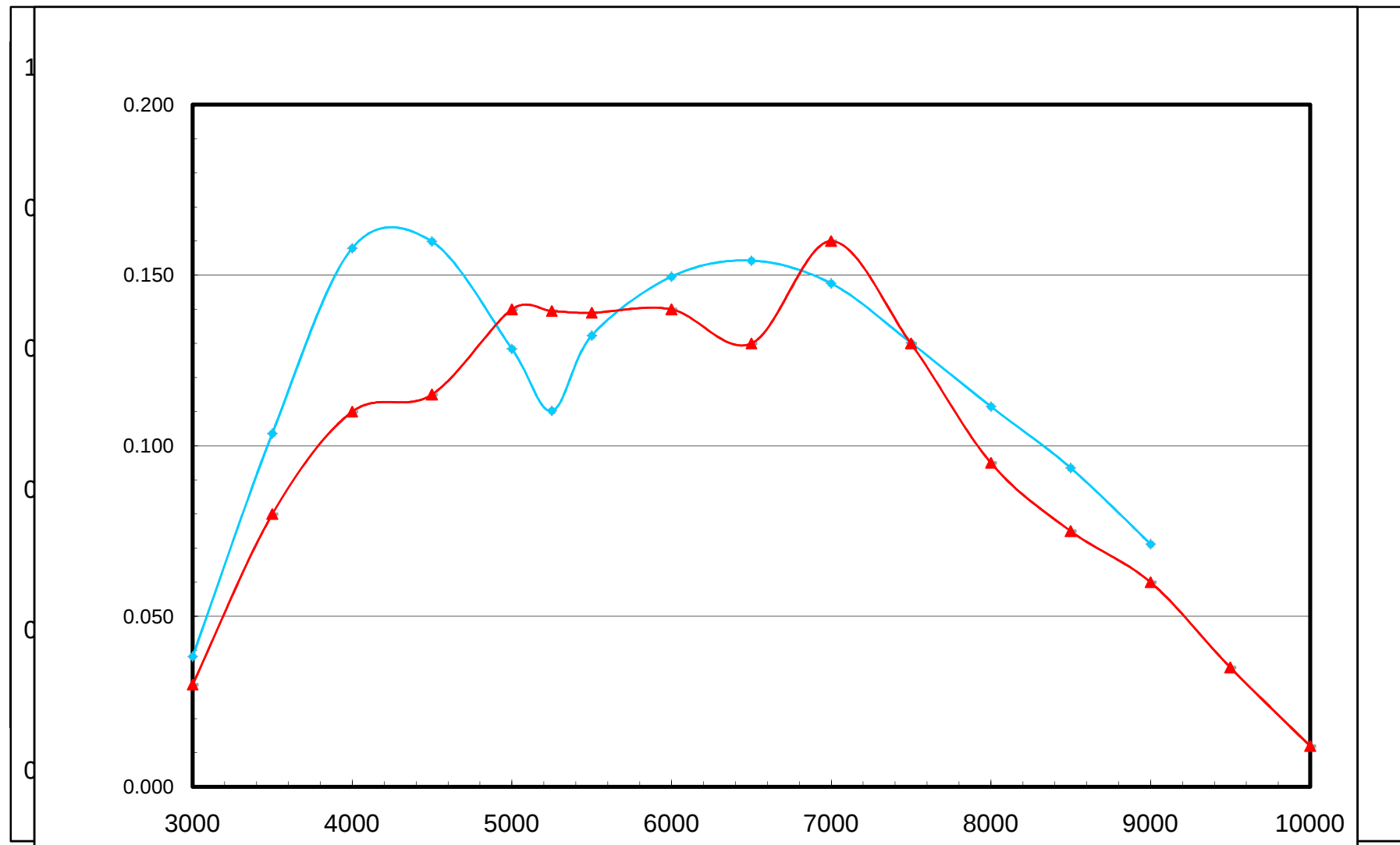
Expected QEf



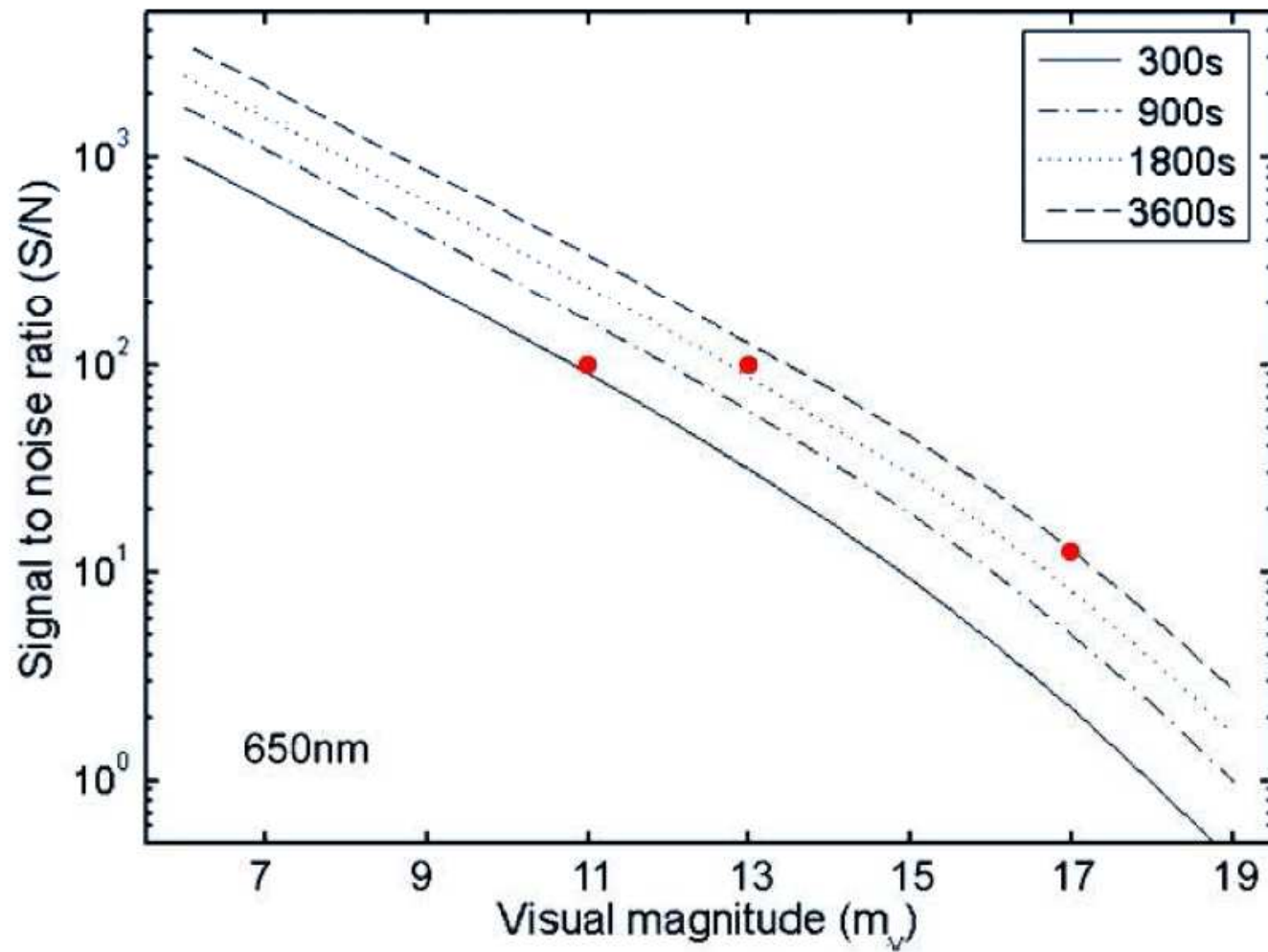
	E2V 44-82 -100deg				
	SS broad	SS grad	SS mid	DD mid	DD grad
300	36	58	10	7	
350	56	63	20	18	
400	80	83	56	55	
450	84	88	72	73	
500	84	90	89	89	91
550	82	92	92	92	92
600	78		91	91	92
650	75		89	89	92
700	73		86	87	92
750	68		77	86	91
800	62		64	81	89
850	40		45	70	78
900	25		27	50	57

Based on E2V test sheets
and previsions

Efficiency



Efficiency



M v	S/N	T(min)
11	100	5
13	100	30
17	12	60

650nm
0.8"
dark time

STEELES

Exposure Time Calculator

SOAR Telescope Echelle Spectrograph

[Online help for SPECTIME](#)

Signal-to-noise ratio: Integration time in seconds:

Maximum time per single exposure (seconds):

Central wavelength: From 3600 to 5399 [Å] data is computed for Channel A
From 5400 to 9000 [Å] data is computed for Channel B

Slit width in arcsecs (+) or pixels (-): Dichroic:

Filter:

Detector binning (dispersion):

Seeing (arcsecs): Airmass:

Lunar Phase (0-14 days from new moon, 0 = new, 14 = full):

Describe the source with the following parameters. The power law index is only used if the source spectrum is set to Fx_power or Fv_power. Setting the power law index to 0 is equivalent to a flat spectrum for that particular power law function. If the value of the wavelength of the source flux is set to INDEF then the central wavelength value above is used. The wavelength of the source flux is ignored for source units = U, B, V, R, I.

Source spectrum: Temperature (K) for blackbody source: Power law index:

Wavelength (Å) of source flux or magnitude: Source flux or magnitude: Units of source flux or magnitude:

E(B-V) (reddening of source in magnitudes):

Detector: [Help for Detector Type](#)

sedivama@efei.br

sptime

Object spectrum: Blackbody spectrum of temperature 6000. K
Reference wavelength: 5500. Angstroms
Reference flux: $V = 12$. ($6.03E-14$ ergs/s/cm²/Å)
Sky spectrum: Cerro Pachon
Moon phase: 0
Extinction:
Airmass: 1.
Seeing: 0.8" (FWHM)

Telescope: Slit
Area: 13.2 m², Scale: 3.144 arcsec/mm
Spectrograph: STEELES
Collimator:

Focal length = 0.85 m
Apertures: Slit with Gaussian profile
Size: 0.80" x 21.63", 0.254 x 6.88 mm, 4.2 x 100.0 pixels

Filter: Dichroic A
Disperser: Echelle
Central order = 104
Central wavelength = 4500. Angstroms
Central dispersion = 2.66 Angstroms/mm, 0.0399 Angstroms/pixel
Ruling = 41 lines/mm
Blaze = 76.0 deg
Grating tilt = 76.1 degrees
Grating magnification = 0.96
Using predicted efficiencies
Crossdisperser: VPH

Central order = 1
Central wavelength = 4500. Angstroms
Central dispersion = 2.66 Angstroms/mm
Ruling = 1200 lines/mm
Blaze = 14 deg
Grating tilt = 16.2 degrees
Grating magnification = 0.86
Using tabulated efficiencies

Camera:
Focal length = 0.215 m, Resolution = 3. pixels
Detector: MIT/LL
Pixel size: 15 microns, 0.19"
Number of pixels = 2048
Read noise = 0.0 e-, Gain = 1.0 e-/DN, Dark = 0. e-/s

Source pixels: 14 pixels (3x FWHM seeing disk)
Background pixels: 89 pixels

WARNING: Orders overlap

---- Results for 919.61s exposure at 4500. Angstroms ----

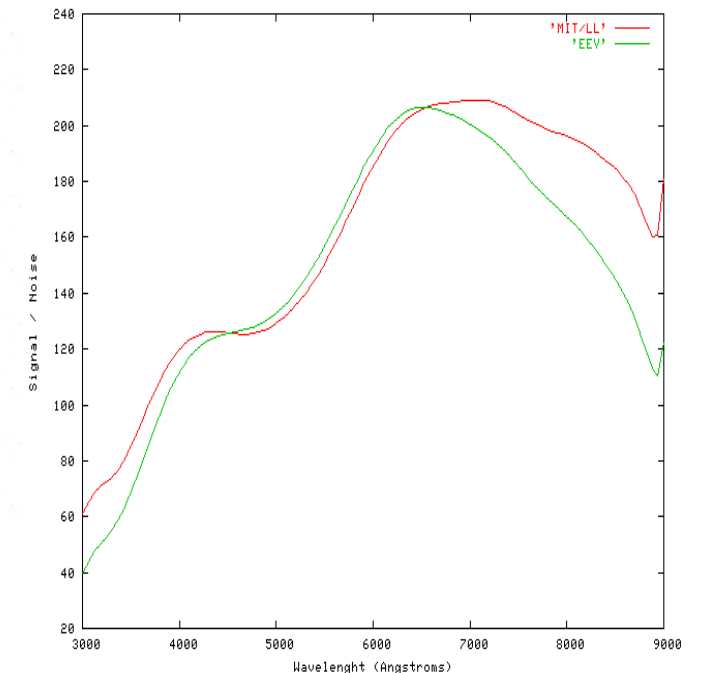
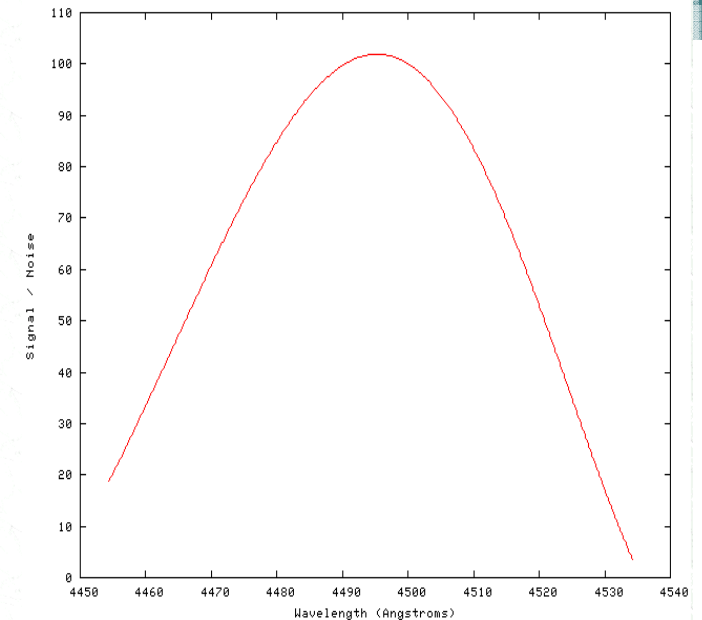
Source flux: 0.0141 photons/s/cm²/Å (AB=12.4)
Background flux: 2.71E-6 photons/s/cm²/Å (over source pixels)

Transmission/Efficiencies:	individual	cumulative
Telescope	75.8%	75.8%
Aperture (seeing=0.8")	62.4%	47.3%
Filter	97.6%	46.2%
Spectrograph	44.3%	20.5%
Disperser	84.2%	17.2%
Detector DQE	87.0%	15.0%

Statistics per exposure per pixel:	e-	sigma
Source	10001	100.0
Background	1	1.0
Dark	0	0.0
Readout	0	0.0
Net	10002	100.0

Signal-to-Noise Statistics:
100 per exposure per pixel
223 per exposure per 0.208 Angstroms (5 pixel) resolution element

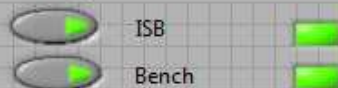
Exposure time: 919.61s (0:15:19.6)





SOAR Telescope Echelle Spectrograph

Energy Status



STELES fold prism **IN**
STELES

TCS Information

R.A.	Dec
17:07:10.261	-30:13:35.702
Airmass	Focus
1.20	-1168.00
Mount AZ	Mount El
67.1234	56.1234
Dome AZ	Dome Shutter
90.1234	0.0036
Date	Sideral Time
2011-05-18	22:51:06.4
LIT Time	Hour Angle
15:45:41.6	-00:05:04.0
Outside Temp	Wind Direction
20.1234	350.1234
Pressure	Wind Speed
742.1234	10.1234
Seeing	Humidity
1.1234	35.1234

Setup Guider Observ

Base Name
Title
Path %
Coment
Observer

Blue channel

Exp. Time 2 1200

Repeat 3 1

Go

Exposure Progress

ReadOut Progress

Write Progress

Stop Sequence Abort

Blue channel

Exp. Time 3 1200

Repeat 3 1

Go

Exposure Progress

ReadOut Progress 2

Write Progress 2

Stop Sequence Abort

Calibration

Telescope

STELES

Calibration mirror

OUT

IN

Milk Flat Lamp

Quartz + De

ThAr + Ne

Calibration Script

Afternoon Call (Bias+Flat+ThAr+Milk)



ExpT Quartz (s) ExpT ThAr (s) ExpT Milk (s)

0

0

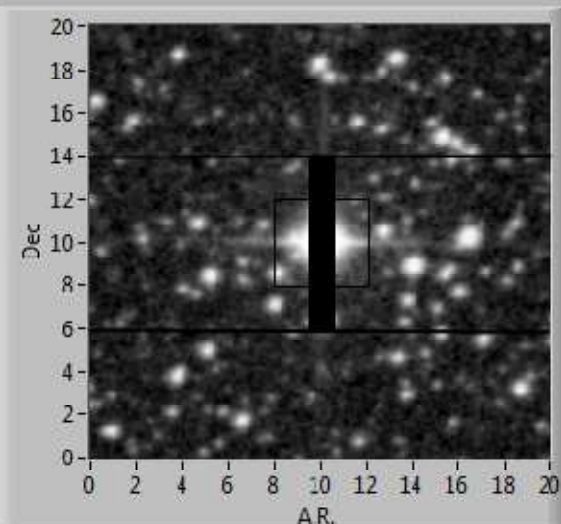
0

Slit Cameras

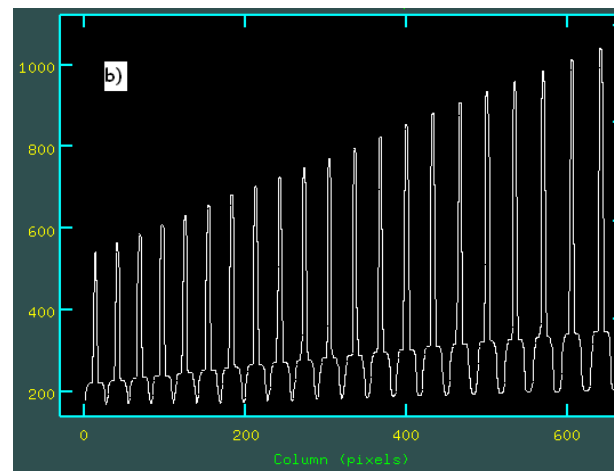
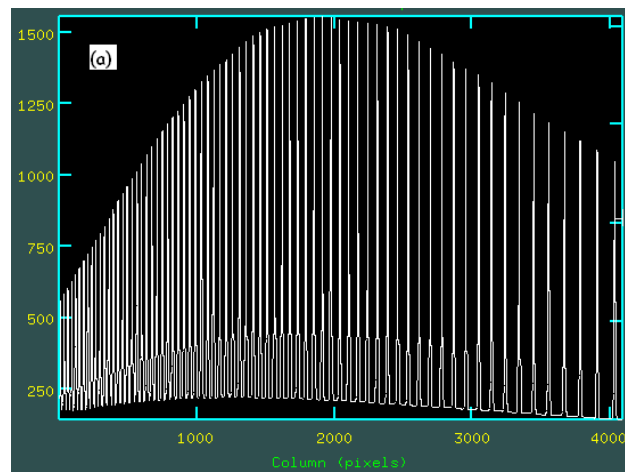
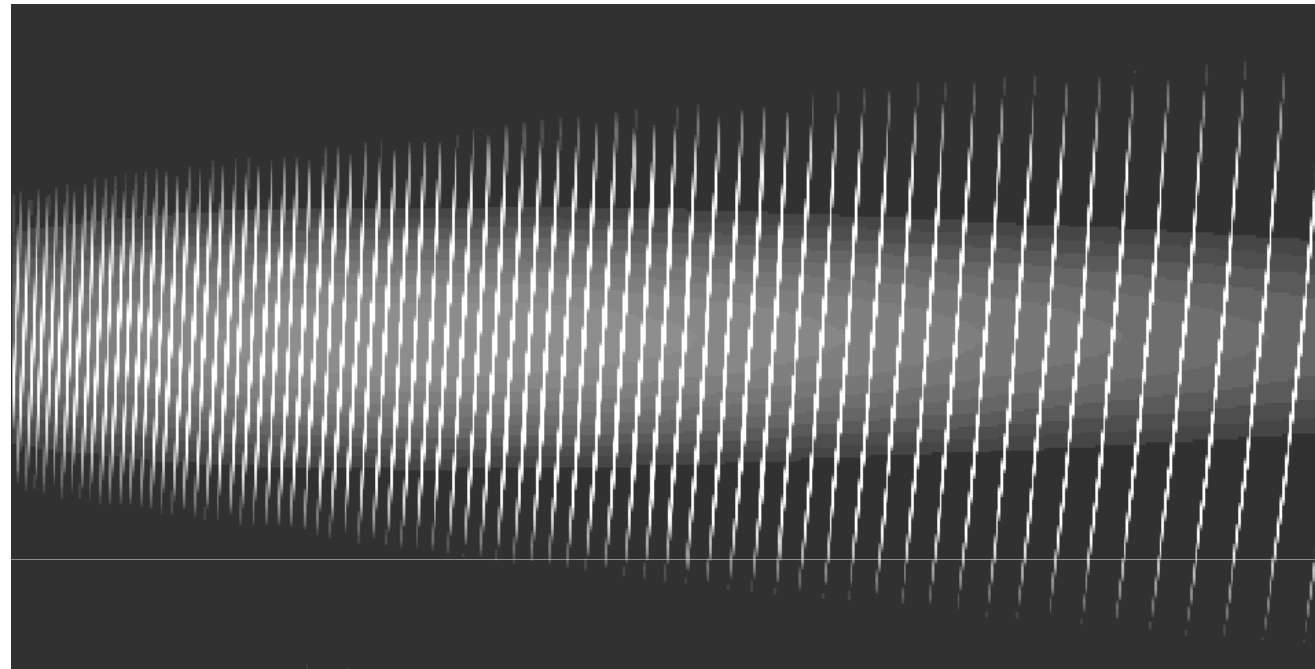
Blue

Red

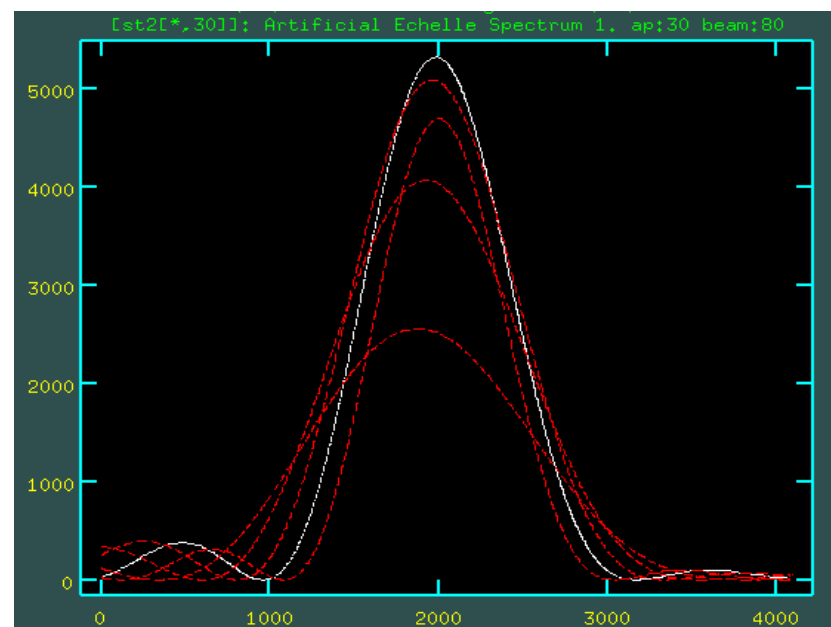
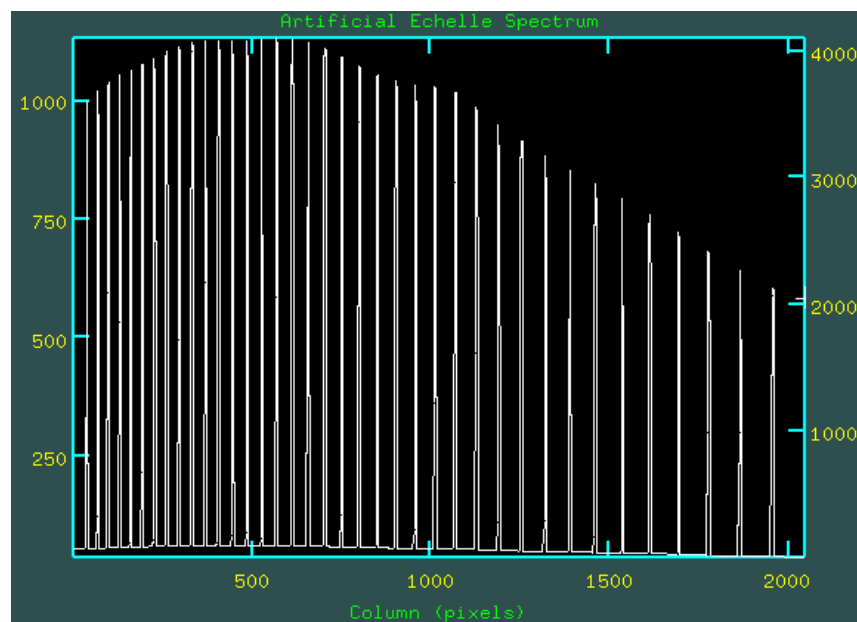
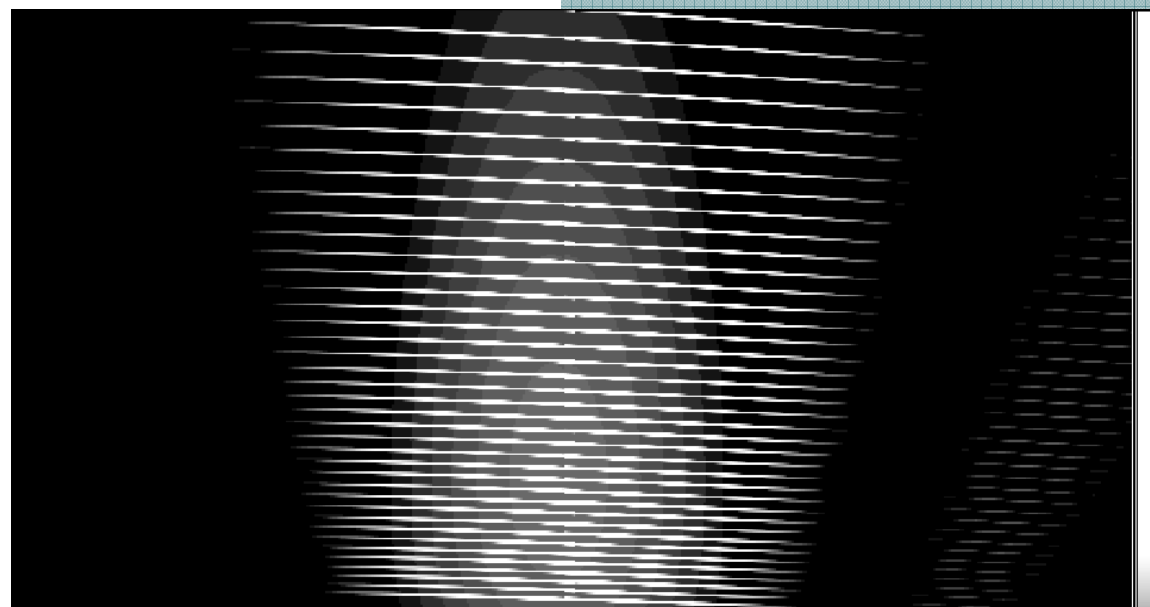
Zoom



Blue



Red

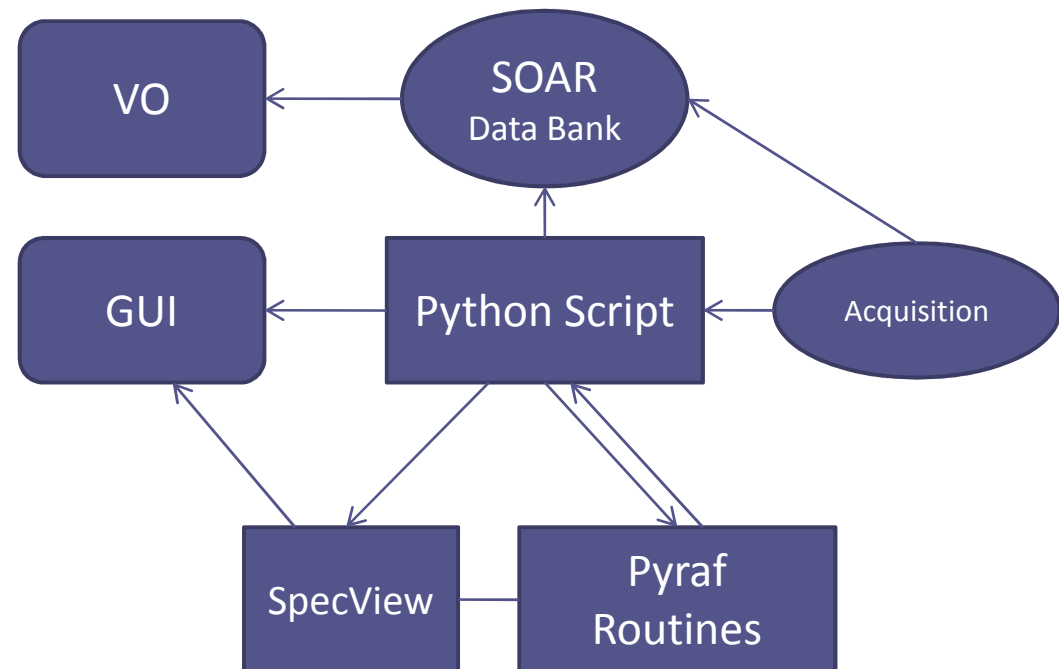


Data Reduction - Structure

First light – IRAF script, tests and early science

Pine line – optimized based on the first data

Poster 36
M.B. Fernandes





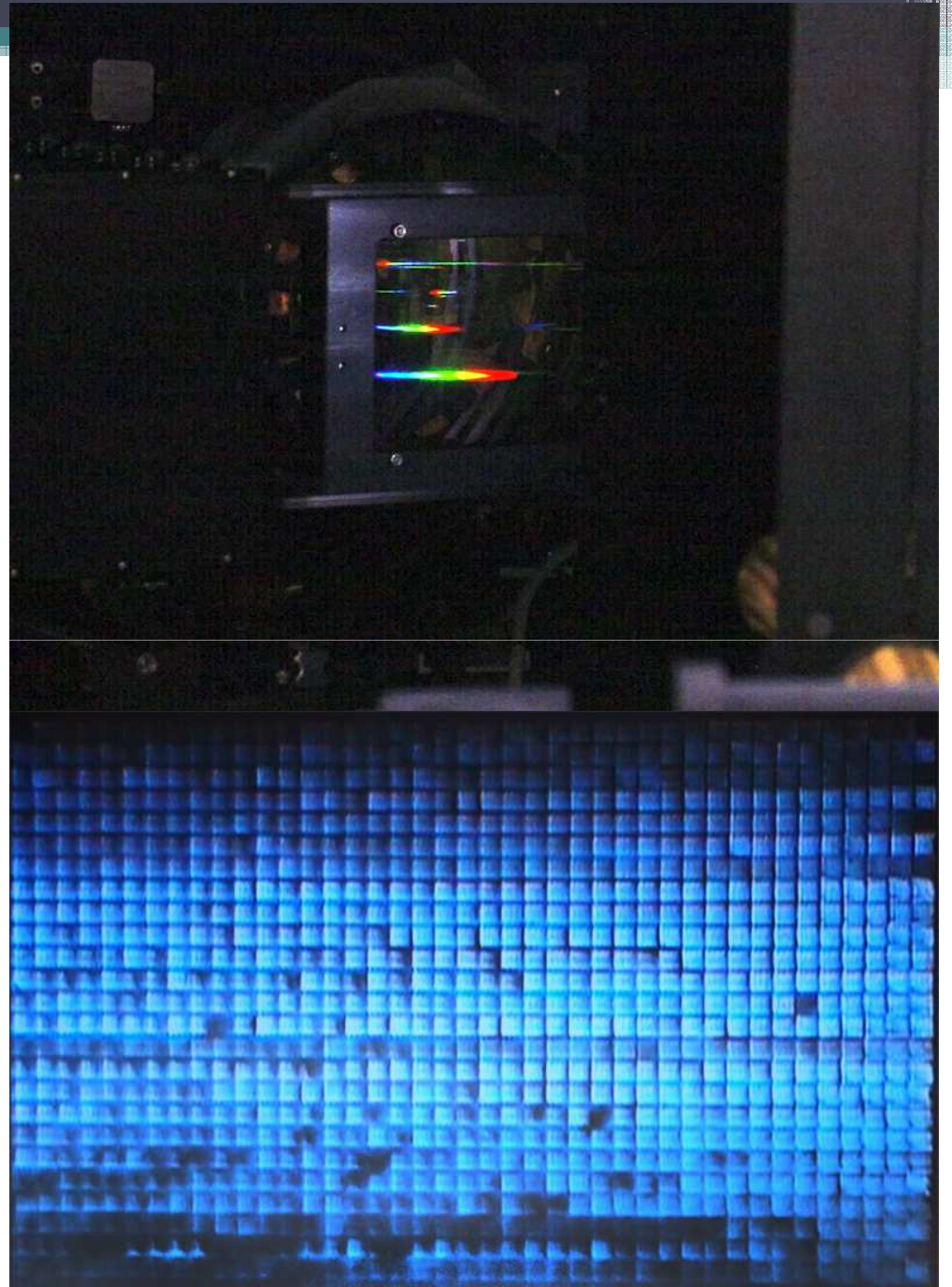


Observing with SIFS

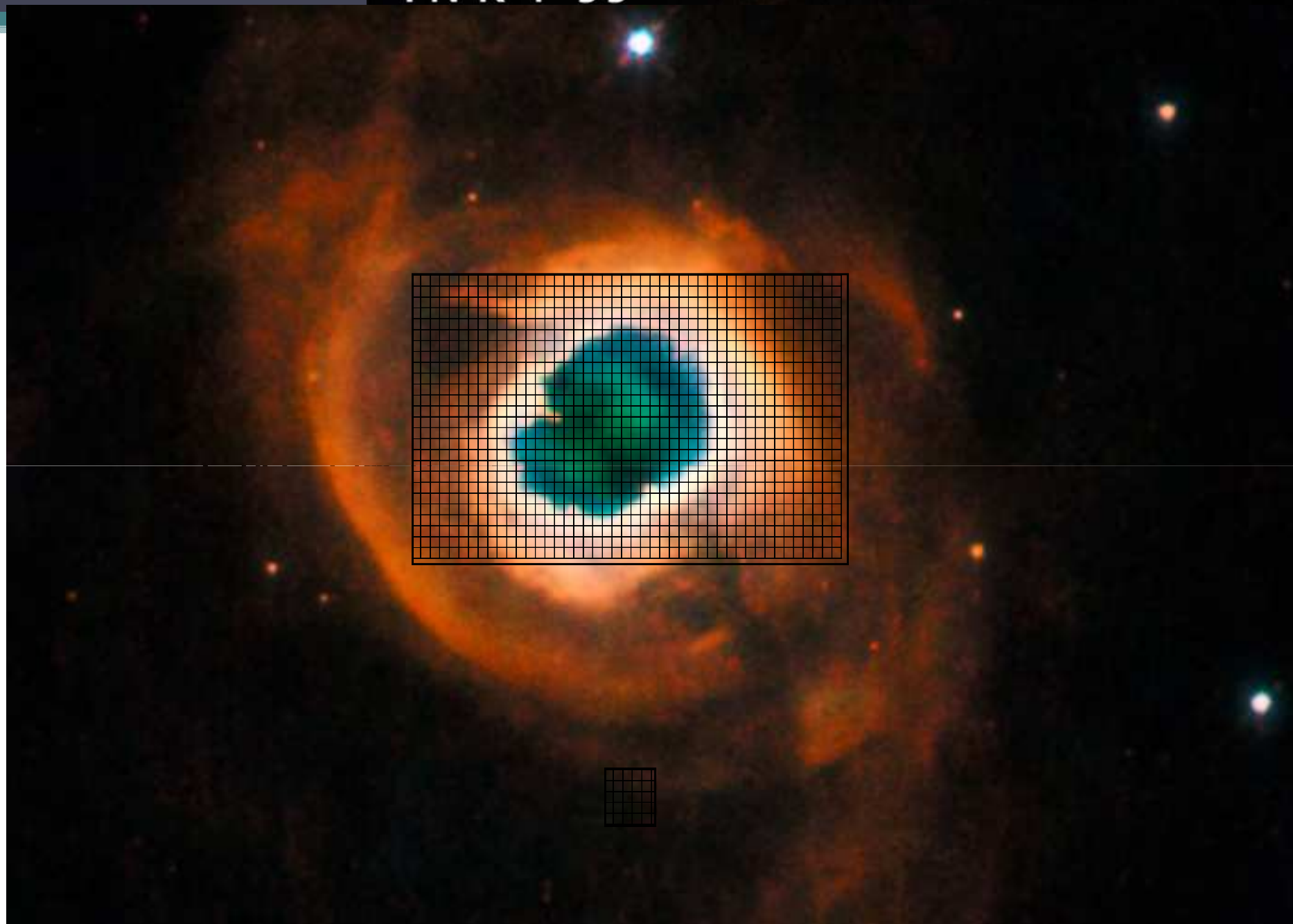
Bruno Castilho

SIFS

- IFU (1300 – 26 x 50)
7.8 x 15 @ 0.3"
3.9 x 7.5 @ 0.15"
- VPH gratings
1,000-40,000
- CCD E2V 4096x4096
- $\lambda\lambda$ 320-850nm



PN K 4-55



0.3 parsec

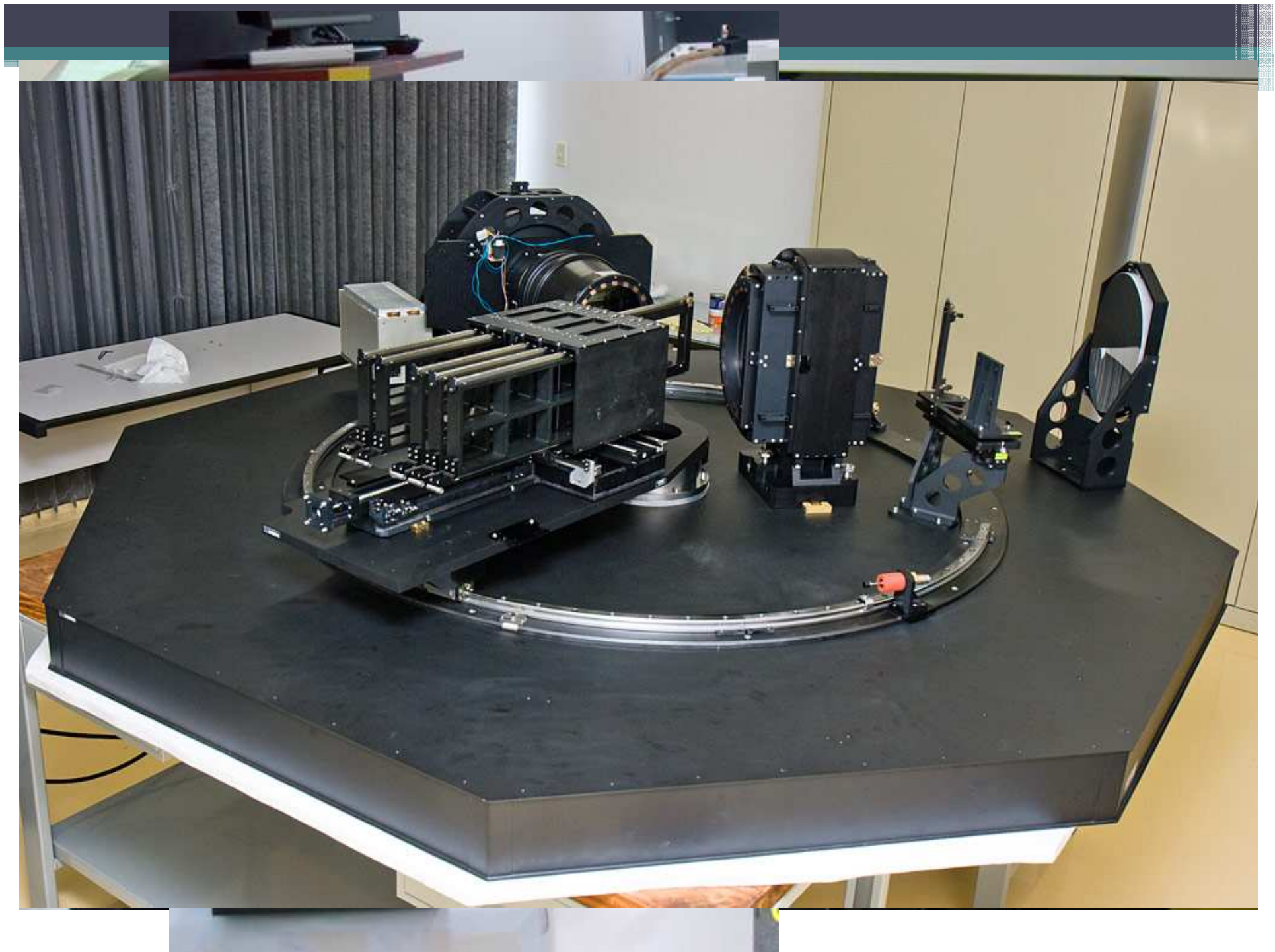
45''



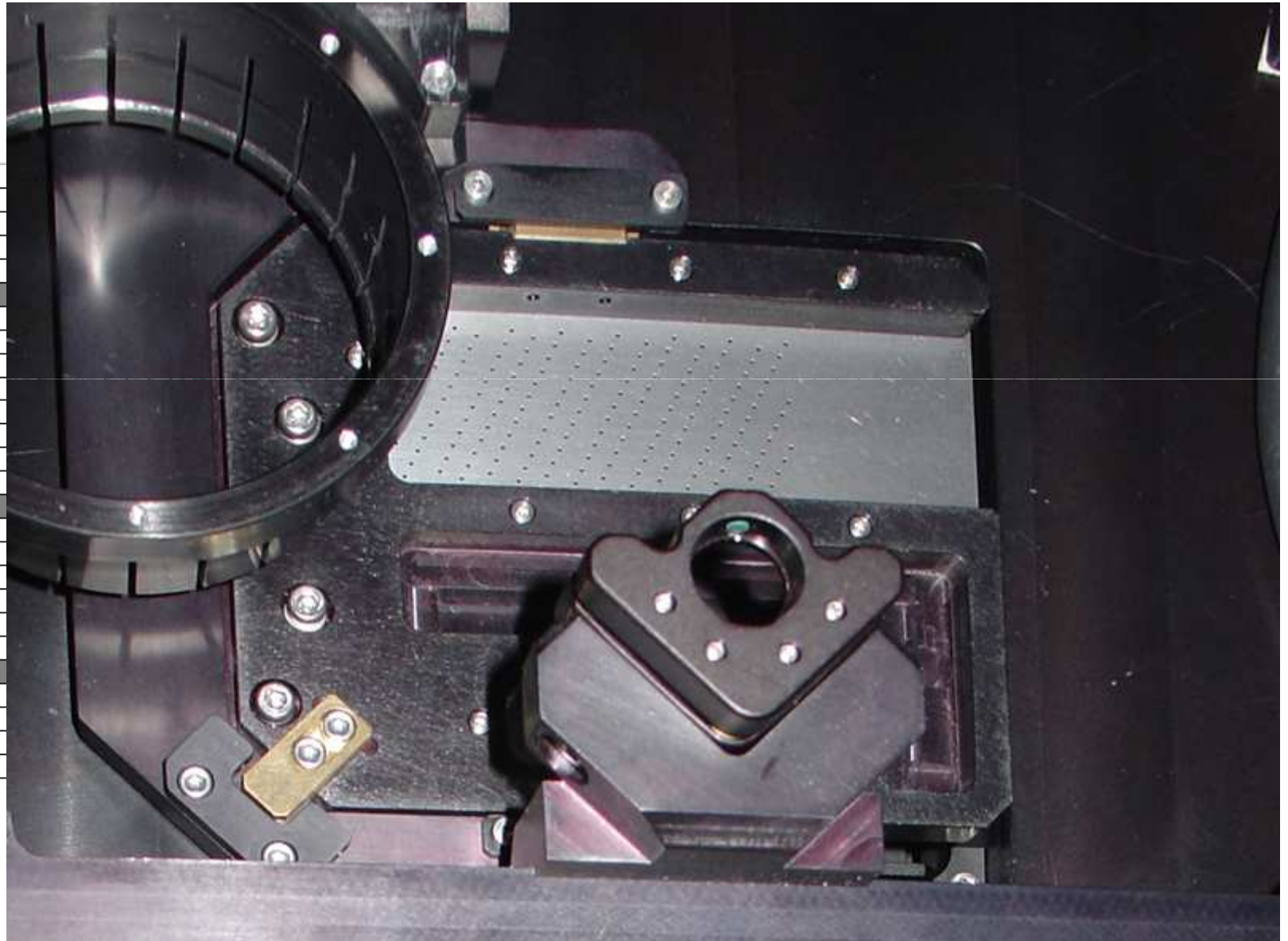
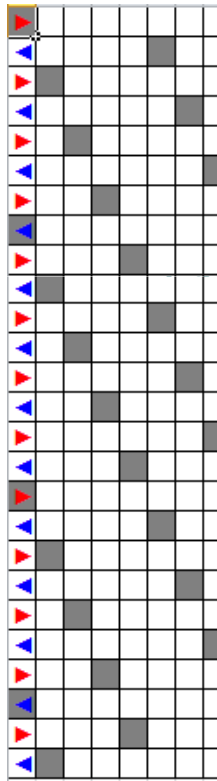
SIFS units

- Fore-optics
 - Magnification 7.8x15 @ 0.3", 3.9x7.5 @ 0.15"
 - Filters 4 positions + focus camera
 - Mask clear, nod & shuffle, 8 flat positions
- Bench Spectrograph
 - Gratings, grating angle 5 gratings, 1 clear (max 53deg)
 - Camera focus, camera angle remotely controlled focus
 - Milk Flat possible upgrade

More on poster 26 Cesar Oliveira



SIFS

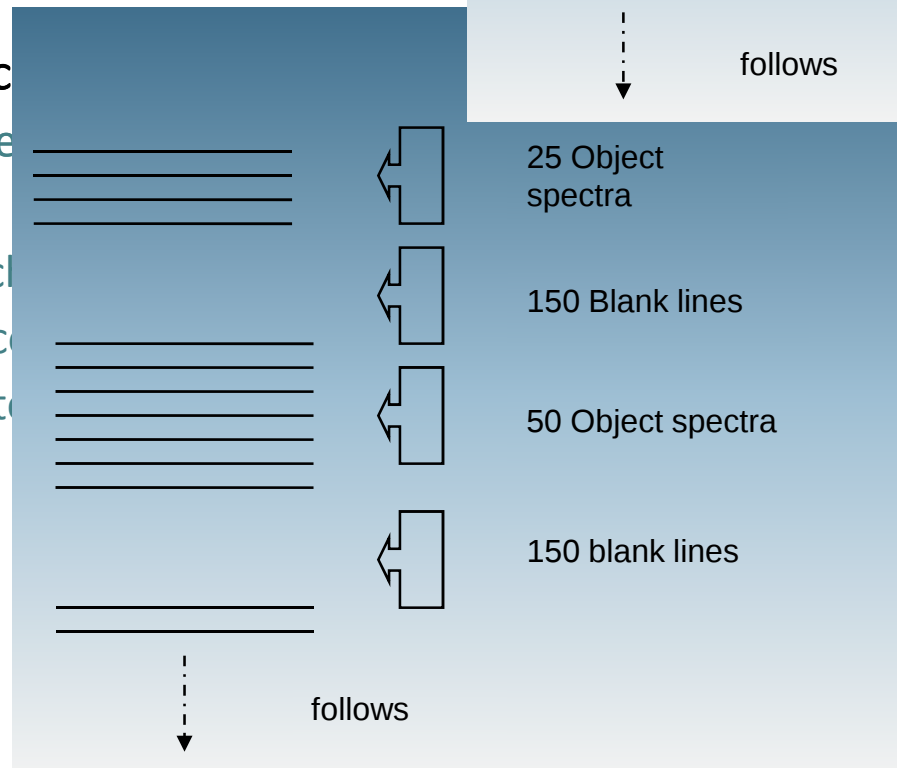
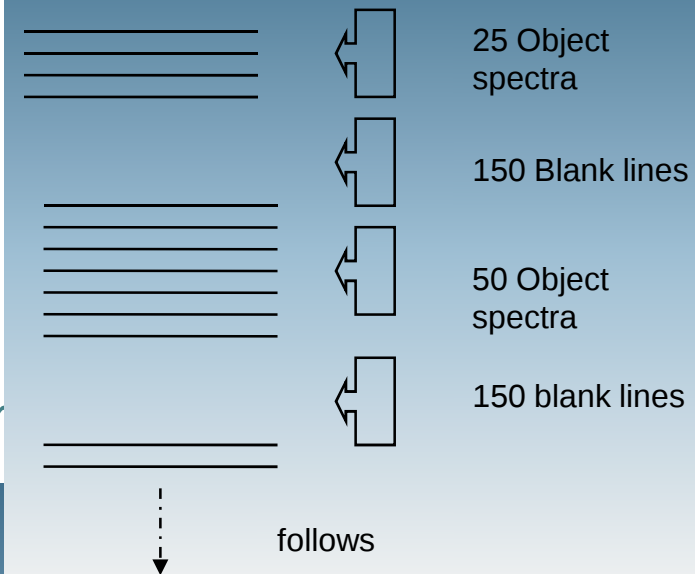


SIFS nod & shuffle

- Electronic nod & shuffle
 - Upgrade to be implemented in the future

- Foreoptics mask c

- Object is recorded
- That give us:
- Then we move cl
- Point now telesco
- Return charges to
- Repeat...



Gratings

- VPHs
 - (1) = 700
 - (2) = 1500
 - (3) = 2200
 - (4) = 2600
 - (5) = 3000
 - (6) = clear





SIFS Phase I & II tools

- Exposure Time Calculator
 - Not implemented – need commissioning data
- Image simulator
 - Simulate necessary files for test data reduction
 - Can be used to simulate results at observation planning
- Gratings software
 - Central Wavelength + Resolution
 - ➔ Grating number + grating angle

Image sim

- Python 2 +

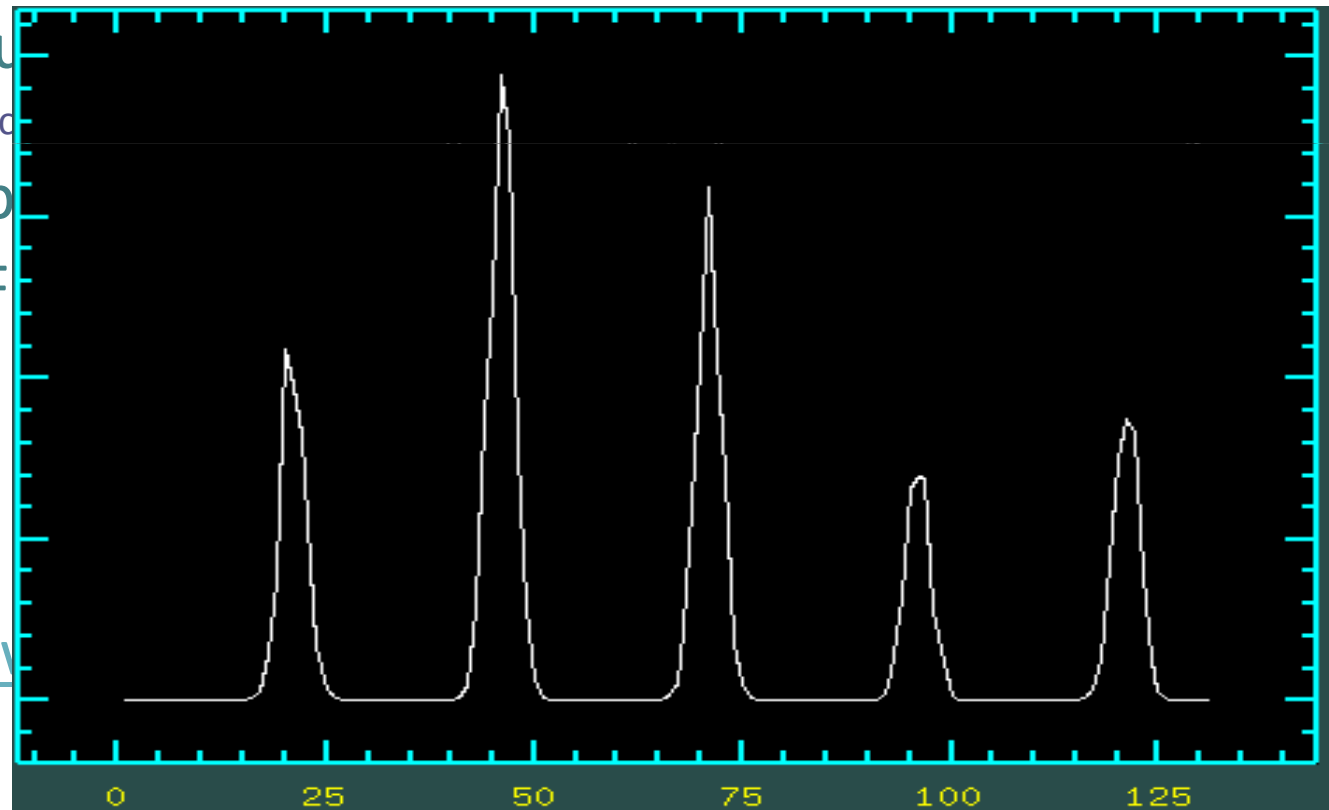
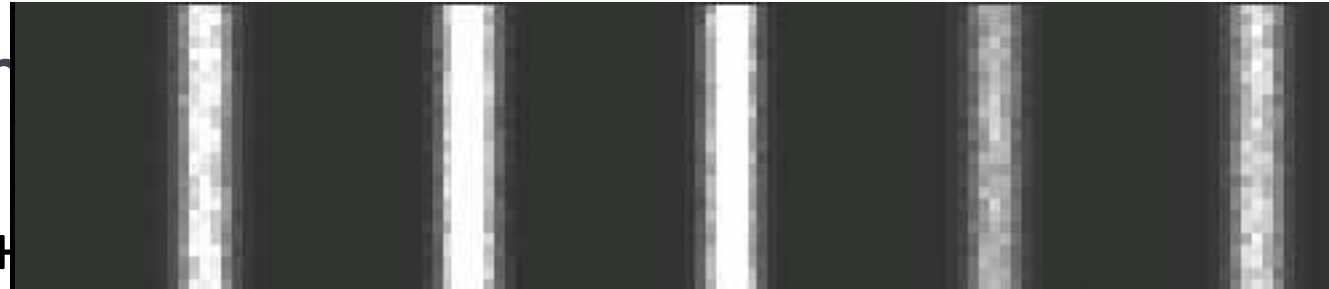
- Open sou

- (Linux, Unix)

- Fiber sep

- Object, F

- <http://www>



SIFS Phase

Please enter the desired

'Chosen Central Wavelength

'Chosen Resolution

'Possible configuration for

'-----'

'| Grating line density

'| Grating angle

'| Resolution (dif. limit)

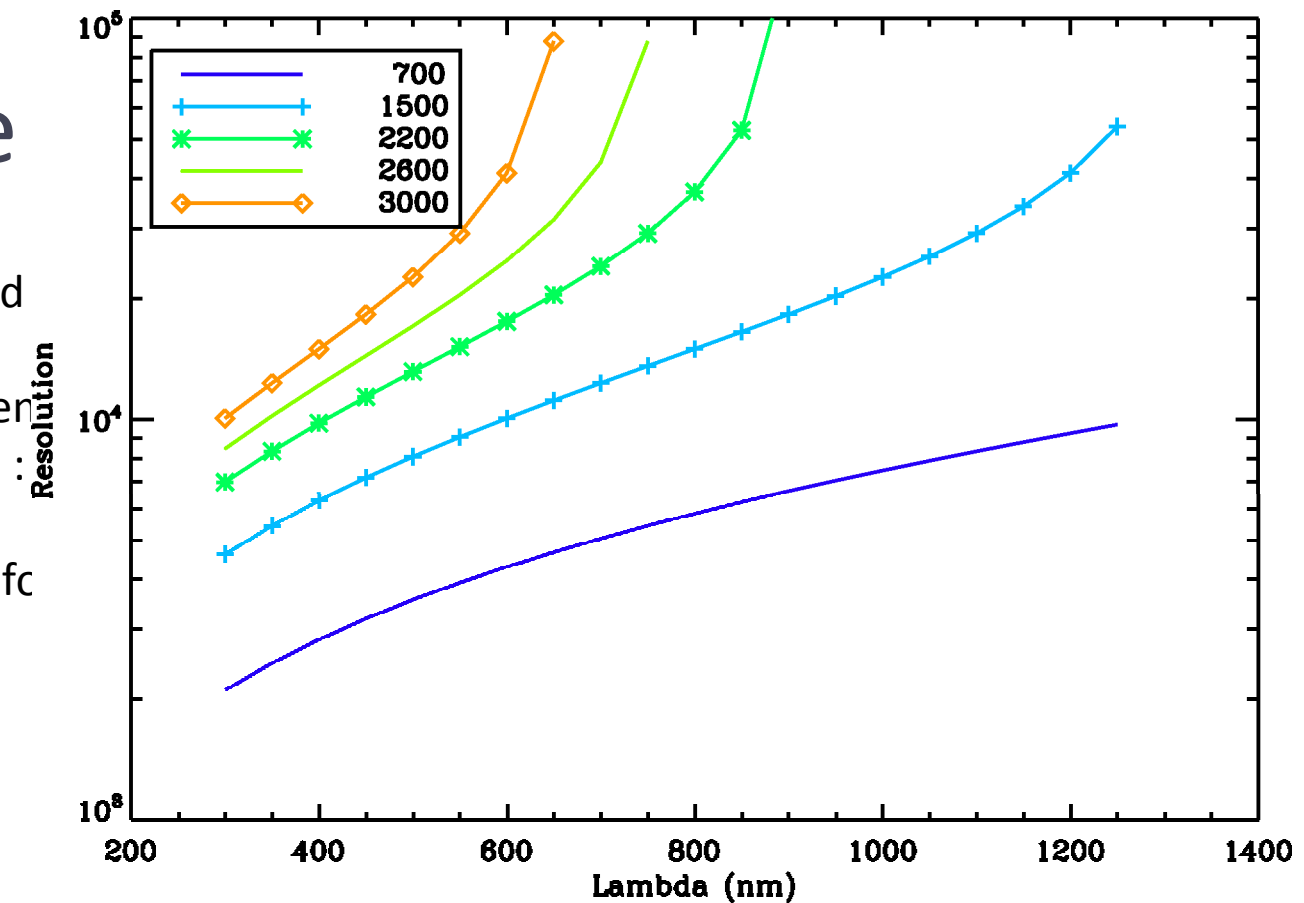
'| Resolution (1 fiber)

'| Resolution (1 pixel)

'| Dispersion (Å/mm)

'| Dispersion (Å/pix)

'| Spectral coverage (Å)





SIFS GUI

- Labview
 - Following SOAR guidelines
 - TCS data for visualization
 - Instrument setup
 - Calibration
 - Centering / Focus
 - Observing



SOAR Integral Field Unit Spectrograph

Energy Status
Fore-optics ☒
Bench ☒

TCS Information

R.A.	Dec
17:35:00.261	-23:45:44.702
Airmass	Focus
1.25	-1342.00

Mount AZ	Mount El
67.0000	56.0000
Dome AZ	Dome Shutter
90.0000	0.0035

Date	Sidereal Time
2011-04-23	22:51:06.4

LT Time	Hour Angle
15:45:41.6	-00:05:04.0

Rotator Angle	Position Angle
0.001000	360.000000

Outside Temp.	Wind Direction
15.5678	327.1234

Pressure	Wind Speed
756.1234	10.1234

Seeing	Humidity
1.1234	30.1234

Instrument Setup

Park Instrument ☒

SIFS READY ☒

Fore-Optics

SIFS fold mirror
☒ IN ☒

Magnification 1-3
1 ☒

Filter 0-4
0 ☒

Mask Position 0-9
0 ☒

Bench

Grating 1-6
1 ☒

Grating Angle
32 ☒

Camera Angle
64 ☒

Grating x Camera Sync
☒ ON

SETUP CATALOG 1 ☒ OK ☒

CAMERA Motor

DETECTOR FOCUS



7.1234

Go

Standard Position 5.0



Observation

Base Name

Type

Title

Path

Comment

Observer

Exp. Time

Repeat

Go Exposure

Stop Sequence

Abort

Exposure Progress

Read Out Progress

Write Progress

Telescope Focus

Focus Camera



Filter Position

0

Telescope Focus



7.1234

Go

XY Graph

Plot 0 ☒

